

Construction of earthquake location uncertainty maps for the Netherlands

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Samenvatting

Een groot deel van Nederland is bedekt met een dikke laag geconsolideerde sedimenten. In het onderliggend vastere deel van de bovenkorst vinden verschillende antropogene processen plaats. Olie en gas worden geproduceerd, zout wordt gewonnen met oplossingsmijnbouw, vloeistoffen en gassen worden opgeslagen, en geothermische energie wordt gewonnen. Al deze activiteiten kunnen het spanningsveld rondom bestaande breuken veranderen en daarmee mogelijk leiden tot seismiciteit.

Gedurende de afgelopen decennia is een uitgebreid sensoren netwerk opgebouwd om seismiciteit te monitoren. Dit netwerk heeft een hoge sensordichtheid op plekken waar veel gasvelden zijn geproduceerd. Op andere plekken zijn er weinig sensoren. In dit rapport is de capaciteit van het netwerk onderzocht met de configuratie van september 2021. Dit is gedaan om te achterhalen in hoeverre het netwerk voldoet om huidige en toekomstige ondergrondse activiteiten seismisch te monitoren. Ook worden de resultaten getoond voor de september 2022 staat van het netwerk. Voor dit onderzoek zijn nieuwe gereedschappen ontwikkeld om vast te kunnen stellen wat de locatie limieten zijn en de verwachte locatie onzekerheden.

De belangrijkste parameter om de capaciteit van een seismisch monitoringsnetwerk te kwantificeren is magnitude van completie (MoC). Deze parameter beschrijft de spatiële variatie van de minimale magnitude waarvan bijna elke beving gelokaliseerd kan worden. De locatie van een beving kan vastgesteld worden als de beving op tenminste 3 sensoren is gedetecteerd. Wanneer een beving gelokaliseerd kan worden, is het van belang om te weten met welke nauwkeurigheid dit gedaan kan worden. Kennis van de diepte-onzekerheid is belangrijk, bijvoorbeeld, om tektonische bevingen te kunnen onderscheiden van geïnduceerde bevingen. Kennis van de epicentrale onzekerheid (onzekerheid in het horizontale vlak) is belangrijk om een beving te kunnen relateren aan een bepaalde breuk en mijnbouwoperatie. Met kennis van de verwachte locatie-onzekerheid van het huidige netwerk kan bepaald worden of extra sensoren nodig zijn.

In dit rapport zijn kaarten ontwikkeld die de verwachte locatie-onzekerheid tonen voor bevingen in de bovenkorst in Nederland. Een belangrijk onderliggend model is de onzekerheidsverdeling van seismische reistijden. Dit model is afgeleid van vele metingen van bevingen in Nederland. Met een Bayesiaanse analyse zijn onzekerheden in reistijden gepropageerd naar onzekerheden in locatie. Er is vanuit gegaan dat locatie gedaan wordt met verschillen in aankomsttijden van de primaire (P) golf bij verschillende sensoren, en met reistijdverschillen tussen de P en schuif (S) golf aankomsten. Deze data attributen bevatten complementaire informatie. Een gezamenlijke inversie van beiden geeft de kleinste locatie-onzekerheden. Met deze gezamenlijke inversie zijn de kaarten berekend voor lokale magnitudes van 0,5 tot 4,0 met stappen van 0,5.

De kaarten zijn berekend over een grid van 1x1 km. Voor elke potentiële beving is de locatie kansendichtheidsfunctie bepaald, als functie van diepte en in het epicentrale vlak. Vervolgens is aangenomen dat deze kansendistributies goed kunnen worden benaderd als zijnde (multivariate) normaalverdelingen, zodat ze beschreven kunnen worden met standaard deviaties. Deze standaard deviaties worden getoond op de kaarten. Daarmee kan dan weer de grootte van het gebied, of dieptebereik, berekend worden dat bijvoorbeeld 68% of 95% van de kansen bevat. De kaarten zijn doorgerekend voor een beving diepte van 3 km. Daarnaast is een gevoeligheidsanalyse gedaan om te evalueren hoe de locatie onzekerheid verandert als functie van verschillende parameters. Bij een grotere diepte dan 3 km neemt de epicentrale onzekerheid toe en zal ook de diepte-onzekerheid iets toenemen.

Signaal-ruisverhoudingen spelen een belangrijke rol in de bepaling welke sensoren een mogelijke beving zouden detecteren. Voor het bepalen van het verwachte signaalniveau is een P-golf grondbewegings-vergelijking afgeleid. Deze vergelijking beschrijft de P-golf amplitudes als functie van brondiepte, epicentrale afstand en magnitude. Het ruisniveau per ontvanger is verkregen met de lokale meting van de seismische achtergrondruis in dezelfde

frequentieband die gebruikt wordt voor detectie (5 tot 40 Hz). Als ruisniveau is de 90ste percentiel van de ontvanger-specifieke ruismeting gebruikt. Dit is het ruisniveau dat 10% van de tijd wordt overschreden.

De MoC en locatie-onzekerheids kaarten tonen een grote variatie van waarden over Nederland. De belangrijkste onderliggende oorzaak is de sterke variatie in sensordichtheid. Een andere belangrijke reden is de sensorplaatsing, te weten nabij het aardoppervlakte of in 200 m diepe boorputten. Gemiddeld genomen zijn de ruisniveaus op 200 m diepte een factor 26 keer lager dan aan het aardoppervlak, in de detectie frequentieband. Anderzijds zijn ook de signaalamplitudes een factor 4 lager op 200 m diepte. Beide effecten samen leveren een winst in signaalruisverhouding van een factor 6.5 met plaatsing op diepte. Daarnaast is de lokale setting van belang. Een regionaal netwerk van sensoren in een druk stedelijk deel van Zuid-Holland presteert minder goed dan een vergelijkbare ontvangerverdeling in een rustig ruraal deel van Twente. De sensoren die het beste presteren wat betreft detectie van seismiciteit zijn sensoren geplaatst op hard gesteente in Zuid-Limburg.

Het rapport bevat magnitude van complete (MoC) kaarten voor 2021 en 2022. De onderliggende waardes zijn te raadplegen via *KNMI* (2023a). De komende jaren zullen er updates blijven verschijnen op het KNMI Data Platform. De kaarten zijn gemaakt met gemeten stations-specifieke ruiscondities en zijn daarmee een verbetering ten opzichte van eerdere kaarten die gemaakt zijn met een aanname van gemiddelde ruiscondities over het seismische netwerk. Uit de nieuwe 2021 en 2022 kaarten volgt dat de MoC lager is of gelijk is aan 2,0 voor heel Nederland. In gebieden met hoge sensordichtheden (Groningen, delen van Friesland, Drenthe, Twente, Zuid-Limburg en de regio rondom Alkmaar) is de MoC lager dan 0,5. Binnen het relatief dichte netwerk in Zuid-Holland wordt 0,5 niet bereikt vanwege hoge ruiscondities. In het midden van het land zijn er niet veel seismische sensoren aangesloten op het KNMI netwerk en bereikt de MoC waardes tussen de 1,5 en 2,0. Voor de berekening van de MoC is de 90ste percentiel van de ruiscondities gebruikt. Dientengevolge zou in ongeveer 10% van de gevallen de MoC (net) niet gehaald worden.

Voor hogere magnitudes is de locatie onzekerheid lager, omdat dan een groter netwerk van sensoren gebruikt kan worden om de locatie van de beving vast te stellen. Binnen de dichtbemonsterde regionale netwerken in Nederland is de epicentrale locatie-onzekerheid veelal kleiner dan een paar honderd meter. Deze onzekerheid groeit tot een paar kilometer, aan de randen van gebieden waar locatie nog net mogelijk is.

De onzekerheid in dieptebepaling gedraagt zich anders dan de epicentrale onzekerheid. Over het algemeen kan de diepte met minder zekerheid bepaald worden dan het epicentrum. Voor het vaststellen van de diepte is het van belang om sensoren te hebben nabij het epicentrum, zodat het hyperbolische deel van de reistijdencurve bemonsterd wordt. Voor het vaststellen van het epicentrum is het belangrijker om een goede azimuthale distributie van ontvangers te hebben. Daarnaast is voor het vaststellen van diepte van belang om informatie te hebben over de lokale snelheidsopbouw van de ondergrond. Voor het vaststellen van het epicentrum kan veelal een gemiddeld of een aardbevings-specifiek reistijdenmodel gebruikt worden.

De berekende locatie onzekerheid wordt minder correct weergegeven voor bron-ontvanger configuraties waarbij het azimuthale gat groter is dan 250° . Het azimuthale gat is de grootste hoek tussen twee ontvangers, zoals gezien vanuit het epicentrum. Wanneer het azimuthale gat groter is dan 250° , is de locatie kansverdeling niet meer goed te beschrijven als een (multivariate) normaalverdeling. De grootste onzekerheden treden op wanneer ontvangers bijna op een lijn staan, hetgeen overeenkomt met een azimuthaal gat van bijna 360° . In dat geval kan de locatie ambigu zijn en is het belangrijk om de polarisatie van P-golven mee te nemen om deze ambiguïteit te verhelpen. In de appendix zijn kaarten opgenomen met daarop de grootte van het azimuthale gat.

In de appendix staan de verkregen locatie-onzekerheidskaarten voor het Nederlands seismisch netwerk van 2021 en 2022. Ook is de onderliggende database publiek beschikbaar met daarin de waardes voor elke gridpunt (*KNMI*, 2023b). Wanneer precieze waarden op een

bepaalde plek van belang zijn, kan deze database geraadpleegd worden. De komende jaren zullen er updates blijven verschijnen op het KNMI Data Platform.

De ontwikkelde hulpmiddelen zijn ook in te zetten om de capaciteiten van scenario netwerken door te rekenen. Voor een mogelijk toekomstige sensorendistributie kan uitgerekend worden wat de verwachte impact is op MoC en locatie-onzekerheid.

Summary

A large part of the Netherlands is covered by a thick blanket of unconsolidated sediments. In the underlying more consolidated part of the upper crust, various anthropogenic activities are taking place. Oil and gas are produced, salt is extracted through solution mining, liquids and gasses are stored and geothermal heat is exploited. All these activities can alter the stress field around existing faults and therewith potentially induce seismicity.

Over the years, an extensive seismic network has been installed to monitor seismic activity. This network has a dense distribution of sensors in regions with past subsurface activity and is more sparse in other regions. In this report, the network capabilities are assessed for the September 2021 network configuration, to find out whether this network suffices to monitor current and future subsurface operations. Also, results are shown for the September 2022 network configuration. New tools are developed to assess location thresholds and expected location uncertainties.

The main seismic network-assessment parameter is magnitude of completeness. This parameter describes the spatial variation of the minimum magnitude for which almost every earthquake can be located. The location of an earthquake can be determined if the signal is detected on at least three sensors. As soon as an event can be located, it is important to know with which precision. Knowing the depth uncertainty is important, e.g., to discriminate induced events from tectonic events. Knowing the epicentral uncertainty is important, e.g., to assign the event to one of multiple nearby mining operations. By knowing the expected location uncertainties pertaining to the current network, decisions can be made on the need of additional sensors.

In this report, maps are developed that show the expected location uncertainty over the Netherlands for upper-crustal seismicity. An important prerequisite is a travel-time uncertainty model, which is derived from a large collection of records from induced events. Uncertainty in travel time is mapped to uncertainty in location using a Bayesian framework. Location is considered using differential P-wave arrival times over the network (P-delays), and also by using the travel-time difference between first P and S waves at each receiver (P-S delays). It is shown that both data attributes yield complimentary information on the location problem and that a joint inversion with both P- and P-S delays yields smallest location uncertainty. For this specific combination of data attributes, the location-uncertainty maps are computed for local magnitudes of 0.5 until 4.0, with steps of 0.5.

The maps have been computed over a grid of 1x1 km. For each scenario event, the location probability density function has been determined, both as function of depth as as function of the horizontal coordinates. The assumption is made that these probability density functions can be approximated well as being (multivariate) normal distributions, such that they can be described with standard deviations. These standard deviations are shown on the maps. From the standard deviations the size of, for example, the 68% and 95% confidence zones can be computed. The maps have been computed for a source depth of 3 km. Besides, a sensitivity analysis has been performed to evaluate how the location uncertainty varies as function of different parameters. When the source is placed at depths larger than 3 km, the epicentral uncertainty increases, and also the depth uncertainty somewhat increases.

Signal-to-noise levels play an important role in the determination at which sensors a potential earthquake is likely detected. The signal is modeled through deriving P-wave ground-motion prediction equations that describe the expected P-wave amplitudes as function of magnitude and distance from the source. The noise level is based on receiver-specific recordings of background seismic noise in the frequency band that is used for detecting induced events (5 to 40 Hz). As the location-specific noise level, the 90th percentile of the recorded noise distributions is used, meaning a noise level that is exceeded 10% of the time.

The magnitude-of-completeness and location-uncertainty maps show large variations over the Netherlands. The main underlying factor is the large variation in sensor density. Another important factor is the position of the sensor, i.e., on the Earth's surface or 200 m below the

surface in a borehole. On average, in the detection frequency band, noise levels are a factor of 26 lower at 200 m depth than at the Earth's surface. On the other hand, signal levels are a factor of 4 lower at 200 m depth. Combining the two effects, this results in a signal-to-noise gain of 6.5 at the 200 m depth sensors. Moreover, the local setting is important. A regional network in South-Holland performs worse than a similar network in Twente because of much higher seismic noise conditions near Rotterdam and the Hague than in the north of Twente. The sensors with the best detection capabilities are located on hard-rock conditions in the southernmost part of the Netherlands.

The report contains magnitude-of-completeness (MoC) maps for 2021 and 2022. The underlying values can be obtained at *KNMI* (2023a). The coming years, updates will continue to appear at the KNMI Data Platform. Previous maps were made with assumed average noise conditions over the seismic network. The new maps have been made with recorded station-specific noise conditions, which yields a more accurate representation of the MoC that can regionally be achieved. On the new 2021 and 2022 maps, MoC is smaller or equal to 2.0 in the Netherlands. In areas with dense sensor distributions (Groningen, parts of Friesland, Drenthe, Twente, south of Limburg and region around Alkmaar) the MoC is lower than 0.5. The network in South-Holland does not reach MoC=0.5 due to high local noise conditions. In the middle of the country, there are not many seismic sensors and the MoC has values between 1.5 and 2.0. For the computation of the MoC, the 90th percentile of the noise distribution has been used. As a consequence, in about 10% of the cases the stated MoC can (just) not be reached.

Location uncertainty reduces with event magnitude as a better spatial distribution of receivers detects the event for higher magnitudes. Within the dense regional networks, the epicentral uncertainty is within a few hundred meters. The uncertainty grows to a few kilometers at the edges of the regions where location is just possible.

Depth uncertainty behaves quite differently from epicentral uncertainty. In general, the depth is less well constrained than the epicenter. For constraining the depth, it is important to have receivers nearby the epicenter, so that the hyperbolic part of the travel-time curve is sampled well. For constraining the epicenter it is more important to have a good azimuthal distribution of receivers. For accurate depth estimation, local velocity information needs to be available. For estimating epicenter, one can mostly rely on background or event-derived travel-time models.

The computed location uncertainty is less well represented for source-receiver configurations for which the azimuthal gap is larger than 250° . The azimuthal gap is the largest angle between two receivers as seen from the epicenter. When the azimuthal gap is larger than 250° , the location probability density function cannot be approximated well as being a multivariate normal distribution. The largest uncertainties occur when the receivers are nearly on one line, which corresponds to an azimuthal gap of nearly 360° . In this case, location can be ambiguous and the polarisation of P-waves is needed as an additional constraint to remove this ambiguity.

In the appendix, the location-uncertainty maps can be found for the Netherlands seismic network state of 2021 and 2022. Also the underlying database, with the values at each grid point, is publicly available (*KNMI*, 2023b). This database is to be consulted if the precise values at a certain location are of interest. The coming years, updates will continue to appear at the KNMI Data Platform.

The developed tools also provide a means to study design options for future extensions of the network. A scenario layout can be checked for its capability in terms of the minimum magnitude of locatable earthquakes and their associated uncertainty in location.

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1 Introduction

In the Netherlands there is a growing number of mining operations. Comprehensive knowledge has been built up of the upper crust. E.g., detailed velocity models (*Van Dalfsen et al.*, 2006) and fault maps (*Ten Veen et al.*, 2022) are openly available. This allows precise planning of subsurface operations, ranging from hydrocarbon production to salt extraction. These subsurface operations perturb the existing stress field and therewith have a potential to induce seismic activity on pre-existing faults (e.g., *Buijze et al.*, 2023). This seismicity is to be monitored, especially if it reaches levels that can be felt or inflict damage. Monitoring is the first step to find, and eventually mitigate, the underlying cause. This cause could be natural—there is tectonic stress build up and release in the Netherlands—or could have an anthropogenic origin. It is important that the location accuracy is sufficient to pinpoint the natural process or mining activity responsible for the seismic event. The Royal Netherlands Meteorological institute (further: KNMI) was asked by the The Dutch State Supervision of Mines (further: SodM) to assess the location accuracy of the current seismic network.

Some gasfields are amenable to seismicity (e.g. Roswinkel and Groningen) while others have not shown any seismic event over their lifetime (e.g. gasfields in Twente). The occurrence or absence of seismicity over many different gasfields has been used to identify reservoir parameters that correlate with seismicity (*Van Eijs et al.*, 2006; *van Thienen-Visser et al.*, 2012). Despite the knowledge accrued, however, forecasting the occurrence of seismicity remains a challenge. For that reason, mining operations are to be seismically monitored to some degree, even if the chance of seismicity is thought to be small.

The monitoring capabilities largely depend on the seismic network and noise conditions. The Netherlands has an extensive network of borehole geophones, surface accelerometers and broadband seismometers (*Dost et al.*, 2017). Prior to 1990, most sensors were located in the south of the Netherlands, where tectonic activity occurs related to the Lower Rhine Graben (e.g., *Hinzen et al.*, 2021). During the end of the eighties, the first induced seismic events were detected in the north of the Netherlands, which were attributed to gas extraction (*Haak*, 1993). Since then, there has been a continuous expansion of the network to monitor induced seismicity.

Most sensors that are deployed in the Netherlands Seismic and Acoustic Network (*KNMI*, 1993) have instrument noise that is lower than the seismic background noise, at least in the frequency band that is relevant for detecting induced seismicity. As such, a bottleneck in detection of events is the seismic noise, which varies largely from site to site. For example, seismic noise levels are much higher near a highway than in the middle of a forest. To properly assess the network capabilities, this location-specific background noise needs to be taken into account.

For the Netherlands, a relatively new potential source of induced seismicity is related to geothermal energy production (*Buijze et al.*, 2019). Up to date, a magnitude 0.0 event has been detected near an operation in South-Holland (*Naranjo et al.*, 2022) and a magnitude

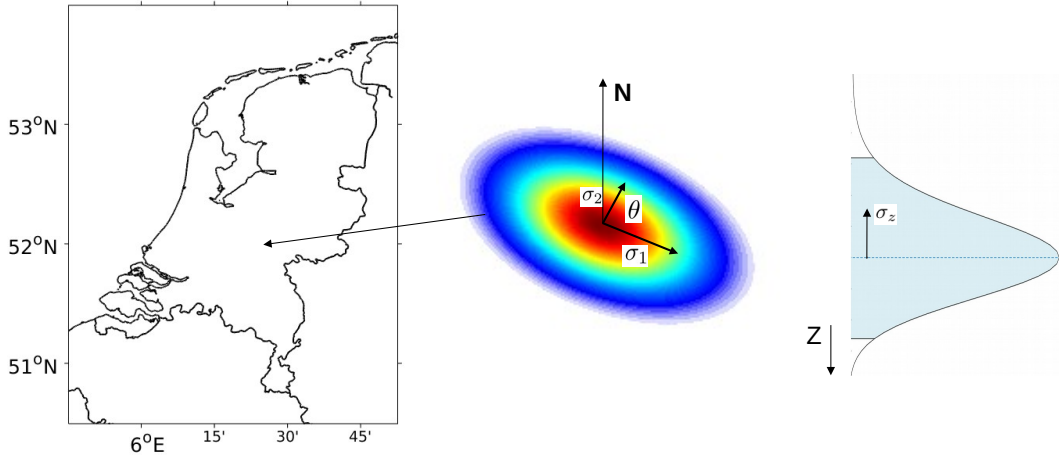


Figure 1.1: In this report, for each point in the Netherlands (left-hand figure) the epicentral probability-density function (PDF, middle figure) and depth PDF is derived (right-hand figure). The epicentral PDF is described with three parameters, σ_1 , σ_2 and θ . The depth PDF is parameterized with σ_z . The coloured areas of the epicentral and depth PDF denote the 95% confidence areas.

1.7 event near an operation in Limburg (*Vörös and Baisch, 2022*). Exploration and exploitation also takes place in areas of the Netherlands where the seismic network is sparse (*Mijnlieff, 2020*), because of an absence of gas production or tectonic activity. SodM is involved in the permitting process as advisor to the Ministry of Economic Affairs and Climate Policy. One of the questions that needs to be answered in light of safe operations is whether the nationwide existing network suffices in terms of monitoring, or that new stations are needed to adequately monitor a geothermal operation. SodM has asked the seismic network operator KNMI to assess the current network capabilities. To monitor seismicity and to connect detected seismicity to an underlying cause, two network capabilities are relevant:

Epicentral uncertainty. This is a description of location uncertainty in the horizontal plane (map view). With this uncertainty it can be assessed whether events from nearby operations can be well separated. Or, it can be assessed whether the uncertainty is low enough to pinpoint the specific fault on which the rupture must have taken place.

Depth uncertainty. This parameter describes the location uncertainty in the vertical direction. Knowing the likely depth uncertainty is important to find out whether the mining operation (in the upper crust) caused the event, or whether a tectonic release of stresses in the deeper crust is more likely. Another option is that the mining activity on a shallow depth triggered an event on a somewhat larger depth. Moreover, in case multiple mining operations are stacked, it can be assessed whether the depth-location accuracy would be sufficient to distinguish events from these different operations.

The location accuracy that can be achieved is a function of event magnitude. For higher magnitudes, more stations detect the event above the noise floor, which enables a better constrained source solution. Fig. 1.1 illustrates the type of maps that are developed within this report. For each point in the Netherlands and for a range of magnitudes ($M=0.5, 1.0, 1.5, 2.0, 2.5, 3.0, 3.5$ and 4.0), the location uncertainty is modelled and mapped. This amounts to generating maps for the Netherlands showing the epicentral uncertainty (which is parameterized with σ_1 , σ_2 and θ) and maps showing depth uncertainty (which is parameterized

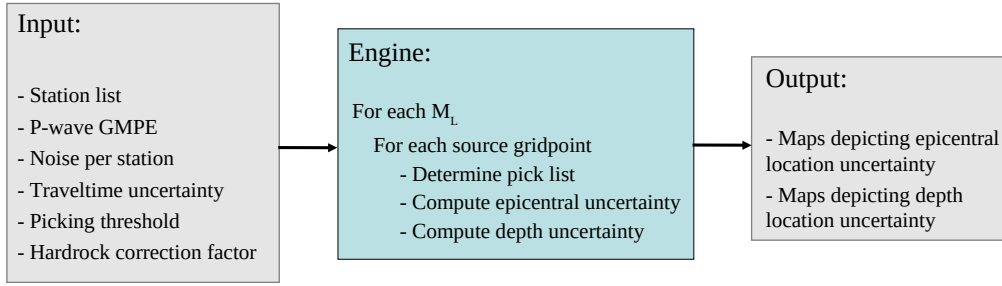


Figure 1.2: A flowchart summarizing the different components that are developed within this project. The required 'input' or prerequisites are listed on the left. The 'Engine' is an implementation of a Bayesian method to compute location probability density functions. The input and engine are developed to yield the desired 'Output': epicentral and depth uncertainty maps for the Netherlands.

with σ_z). The exact meaning of these uncertainty parameters is discussed in Chapter 3.

Fig. 1.2 shows a flow-chart for producing the location uncertainty maps. The outline of the report follows this flow-chart. In Chapter 2 (Prerequisites) all the different inputs for the analyses are derived and discussed. E.g., a P-wave ground-motion prediction equation is derived in order to model the amplitude at a station at a given distance from a scenario event. For that station, the noise conditions need to be known to determine whether it is likely that this station detects the scenario event. In Chapter 3 the methods are developed to propagate errors in timing of P- and S-waves (due to unknowns in the velocity structure, picking errors or modelling errors) to location uncertainty. The implementation of this 'method' forms the 'engine' for computation of the location-uncertainty maps. In Chapter 4 it is assessed what to be the influence on location uncertainty of different parameters, like depth of the source, azimuthal gap and number of receivers. In Chapter 5 the produced maps are shown and discussed, followed by a final Discussion and Conclusions chapter and an appendix that contains the NL-wide location-uncertainty maps.

As a by-product, in this report the magnitude of completeness (MoC) map is updated. In order to find out that an earthquake has happened, it first needs to be detected. Detection can be done already with a single sensor. However, a single detection is not sufficient to constrain both depth and epicenter of the event, or to be sure that the recorded ground motion was caused by a fault rupture in the upper crust. In order to determine a location with some confidence, detection at a minimum of three receivers is required. The MoC at a certain point is the minimum magnitude for which these three detections are likely reached. A MoC map shows how this minimum magnitude varies over an area. For the Netherlands, a first MoC map was derived in *Van Eck et al. (2004)* and later updated in *Dost et al. (2012)*, *Dost et al. (2017)* and *Kruiver et al. (2021)* to incorporate network extensions. These maps were based on average noise conditions at 200 m deep geophones and surface receivers that existed before 2004. Besides modeling the MoC spatially, one can also, on hindsight, evaluate which MoC was achieved over a certain region. Such estimates, using the KNMI induced event catalogue, can be found in *van Thienen-Visser et al. (2016)*, *Dost et al. (2012)* and *Muntendam-Bos and Grobbe (2022)*. In this report, a thorough update is made of the anticipated MoC by incorporating recent developments of the network, the newly derived P-wave ground-motion prediction equations and site-specific noise curves.

2 Prerequisites

The construction of a location-uncertainty map requires many elements. These prerequisites are assembled and developed in this chapter:

- **List of stations** that are included in the analysis. This must be a list of stations that can actually be processed within the operational chain at KNMI. Then still, different choices are possible. Are Belgian and German stations included to cover the border regions? Do we include temporary networks? Do we only include stations that yield real-time data? Etc.
- **P-wave ground-motion prediction equations** or GMPE are relations that model the expected maximum amplitude for incoming primary (P) waves, as function of distance and magnitude. Detection of induced seismicity is done based on P-waves using relatively high frequencies (5 to 40 Hz). P-wave ground-motion prediction equations (GMPEs) are derived, specifically for this frequency band and (induced) seismicity in the upper crust.
- **Station-specific noise distributions** are needed to assess whether incoming earthquake-related waves would likely stand out from the noise. For the detection frequency band and using at least half a year of data, the seismic noise distribution is obtained for each station from the list.
- **Picking threshold** is the average minimum factor needed between signal and noise for an automatic detection to be made. This factor is empirically derived using a collection of recorded seismograms.
- **Hard-rock correction factors** are factors used to adapt the P-wave GMPEs to hard-rock sites. They are needed because the P-wave GMPEs are calibrated purely with recordings on soft sites (unconsolidated sediments). A few stations in the Netherlands record on hard rock, which results in lower recorded amplitudes for the same earthquake and distance. Based on assumed average sediment and hard-rock elastic properties, it is derived how much lower these amplitudes would be.
- **Travel-time models** describe the arrival times of P- and S-waves as function of distance, and the variability therein for induced seismicity in the Netherlands. This variability, leading to uncertainty of the actual traveltimes, is derived from a large database of recordings. In the following chapters, this uncertainty in time is mapped uncertainty in location.

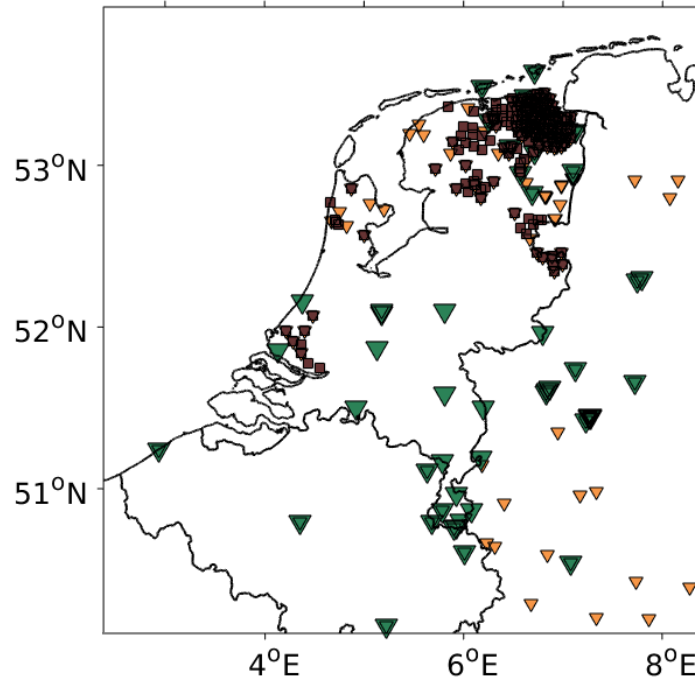


Figure 2.1: Stations in the Netherlands and direct surroundings, which data can be obtained through <https://www.orfeus-eu.org/fdsnws/> or <http://eida.bgr.de/fdsnws/>. The overview is from September 2021. Green triangles denote broadband stations, orange triangles are short-period stations and purple squares are accelerometers. Below many of the accelerometers, there is a co-located string of geophones for which the deepest node is at 200 m depth.

2.1 List of stations

In the Netherlands and surroundings, different seismic networks are active. We use FDSN web services (<https://www.fdsn.org/webservices/>) to collect all publicly available stations for the following networks

- BE: Belgian Seismic Network
- BQ: Bensberg Earthquake Network
- GE: GEOFON Seismic Network
- GR: German Regional Seismic Network
- NL: Netherlands Seismic and Acoustic Network
- NR: NARS - The Network of Autonomously Recording Seismographs
- RN: RuhrNet – Seismic Network of the Ruhr-University Bochum

The locations of these stations are depicted in Fig. 2.1. Within the NL network, not all stations are operated by KNMI. Data from stations operated by the industry (Vermilion, TAQA, Nouryon and NEDMAG) is flowing to the KNMI in real time and further distributed within the NL network. Most of the (former) NAM stations have been transferred to KNMI.

In the first iteration, all stations with a temporary nature were removed. From the overview, this only concerns the NARS network stations, which is a research network that moves every

few years to a new location. An S-wave crustal model was obtained with a former NARS deployment in the Netherlands (*Yudistira et al.*, 2017).

In the second iteration, it was decided to keep as close as possible to the station list that is used at KNMI for detection of earthquakes. All these stations provide real-time data. Moreover, for the KNMI detection list already a selection has been made of the stations with the highest signal-to-noise ratios (SNRs). For instance, only the deepest of the five sensors in a borehole setup is used during detection and is therefore included in the station list. If sensor issues are known for this specific geophone, it will be the second deepest, etc. Stations from the BQ network are not included in the detection list, because the data is not available in real-time. It takes seven days before data becomes available to KNMI.

Next, the list is restricted to stations within 100 km from the Dutch onshore border. This then includes a set of stations in Germany and Belgium that are sometimes used for location of events in the Netherlands. Stations further away may still detect the larger events in the Netherlands. However, these events are then already detected by a good azimuthal coverage of stations within the Netherlands and close to border, which obviates the need for more distant stations. Finally, sensors are removed where still multiple sensors exist at the same (borehole) location. This is the case for many of the pre-2015 borehole stations.

In the third iteration, the list is made specific for the state of the network in 2021. This is done by

1. Removing stations that have been closed by September 1, 2021. This concerns stations NL.GUR1 and NL.BING.
2. Removing stations that were installed prior to 2021 but have had less than 50% data availability over the time period September 1, 2020 until August 31, 2021. This concerns stations NL.G024, NL.STR, NL.SUH4, NL.T014, NL.WDB4, NL.WMH4, NL.WYN4 and GR.AHRW.
3. Removing stations that were (re-) installed in 2021 but have delivered less than two months of data by September 1, 2021. This concerns station NL.G314.

The final list yields 200 stations with unique locations. The sensor locations are shown in Fig. 2.2 and are listed in Appendix A. The sensor locations and specifications of the September 1, 2022 network are listed in Appendix C.

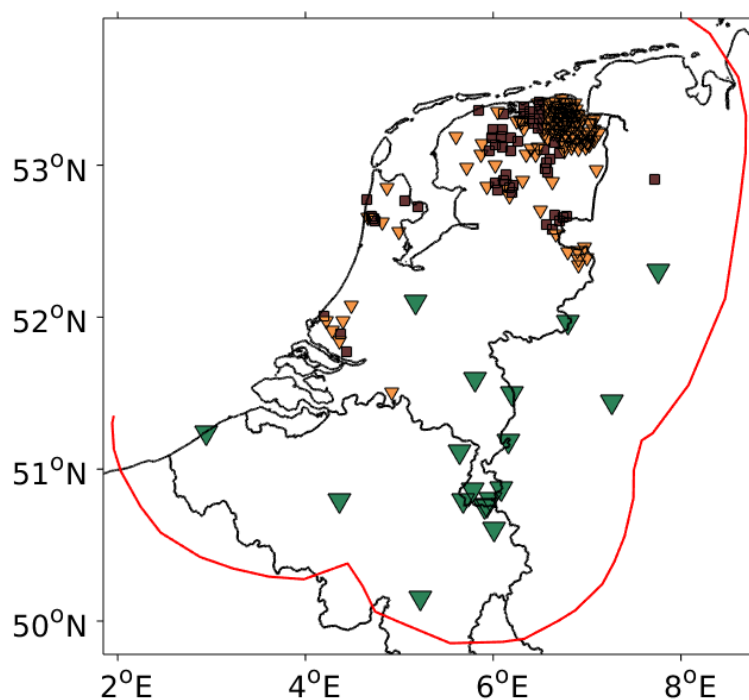


Figure 2.2: Stations that are used for earthquake detection by the KNMI, within 100 km of the Dutch onshore border (red line). The overview is from September 2021. Only stations are included that provided sufficient real-time data over the previous year (counting from September 2021). This same list of stations is used for computation of earthquake location uncertainty maps. Green triangles denote broadband stations, orange triangles are short-period stations placed at depth (> 40 m) and purple squares are accelerometers or geophones placed at the Earth's surface.

2.2 P-wave ground-motion prediction equations

To determine whether earthquake waves are likely detected, a model is needed that describes the maximum event-related amplitudes as function of (hypocentral) distance and magnitude. A model that does exactly this is called a ground-motion prediction equation (GMPE). Detection is performed with P-waves and hence P-wave amplitudes are important when studying detection capabilities. In the literature, however, the emphasis is on S-wave GMPEs also sometimes called attenuation relations. S-wave attenuation models find their use, e.g., in local magnitude determination (e.g., *Kanamori et al.*, 1993) and S-wave GMPEs are used especially for hazard studies (e.g., *Douglas et al.*, 2013). P-waves typically play a minor role in hazard studies and, as such, P-wave GMPEs usually receive little attention. Similarly, most magnitude estimation is done with S-waves, which obviates the need of a P-wave attenuation model.

In the following, we derive a P-wave GMPE for seismicity in the upper crust of the Netherlands, specifically for a detection frequency band of 5 to 40 Hz. In the Netherlands, detection is done either with sensors installed at, or close to, the Earth's surface, or with the lowermost sensor in a vertical array of geophones. These lowermost geophones are typically installed at 200 m depth. Therefore, two empirical relationships are derived, one for motions at the Earth's surface (Surface Level or SL model) and one for motions at 200 m depth (Depth Level or DL model). Calibration is done with more than 20,000 recordings from the Netherlands Seismic and Acoustic Network (*KNMI*, 1993).

As a measure, the peak-ground velocity (PGV) is used. In the remainder of the report, PGV refers to the vertical-component PGV related to incoming P waves. Since P-waves arrive at (unconsolidated-sediment) recording sites with small angles of incidence, the vertical-component PGV is a good approximation of the P-wave related maximum PGV that would be obtained from the three-dimensional ground motion.

2.2.1 Database

The ground motions database is expressed in vertical-component peak-ground velocities (PGVs) in the detection frequency band of 5 to 40 Hz. The maximum absolute amplitude is selected within a small time window around the expected P-wave arrival. The starting time of this time window is based on the average velocity structure. For distances smaller than 30 km, the time window is quite short (1.0 s) to make sure that first arriving P-waves are selected instead of, for instance, an S to P conversion. At distances larger than 30 km, the time window is increased to a duration of 2.5 s to account for heterogeneity in the velocity structure that can amass to sizeable timing differences between expected (mean model) and actual P-wave arrival times. Signal time windows are only used if their SNR is equal to, or higher than, 12 dB. This large threshold value would reject some valid earthquake ground motions. However, because of the large size of the database, this is not a big concern.

Despite a high SNR threshold, some noise-related motions do remain in the automated database. The SNR threshold works well in the presence of semi-continuous noise. The SNR does not suffice when strong transient noise signals are present, which are stronger than the earthquake P-wave arrivals in the signal time window. This especially occurs at accelerometers, many of which are located close to local roads with infrequent car traffic. A noise-source pollution of the signal time window may result in an unexpectedly strong PGV value. These can be identified by computing the misfit with the P-wave GMPE. Values that have an absolute misfit $|\ln(PGV) - \ln(model)| > 2$, are removed, where \ln is the natural logarithm. Note that this step can only be done in a second iteration, i.e., after fitting the first version of the GMPE. The rejection step results in the removal of 192 values (2.43%) from the SL database, with 7709 measurements remaining. It results in the removal of 26

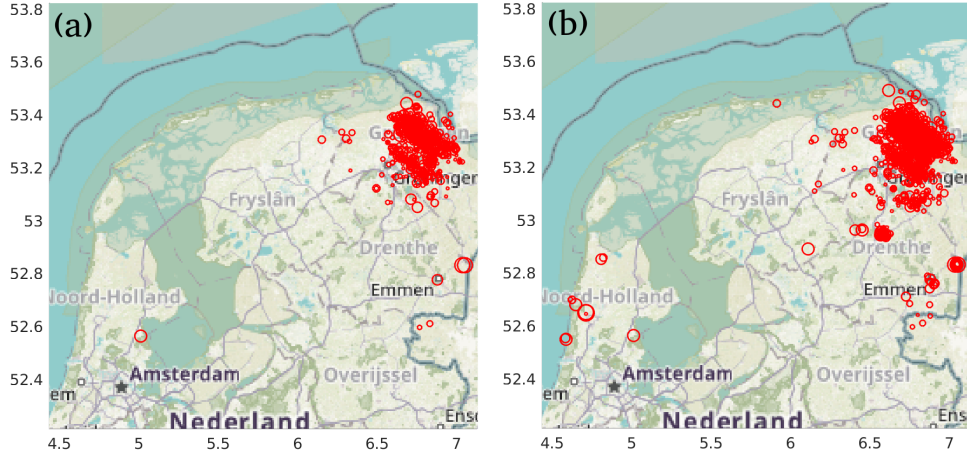


Figure 2.3: Induced events that have resulted in usable PGV records at a) the Earth's surface and b) at around 200 m depth. The size of the circles is indicative for the local magnitude, which varies from 0.4 to 3.6. The background map is obtained from www.openstreetmap.org.

values (0.16%) from the DL database, with 15753 measurements remaining, illustrating the lower noise levels at the deep geophones. Fig. 2.3 shows the source locations of the final PGV database, which contains 683 and 1382 events for the SL and DL database, respectively.

2.2.2 Ground motion model

Bommer et al. (2019) developed a GMPE specifically for the Groningen gas field (further: BMR2019). Their geometrical spreading term is partly based on numerical simulations that use the detailed 3D velocity model that is available in the Groningen region (*Romijn*, 2017). The functional form of our GMPE is largely adopted from *Bommer et al.* (2019). The main difference is that we include the event depth as an additional term, similar to *Atkinson et al.* (2016). Our model is parameterized as

$$\ln(Y_{mod}) = c_1 + c_2 M + g(R^*), \quad (2.1)$$

where Y_{mod} represents the modelled PGV expressed in mm/s, M the local magnitude and $g(R^*)$ a function that describes the geometrical spreading. c_1 and c_2 are fitting parameters. The effective point-source distance R^* is defined as

$$R^* = \sqrt{R^2 + D^2 + [e^{e_1 M + e_2}]^2}, \quad (2.2)$$

where R is the epicentral distance in km and D is the event depth in km. Eq. 2.2 contains the magnitude-dependent distance saturation term $e^{e_1 M + e_2}$ with fitting parameters e_1 and e_2 ; for small hypocentral distances ($\sqrt{R^2 + D^2}$) the exponential term is significant which results in a non-linear scaling of Y_{mod} with magnitude.

For the S-wave as developed in *Bommer et al.* (2019) the attenuation is divided in three distance ranges with different attenuation. P-waves, however, behave differently than S-waves (as in BMR2019). Moreover, the BMR2019 model has been calibrated for the Groningen region only, whereas we include all induced seismic events in the Netherlands and thus sample over more varying velocity structures. For our database we find that two distance ranges suffice. That is, the misfit is not degrading when we go from three to two distance ranges. With this simplification into two distance ranges, the attenuation terms read

$$\begin{aligned} g(R^*) &= c_4 \ln(R^*) & R^* \leq d \\ g(R^*) &= c_4 \ln(d) + c_{4a} \ln(R^*/d) & R^* > d, \end{aligned} \quad (2.3)$$

where c_4 and c_{4a} are fitting parameters. d is the effective point-source distance at which the transition occurs from strong attenuation near the source to milder attenuation at larger offset. In the following section the model is calibrated with the (P-wave) PGV database. This results in an estimate of all fitting parameters that appear in Eqs. 2.1-2.3: c_1 , c_2 , c_4 , c_{4a} , e_1 , e_2 and d . All fitted parameters are listed in Table 2.1. More details on the calibration procedure are provided in the next section.

Coefficient	SL model	DL model
c_1	-0.20	-1.60
c_2	1.96	1.96
c_4	-3.44	-3.44
c_{4a}	-1.62	-1.62
e_1	0.45	0.45
e_2	-0.80	-0.80
d [km]	8.0	8.0

Table 2.1: Parameters of the P-wave GMPEs for PGV values at the Earth’s surface (SL model) and at approximately 200 m depth (DL model). Both models are for unconsolidated-sediments sites.

A model is derived for recordings at the free surface (SL model) and a model for motions at approximately 200 m depth (DL model). Both models are the same, except for the static offset term c_1 . This term results in a 4.06 times larger PGV for the free-surface model than for the deep model. A factor of two can be explained by the free-surface effect; at the Earth’s surface the up-going and down-going waves are recorded at the same time, which results in an amplitude doubling. The other factor of approximately two is caused by near-surface amplification, which is mainly caused by a reduction of seismic impedance towards the Earth’s surface. Fig. 2.4 shows an example borehole recording where both effects can be seen on the first up-going P-wave arrival. *van Ginkel et al. (2022b)* describe the behavior of P-waves in the unconsolidated near-surface setting in Groningen.

2.2.3 Model calibration

In *Ruigrok and Dost (2020)* an S-wave GMPE was derived for induced seismicity in the Netherlands, next to the Groningen-only model (BMR2019). The Netherlands-wide S-wave GMPE model for PGV has nearly the same functional form as the P-wave model introduced in the previous section. Also, the geological setting is the same for both models. For these reasons, the parameters found for the S-wave GMPE are used as a starting point for the P-wave model calibration.

Fig. 2.5(a)-(b) shows the recorded PGV values within 2 km distance and an approximate source depth of 3 km. Maximum P-wave amplitudes are expected within 2 km epicentral distance. The maximum amplitude may occur at the epicentrum, it can also occur at a larger distance, depending on the source mechanism. The model is fitted such that it well follows the mean of the recorded values. The magnitude range of the measured PGVs is sufficiently broad to estimate magnitude-dependent saturation ($e^{e_1 M + e_2}$ in Eq. 2.2). It is clear from Fig. 2.5(a)-(b) that only curved lines fit the data points adequately. The strong saturation is probably related to the relatively high frequencies ([5 40] Hz). At epicentral distances of 10 km the magnitude-dependent scaling is approximately linear (Fig. 2.5(c)-(d)). Fig. 2.5(c) has no PGV datapoints below $1.0 \mu m/s$. Due to high noise levels on the surface instrument, lower ground motions have been rejected through the SNR threshold.

In the calibration process, care is taken that the models perform well for different magnitude

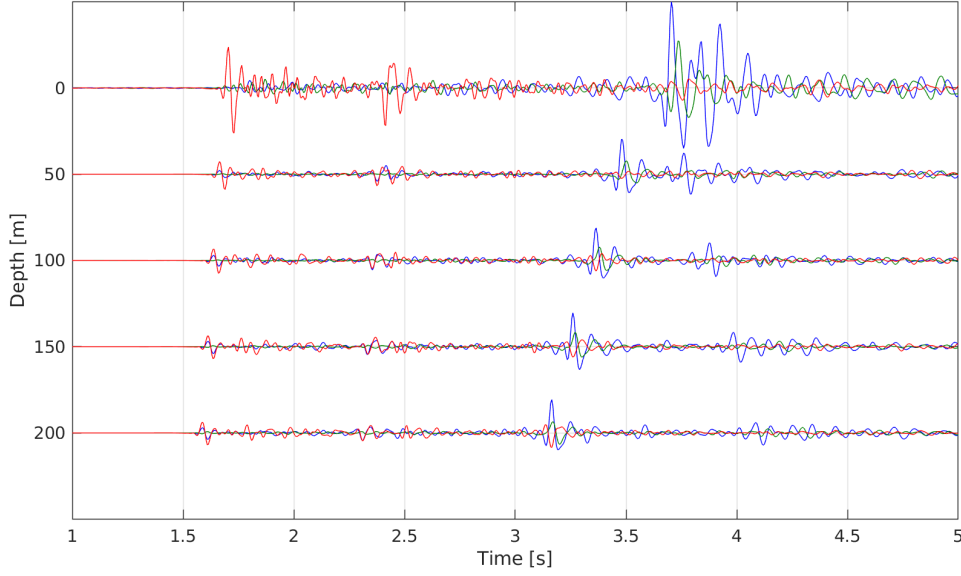


Figure 2.4: Seismic recording of the Westeremden M=1.7 event at November 8, 2011 at borehole G22, which is situated at 2.98 km epicentral distance from the source. The red, blue and green wiggles show the vertical, radial and transverse components, respectively. To all seismograms, band-pass filtering has been applied between 5 and 40 Hz.

and distance bins. In the parameter-fitting exercise, the DL database has more weight, because it contains most PGV values. Moreover, the amplitudes recorded at depth show less variability because the site response plays a relatively minor role at these depths. Fig. 2.6 shows the distance-dependent misfit between the PGV databases and the fitted models for a source depth of 3 km. The DL model shows very little bias until 40 km distance. The mean misfit from the SL model is slightly undulating as function of distance. The SL database has a lack of measurements for epicentral distances larger than 32 km, which explains the larger variability from that point onward. The DL database still has records beyond 120 km. The amplitude decay at distances beyond 40 km has been solely based on the DL database. Until 40 km the model performs very well, between 40 and 70 km the model on average overpredicts PGV whereas between 70 and 100 km there is on average an underprediction.

Fig. 2.7 shows the estimated mean models. The model values at M=4.0 are an extrapolation since the largest event in the database has M=3.6 (Huizinge event at August 16, 2012). The SL model is shown together with the average 90th percentile of the v_{RMS} noise as recorded over surface accelerometers (Section 2.3). This gives, for each magnitude, an indication of the distance range at which a detection is likely. The DL model is shown together with the v_{RMS} noise recorded at vertical component geophones at 200 m depth.

Variabilities in the model are split up into within-event variability ϕ and between-event variability τ . This yields the total standard deviation of the model $\sigma = \sqrt{\phi^2 + \tau^2}$. The variabilities are expressed in \ln and are listed in Table 2.2.

τ expresses, among others, that there is a variability in the isotropic component of the sources. τ is thus a property of the sources, although also the network configuration has some influence on how well it can be estimated. For example, the network configuration determines how well the radiation pattern is averaged out. Both the DL and SL networks yield an independent estimate of τ . The DL dataset contains a larger amount of sources (Fig. 2.3). Using recordings at 200 and 0 m depth, values of 0.2936 and 0.2872 are obtained,

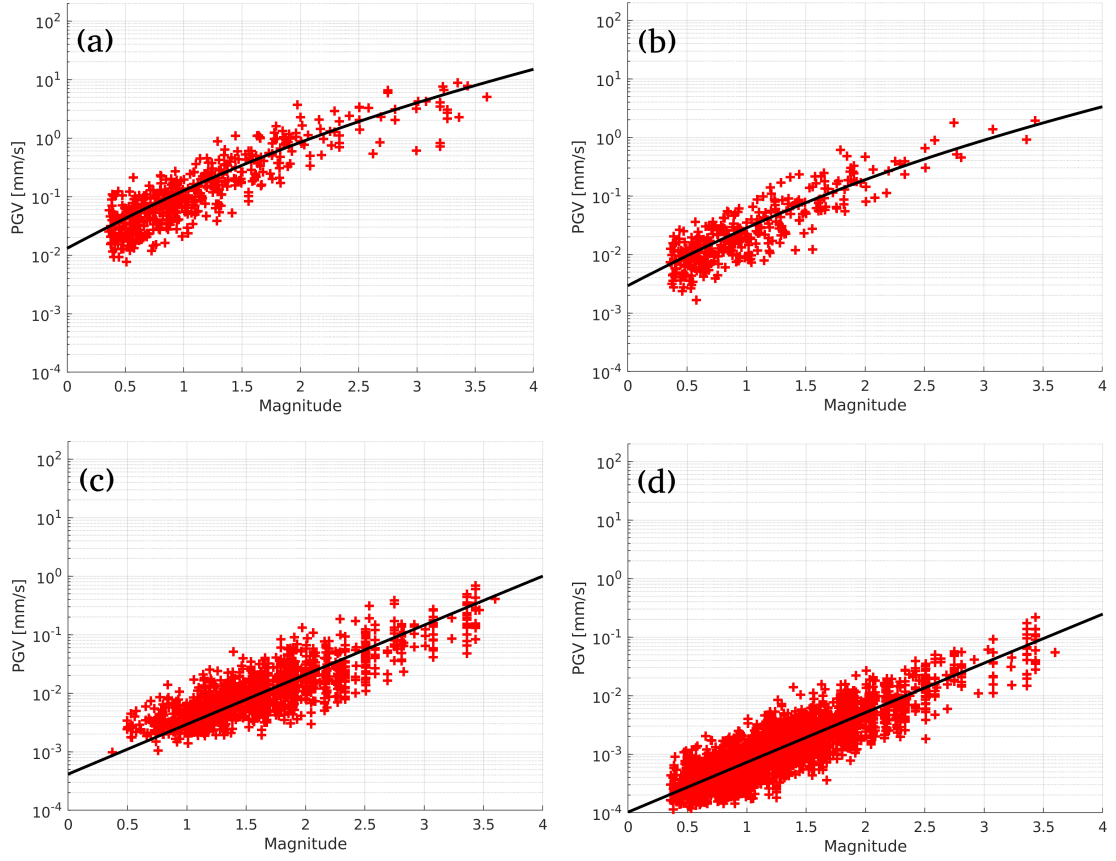


Figure 2.5: Measured PGV values (red crosses) and fitted mean models (black lines) as function of local magnitude. Panels (a) and (b) are for recordings within 2 km epicentral distance (and a model for a source at 3 km depth and 1 km epicentral distance), panels (c) and (d) are for recordings between 8 and 12 km epicentral distance (and a model for a source at 3 km depth and 10 km epicentral distance). The left panels (a,c) show surface recordings and models, the right panels (b,d) show recordings at approximately 200 m depth.

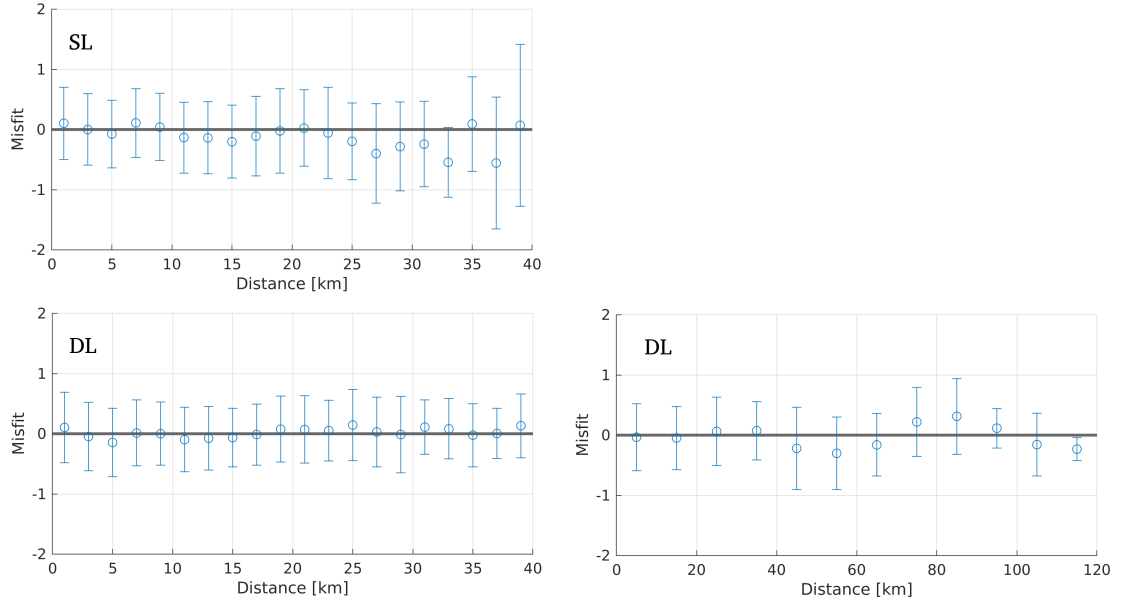


Figure 2.6: Misfit ($\ln(Y_{obs}) - \ln(Y_{mod})$) distributions between surface level (SL) and depth level (DL) recordings and their respective PGV database and the surface level (SL) and depth level (DL) GMPEs. The mean misfit (circles) and standard deviation (the bars are from -1 to +1 standard deviation) are displayed over distance bins of 2 km for the left panels. The right panel shows distance bins of 10 km. Only for the DL database there is a sufficient number of recorded PGV values to show misfits until 120 km epicentral distance.

respectively. These values are averaged to find the τ as listed in Table 2.2.

ϕ expresses the variation of amplitudes for a single earthquake and distance. These variations are largely caused by the radiation pattern of the source and variability in site amplification. Prior to obtaining ϕ , first the total variability σ is determined by 1) computing the residual between each data point and the mean model and 2) by taking the standard deviation of this residual distribution. The within-event variability ϕ is then obtained by de-averaging τ from the total variability σ . This total variability is higher for the 0m depth level and hence ϕ is larger at the Earth's surface. This is to be expected since the variability in the site effect is an important constituent in the SL model, which is hardly present in the DL model.

Coefficient	SL model	DL model
τ	0.2904	0.2904
ϕ	0.5307	0.4650
σ	0.6050	0.5482

Table 2.2: Total σ , intra-event ϕ and inter-event τ standard deviations in the PGV empirical model.

2.2.4 Quality checking

As a quality check of the model, we plot the measured and modelled PGV levels for several selected events. An example is shown on Fig. 2.8. It shows a recent event induced at

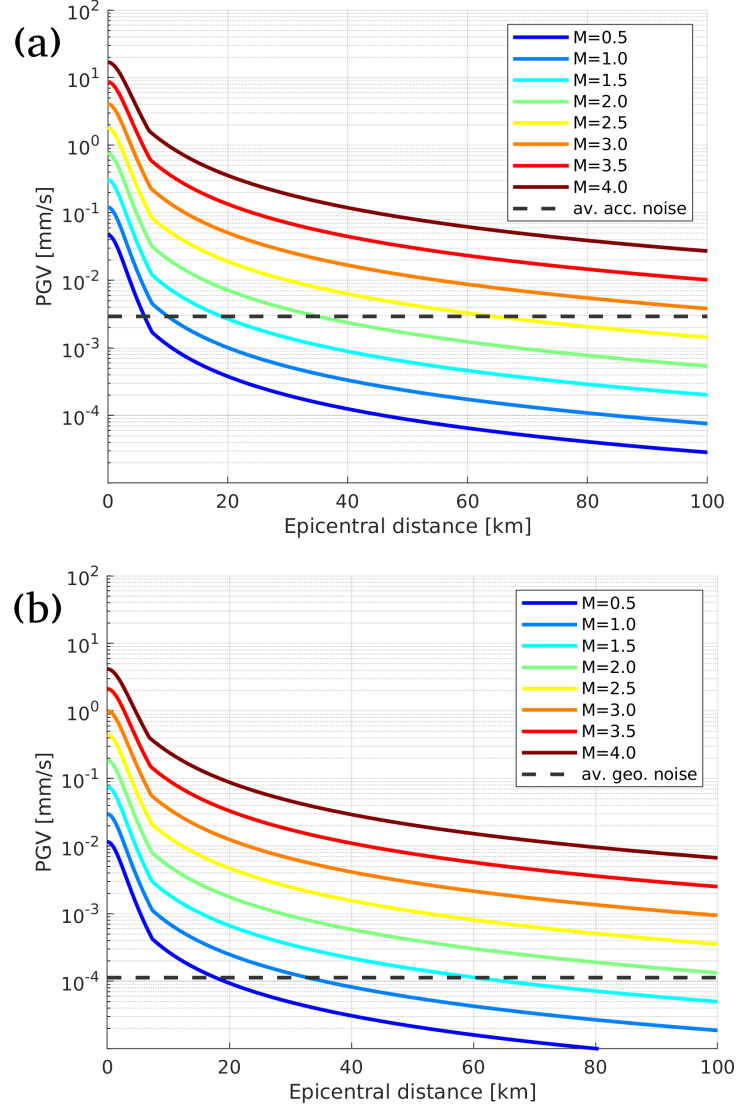


Figure 2.7: The PGV GMPEs for (a) surface level (SL) motions and (b) depth level (DL) motions. The models have been calibrated with seismicity between local magnitudes of 0.4 and 3.6. Dashed lines show, on (a), the average 90th percentile surface (accelerometer) v_{RMS} noise level and, on (b), the average 90th percentile 200 m depth (geophone) noise level.

the Groningen gas field. Nearly all measured PGVs lie within the 98th percent confidence area of the SL model (which is the area in Fig. 2.8(a) from the 1st to the 99th percentile), whereas all measured values at 200 m depth lie within the 98th percent confidence area of the DL model (Fig. 2.8b).

The derived GMPEs have been calibrated with events inside and outside Groningen (Fig. 2.3). Most of the measurements, however, are from Groningen events, which could result in a bias to Groningen-type events and subsurface conditions. If there is a profound difference in induced events inside and outside Groningen, this would result in the models being less suitable outside of Groningen. In Fig. 2.9 we check the model performance for events outside of Groningen. Non-Groningen data points are shown in black; on the background, Groningen-event data points are shown in red. The GMPEs are displayed for $M=1.5$. All data points have been mapped to $M=1.5$. This is done by

1. computing Y_{mod} (Eqs. 2.1–2.3) for the actual distance and magnitude pertaining to each data point
2. computing the Y_{mod} in case it were an $M=1.5$ event, and
3. subtracting the difference in the values computed in Step 1&2 from the measured PGV.

Fig. 2.9 shows that both the Groningen and non-Groningen data points appear well described by the derived GMPEs. Groningen and non-Groningen data points have a similar (distance-dependent) mean and variance.

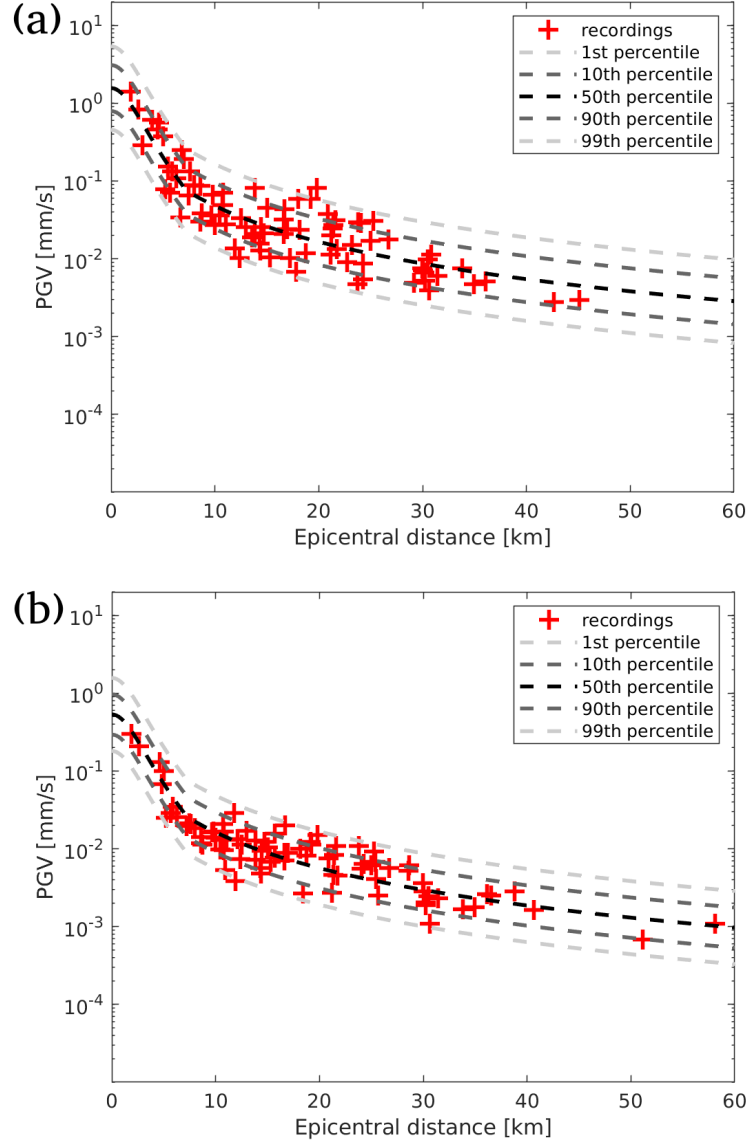


Figure 2.8: Recorded vertical-component PGV values (red crosses) and the GMPE percentiles (dashed lines) for (a) sensors at the Earth's surface and (b) sensors at depth. Data is shown for the Zeerijp M=2.51 event at October 4, 2021.

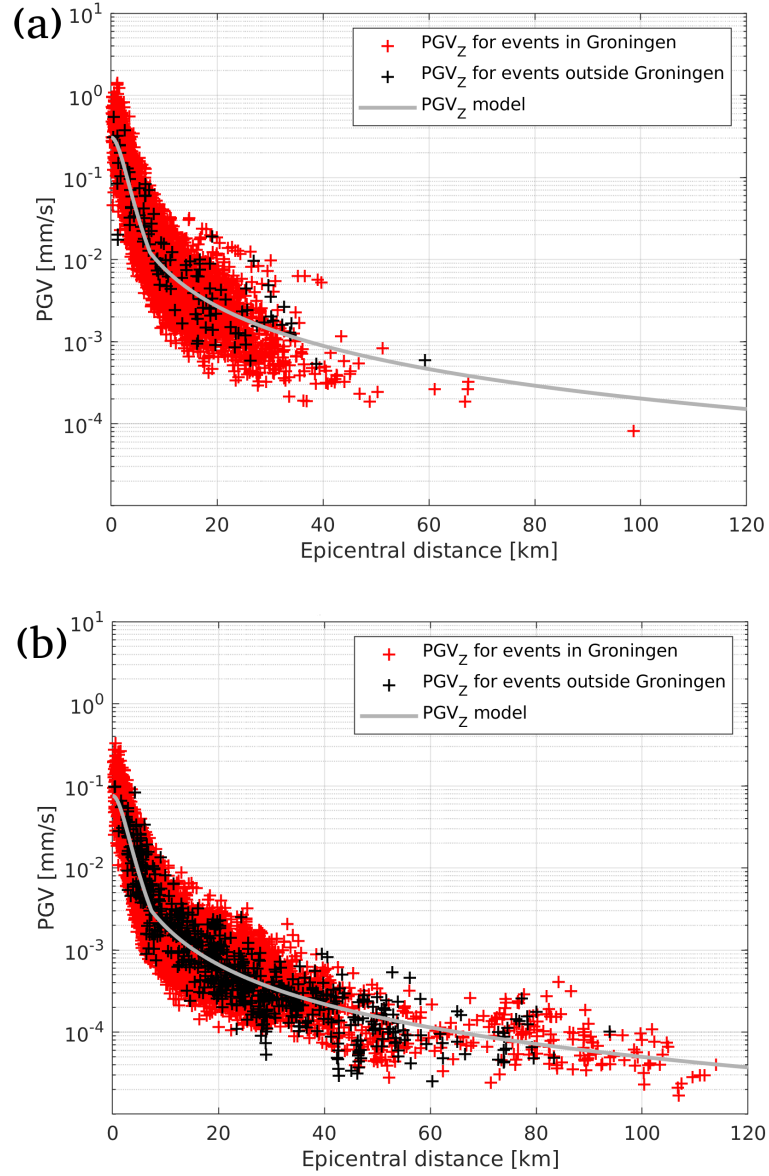


Figure 2.9: Recorded vertical-component PGV values mapped to $M=1.5$ using the derived P-wave GMPEs (Eqs. 2.1–2.3 and Table 2.1) for (a) recordings at the Earth's surface and (b) recordings at 200 m depth. At the Earth's surface there are 7592 recordings of Groningen events (red crosses) and 117 recordings of non-Groningen events (black crosses). At 200 m depth there are 15078 recordings of Groningen events (red crosses) and 675 of non-Groningen events (black crosses). The mean PGV models (GMPEs) are shown in grey.

2.3 Station-specific noise distributions

There are many different types of sources that produce seismic waves: cars, trains, explosions, wind turbines, generators, trees moving back and forth in a wind field, earthquakes, etc. From these different sources, a part is classified as transient sources, meaning that they have a short duration. Sources with a long duration, on the other hand, are called noise sources. E.g., a car passing by a seismic sensor can be seen as a transient source near this sensor. If the sensor is standing close to a highway, there will be a continuous succession of transient sources that merge into a constant source. As such, the highway is a noise source. In the current work, we are only interested in earthquakes. In this context, all other sources that produce seismic waves are characterized as noise.

In this study, we are only interested in vertical-component noise levels, because only the vertical component is used for detection. The noise levels are expressed in root-mean-square particle velocities v_{RMS} . We are especially interested in the 90th percentile levels (P90) levels. These are the levels are exceeded 10% of the time and which are used later on for generating the maps. In the following we derive, in a few steps, the P90 vertical-component v_{RMS} levels of all stations on the detection list.

2.3.1 Power-spectrum density

Seismic sensors in the NL network record continuously. The recorded wavefield is quantified by computing the power-spectrum density (PSD) (e.g., *McNamara, 2004*). At KNMI, PSDs are routinely computed, for example to quality check the instrumentation (*Koymans et al., 2021*). The approach described in the aforementioned references is used to compute and efficiently store the PSDs. These PSDs form the basis of our noise analysis.

The database described by *Koymans et al. (2021)* has since been updated to a more precise version. The former database was computed using a 1-octave smoothing window, sampling the decibel dimension with steps of 1 dB and sampling the frequency dimension with steps of 1/8 octave. The new database is computed using a 1/8-octave smoothing window, sampling the decibel dimension with steps of 1 dB and sampling the frequency dimension with steps of 1/24 octave. This updated database is used in this study.

First we retrieve continuous vertical-component time-amplitude recordings, remove the instrument response and transform the data to particle acceleration: $a(t)[m/s^2]$. Second, we use the Fast Fourier Transform (FFT) to compute the one-sided (positive frequency only) amplitude spectrum $|A(f)|$. Subsequently, the PSD is obtained with

$$PSD(f) = 2|A(f)| \frac{\Delta t^2}{T}, \quad (2.4)$$

where Δt is the sample duration of the seismic recording and T is the duration of the seismic (noise) record. The factor of two is needed when only positive frequencies are kept with $|A(f)|$. To compute a stable PSD, it is averaged over multiple partly overlapping time segments. Additionally, smoothing over the frequency dimension is applied. This results in one estimate of $PSD(f)$ for each hour. The process is repeated for all hours in the day, using a 50% overlap. Hence, 48 $PSD(f)$ functions are found (using also the first half hour of the next day). These 48 PSDs give the distribution of frequency-dependent noise levels over that day. By up-scaling this approach, also the distribution of PSDs can be found for a week, a month and a year of data. From these distributions, the probability of occurrence of a certain power level can be computed for each frequency. This yields the probability density function (PDF) of the PSD.

Fig. 2.10 show examples of PDFPSDs for one year of data. The sites chosen show broadly the noise level variation over the Netherlands.

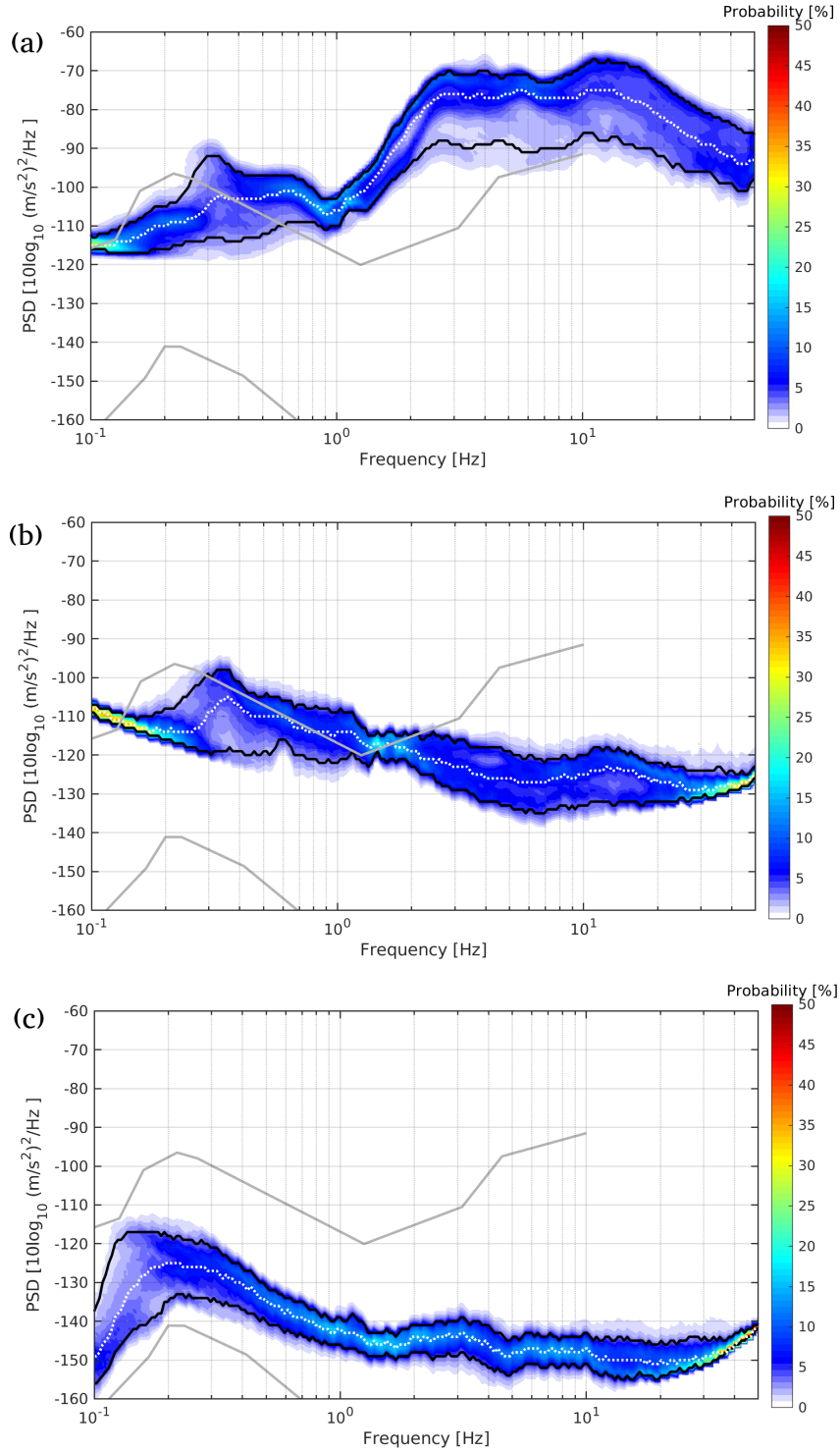


Figure 2.10: PDFPSDs (probability-density functions of power-spectrum density) for the vertical component at stations (a) VPR (surface accelerometer at Rotterdam harbour), (b) G204 (borehole geophone at 200 m depth in Groningen) and (c) TERZ.01 (broadband sensor at 250 m depth in South Limburg). Data from 2020-09-01 until 2021-08-31 is shown. From the seismic noise distribution the 10th percentile and 90th percentile are depicted with black lines, the white dotted line denotes the mean (P50). The grey lines denote a higher and lower noise reference model as derived in *Peterson* (1993).

The upper plot is from an accelerometer that is located at the surface at an industrial site in Pernis, within the Rotterdam harbour area. At this site, ground vibration levels are very high due to nearby roads (A4 within a few hundred meters), water ways (Nieuwe Maas) and large oil refineries. It can be seen that for frequencies larger than 0.4 Hz, the average noise condition (white-dotted line) is higher than the upper grey curve, which depicts the higher noise reference model from *Peterson* (1993). This shows that noise conditions in the Netherlands can be very severe in comparison with a worldwide distribution of sites with (broadband seismic) instrumentation.

The middle plot is from a sensor in Groningen, at 200 m depth, recording much lower vibration levels. A large part of the anthropogenic noise (> 1 Hz) does not reach the 200 m depth level.

The lower plot is from a broadband station in the south of Limburg, at 250 m depth and situated within hard rock. At this site, very low background ground-motion levels are recorded, which makes it quite ideal to detect (low-magnitude) seismicity.

The lower black lines in Fig. 2.10 denote the 5th percentile (P05) of the PDF; the upper black lines denote the 95th percentile (P95). At frequencies where both lines have a large separation, there is much variation in noise level over the year. Fig. 2.10 also serves to show the main differences between the sensor capabilities. Both the accelerometer and the geophone (upper two panels) reach their recording limits around 0.2 Hz. Below this frequency, the recording is dominated by instrumental noise. This instrumental noise is nearly constant over the year. This results in the P05 and P95 lines coming close together and the PDF showing a large probability (red) for a single power level. The broadband station (lower panel) still records well the actual ground-motion below 0.2 Hz, resulting in a continuing wide band of power levels below 0.2 Hz. The geophone and the broadband station show a more confined distribution at the upper end of the frequency range. This is due to the 50 Hz grid distortion that is picked up at many instruments.

The PSD is computed over a broad frequency band. For detection, a more restricted frequency band is used in which a high SNR may be expected for local and regional seismicity. For borehole geophones and (surface) accelerometers, 4th-order Butterworth filters are applied to select a band between 5 and 40 Hz. In the current operational chain, the recording from many of the broadband stations is band-pass filtered between 0.7 and 5.0 Hz prior to running a detection algorithm. These broadband stations are primarily located in the south (Fig. 2.2) with the aim to detect seismicity of a tectonic origin, both inside and outside the Netherlands. When more geothermal operations are planned in the south, a parallel production chain will be activated to gear the broadband stations also towards detecting induced seismicity, while keeping the 'tectonic' chain running at the same time. In the following, we use the 'induced' detection band-pass filter from 5.0 to 40 Hz for all sensors.

2.3.2 Root-mean-square velocity

The operational PSDs are expressed in squared acceleration per Hz and plotted in dB: $10\log_{10} [(m/s^2)^2/Hz]$. For detection, the time-amplitude recording is used which is expressed in particle velocity (m/s). In Section 2.2 models were derived that estimate a PGV level, which is also expressed in particle velocity. Thus, in order to find whether an incoming P-wave likely stands out from the noise, the same unit needs to be used, i.e. particle velocity. First, the PSD is expressed again in a linear scale: $(m/s^2)^2/Hz$. Subsequently, a division by ω^2 is needed to go from acceleration squared to particle velocity squared, then an integration over the relevant part of the frequency domain is done, and a square-root to go back from

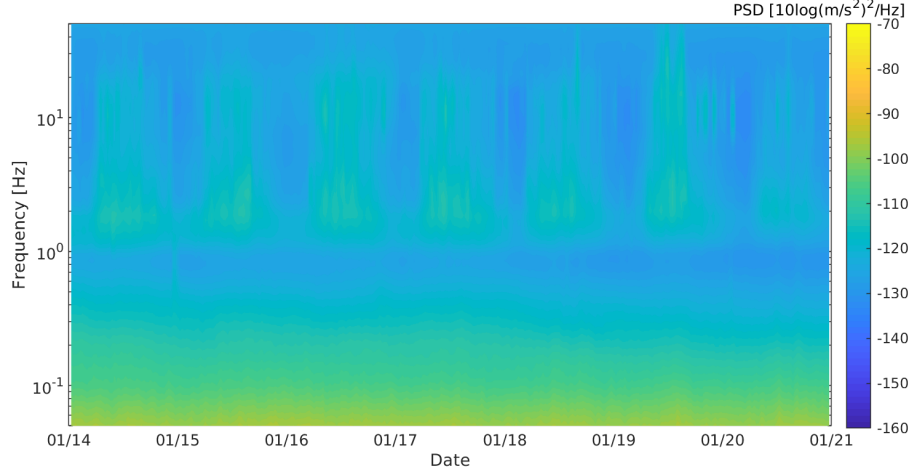


Figure 2.11: Spectrogram showing one week of power spectrum densities for the vertical component of geophone T061. Data is shown for Monday the 14th of January until Sunday the 20th of January, 2019. UTC time used.

a power to an amplitude value:

$$v_{rms} = \sqrt{\int_{f_1}^{f_2} \frac{PSD(f)}{\omega^2} df}, \quad (2.5)$$

where v_{rms} is the root-mean-square particle velocity. A similar formula can be found in *Bormann* (1998). The main difference is that we do not include a factor of 2 in Eq. 2.5 since we already implemented this factor in Eq. 2.4 to compute the PSD from a one-sided FFT.

The minimum noise that is to be expected is the instrumental noise. The newer geophones in the network (>2014) are of the type SM6H. In the detection frequency band, their instrumental noise level is $0.008 \mu m/s$. The newer accelerometers that are used in the network are of type Etna2. Their instrumental noise (including digitization etc.) is $0.048 \mu m/s$ in the 5 to 40 Hz band. These values are found by substituting the instrumental noise models of these instruments expressed in $PSD(f)$ (as published in *Koymans et al.* (2021)) in Eq. 2.5.

In the 5 to 40 Hz band, most of the noise has an anthropogenic origin. A large part can be attributed to traffic, which has a clear diurnal pattern as shown in Fig. 2.11, with quiet hours during the night and higher noise conditions during the day. For some stations, the weekend days look markedly different from week days. For that reason, at least one week of data is required to capture the variability in on-site vibration levels. One week of data does not suffice, however, for stations close to wind turbines or large trees. Wind conditions show a seasonal variability. By sampling both a more quiet and a more windy season, much of the resulting variability in vibration levels is captured. For that reason, we use at least half a year of data. For most stations, we use (almost) one year of data for quantifying the seismic noise distribution (PDFPSD), from which we compute e.g. a mean v_{rms} .

There are currently seven broadband stations not having a channel sampling the seismic record at 100 or 200 Hz. These stations in the NL network (ARCN, HGN, HRKB, MAME, OPLO, ROLD and WTSB) do have a channel that samples at 40 Hz, yielding a Nyquist frequency at 20 Hz. Due to the anti-alias filter and its effect to lower frequencies (due to

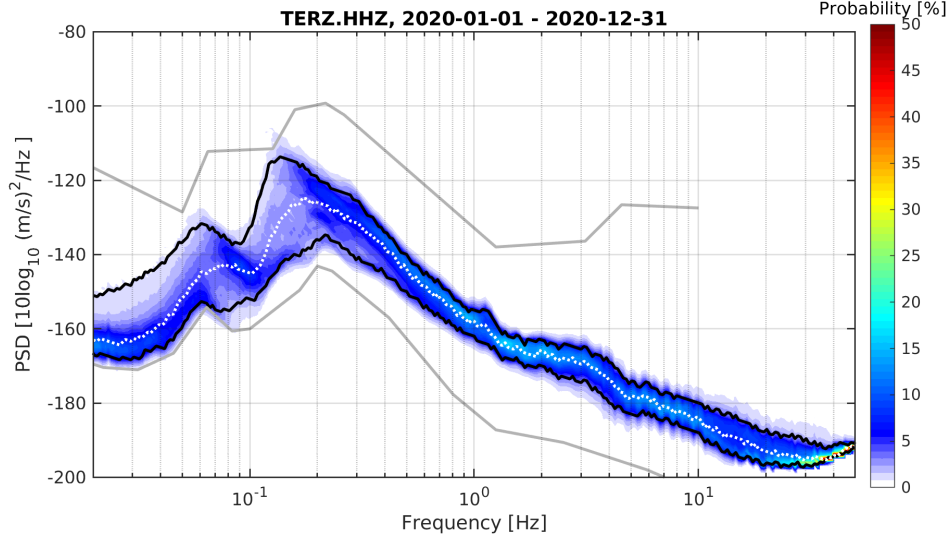


Figure 2.12: PDFPSD for the year 2020 for the vertical component at station TERZ (lower sensor). The PDFPSD is expressed in terms of particle velocity instead of particle acceleration (Fig. 2.10c). From the seismic noise distribution the 10th percentile and 90th percentile are shown with black lines, the white dotted line denotes the mean (P50). The grey lines denote a higher and lower noise reference model as derived in *Peterson (1993)*.

smoothing when computing PSDs) the frequency band for v_{rms} computation is restricted to 13 Hz. Stations G81B, TERZ.01 and DBN are used to test the influence of the restricted frequency range on the 90th percentile of the root-mean-square velocity v_{rms}^{P90} . For the year 2020, noise-level reductions of 3.7, 10.6 and 3.4% are found, respectively.

In Fig. 2.12 a PDF is plotted in terms of particle velocity (instead of particle acceleration). It shows that most of the power resides at frequencies lower than 13 Hz. This explains why the v_{rms}^{P90} is reduced only with a modest percentage when the frequency range 13 to 40 Hz is left out from the 5 to 40 Hz band. Although the frequency limitation does reduce the noise level, we do not correct for this. The reason is twofold. On the one hand, it is unknown what the proper correction for the few different broadband sites should be. On the other hand, the effect is limited.

For all stations the v_{rms} is determined using data of September 1, 2020 until August 31, 2021. Stations that have less than 50% data availability during this time period, are left out from the station list (see Iteration 3 in Section 2.1). These stations would have a less representative v_{rms} . Yet, more importantly, these stations could not be used for detection and location for more than 50% of the time.

For stations installed in 2021, the station-rejection criterion is less strict. The new stations are still included when they have produced at least two months of data. New stations that are hence included for the 2021 version are BLIJE, DON, NLDW, ZDL, ZH04. Note that the v_{rms} of especially these stations might change somewhat in 2022 when more data becomes available.

For all stations from the list (Section 2.1) we compute and store the P05, P10, P50, P90 and P95 levels of v_{rms} . Fig. 2.13 shows histograms of the resulting values. The accelerometer and geophone distributions are clearly separated. On average, noise levels on accelerometers are much higher. This louder noise condition can be ascribed to placement of the accelerometers at the Earth's surface, whereas all the geophones in the detection list are deeply buried,

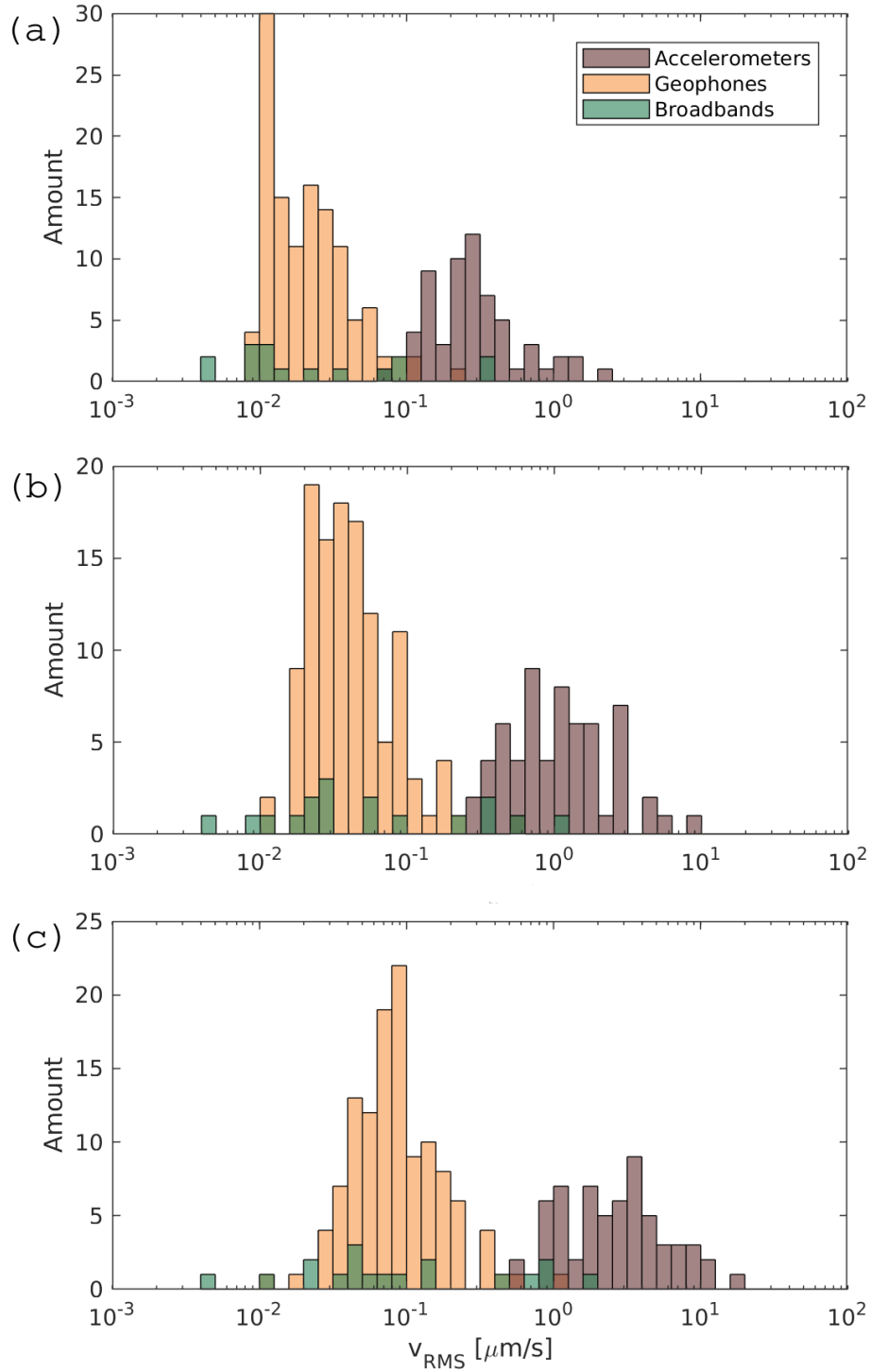


Figure 2.13: Vertical-component v_{RMS} distribution for stations from the detection list (Fig. 2.2), with the exception of stations J01 and J02. In (a) the 10th percentile of the root-mean-square particle velocity is shown, in (b) the 50th percentile (mean) and in (c) the 90th percentile.

most of them at 200 m depth. The surface short-period sensors J01 and J02 have been left out from the distribution. The broadband sensors show a mixed picture. The most quietly situated is TERZ.01, which is buried deep and in hard-rock ($v_{rms}^{P90} = 0.004 \mu\text{m/s}$, Fig. 2.10a). The highest seismic noise levels, at a broadband station, are observed at DBN, which is at the Earth's surface, on top of soft sediments and close to several busy roads ($v_{rms}^{P90} = 1.812 \mu\text{m/s}$). The overall highest noise levels are observed at a station in Pernis ($v_{rms}^{P90} = 19.55 \mu\text{m/s}$, Fig. 2.10c) which is situated within a few hundred meters of an oil production plant, a refinery, a logistical hub and a highway.

Fig. 2.13(a) shows the P10 v_{rms} distributions. The noise levels on the geophones do not show a normal distribution, but a skewed distribution with a large peak around $0.01 \mu\text{m/s}$. This can be ascribed to the instrumental-noise limit of geophones of $0.008 \mu\text{m/s}$. This implies that for some of the instruments, for a few percent of the time, not the environmental noise is the limiting factor for measuring very small earthquakes, but the instrumental noise.

Fig. 2.13(c) shows the P90 v_{rms} distributions. Both the accelerometer and geophone noise levels over the network show approximately a log-normal distribution. For the 61 accelerometers, the logarithmic mean, plus-minus one standard deviation is $0.423 \pm 0.350 \log_{10}(\mu\text{m/s})$, which corresponds to $2.646 \mu\text{m/s}$ for the mean. The logarithmic mean over the 119 geophones, plus-minus one standard deviation is $-1.063 \pm 0.308 \log_{10}(\mu\text{m/s})$, which corresponds to $0.086 \mu\text{m/s}$ for the mean. Thus, on average there is a factor of 30.6 difference (29.7 dB) in seismic noise level between the surface accelerometer and 200 m depth geophone installations, for the 5 to 40 Hz frequency band.

Fig. 2.14 shows the spatial distribution of v_{rms} . Because of multiple orders of magnitude in which noise levels vary, the amplitude scale (colorbar) is chosen as the 10-based logarithm of the noise levels. Again, mainly a difference can be seen between deep geophone sensors (denoted with triangles) and surface accelerometer installations (squares). A part of the 23 broadband stations are in a dedicated vaults and in a hard-rock setting, resulting in low noise conditions. Broadband stations in Utrecht, the east of Noord-Brabant and Limburg (DBN, OPLO, ARCN, HRKB) are on top of soft sediments and have higher seismic noise levels. Also station ROLD has quite high noise levels, which can likely be attributed to the urban setting of this station. The Groningen geophone network (tight triangle distribution at the northeastern edge of Fig. 2.14) has a dense spatial distribution. This makes it possible to notice regional differences in noise levels. The region around Groningen city and the Eemshaven is significantly louder than other parts of the network. In the Groningen city area, roads are the main source of seismic noise. In the Eemshaven area, wind turbines are the main cause of near constant seismic vibrations.

2.3.3 Noise reduction with depth

The accelerometers and geophones in the detection list are on different locations. Most of the geophone stations have been chosen outside urban areas, where there is more space to flush a temporary borehole. Some of the accelerometer sites have been chosen on purpose within urban areas, to have a measurement of peak-ground acceleration (PGA) and PGV that buildings in the area experience. Therefore, the accelerometers are likely on average on more noisy locations than the geophones. As a consequence, the noise reduction values found in the previous section are likely not caused only by differences in depth. In the following we isolate the depth dependence on the noise by analyzing only data from the G-network stations. At each such station there is an accelerometer at the Earth's surface and a string with geophones with nodes at 50, 100, 150 and 200 m depth (Dost *et al.*, 2017).

We select one year of vertical-component PDFPSD data for all 69 borehole stations of the G-network. Numbering in the G-network goes from G01 to G70, but G15 has never been constructed. G010 denotes the 0 m depth level at station G01 and G011-G014 denote

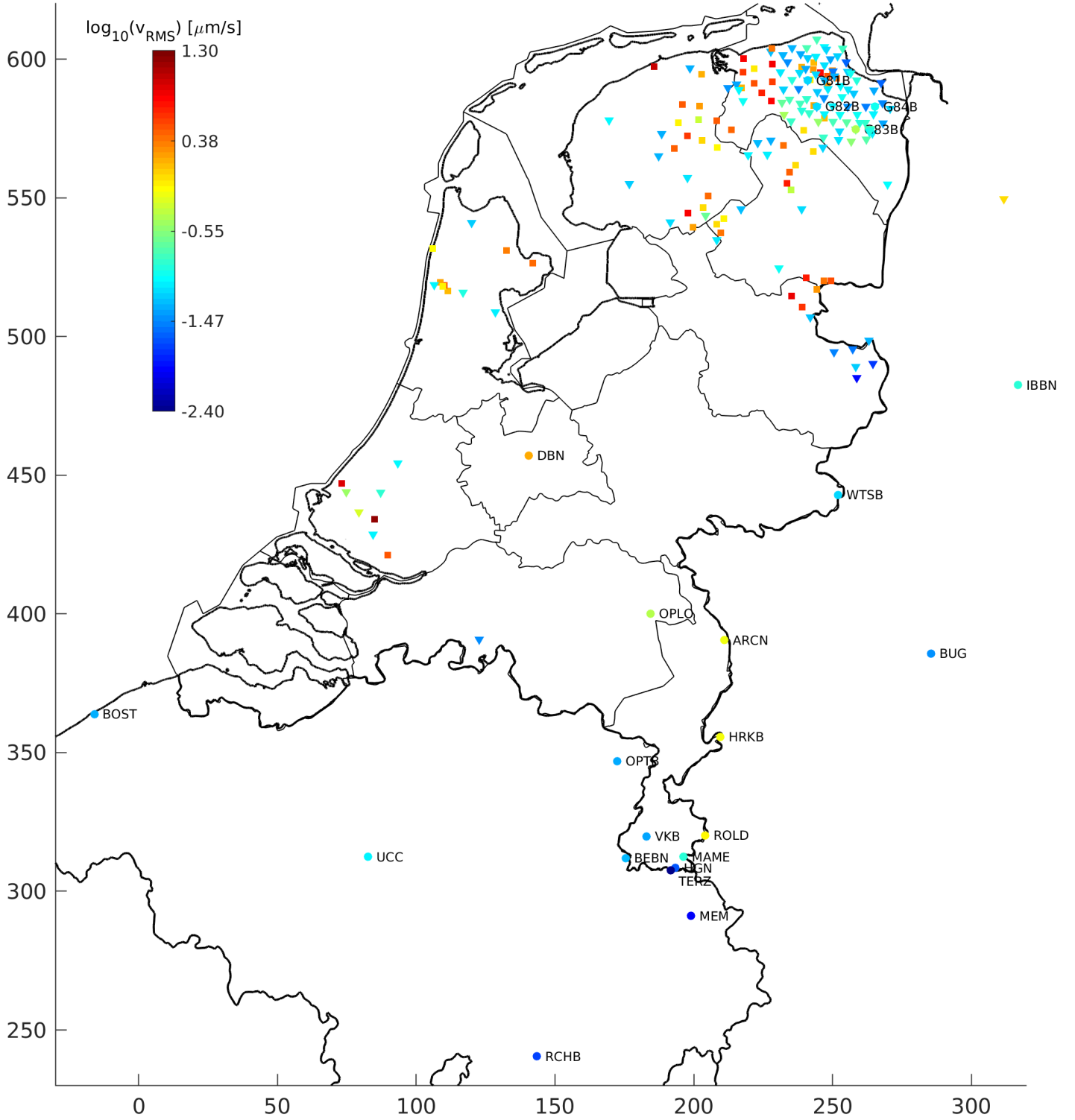


Figure 2.14: Spatial distribution of v_{RMS} . The 90th percentile of the vertical-component is shown, for stations from the detection list (Fig. 2.2). Accelerometers and surface short-period instruments are shown with squares, deep geophones with triangles and broadbands with circles. Only the names of the broadband stations are indicated. Names of other stations can be found at <http://rdsa.knmi.nl/network/NL/>. Thin black lines denote province borders within the Netherlands. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

the 50 to 200 m depth levels. As in the previous section, v_{rms}^{P90} are determined in the 5 to 40 Hz frequency band. Stations are included in the analysis when at least 50% of the data is available from September 1, 2020 until August 31, 2021. Four vertical-component sensors have issues (G050, G061, G320 and G330) or insufficient data over 2021 (G190 and G31). Therefore, data of all depth levels at G05, G06, G19, G31, G32 and G33 are excluded in the analysis, leaving 63 stations with v_{rms}^{P90} values established at all five depth levels.

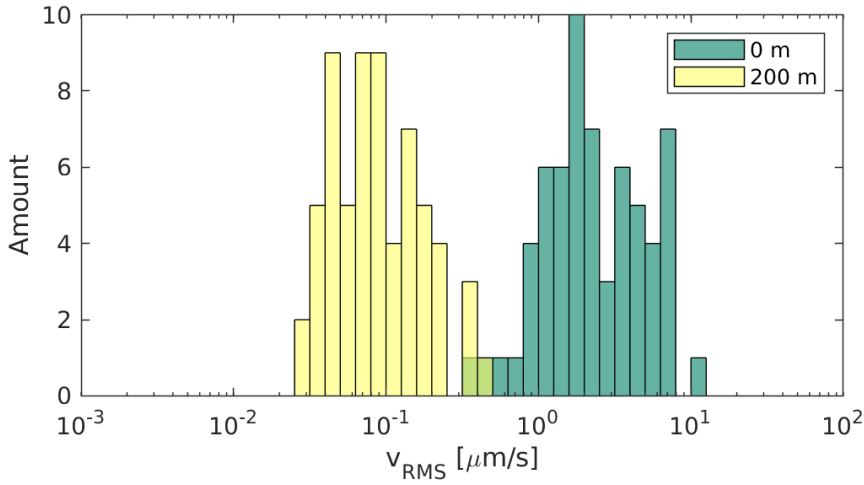


Figure 2.15: Vertical-component v_{RMS} levels for 63 stations of the G-network. The distribution is shown of the 90th percentile level, computed over one calendar year of continuous recordings, for recordings at 0 m (green) and 200 m (yellow) depth.

Fig. 2.15 shows v_{rms}^{P90} distributions for the 63 stations included in the analysis, for two depth levels. The noise levels at the Earth's surface and at 200 m depth are approximately log-normal distributions. The logarithmic means at 0 and 200 m are 2.293 and 0.088 $\mu m/s$, respectively. This yields a noise-reduction gain of 25.9 (corresponding to 28.3 dB). As anticipated, this value is smaller than the factor of 30.6 noise reduction we found in the previous section between accelerometers at the Earth's surface and geophones at 200 m depth in the detection list. Table 2.3 lists the average noise values at all different depth levels, as well as the noise-level reduction gains when installing the sensor at depth. Note that these values are only representative for unconsolidated-sediments sites and the detection frequency band.

Depth [m]	μ	σ	10^μ [$\mu m/s$]	Gain [-]	Gain [dB]
0	0.360	0.324	2.293	1.0	0.0
50	-0.696	0.299	0.201	11.4	21.1
100	-0.863	0.305	0.137	16.7	24.5
150	-0.966	0.292	0.108	21.2	26.5
200	-1.053	0.292	0.088	25.9	28.3

Table 2.3: Average vertical-component 90th percentile v_{rms} levels over different depth levels of G-network stations, in the detection frequency band. Also the standard deviation per depth level σ is listed. μ and σ are the log-normal mean and standard deviation [$\log_{10}(\mu m/s)$]. The fourth column shows the log-normal mean expressed in $\mu m/s$. The last two columns show the noise-reduction gain by placing the sensor at depth with respect to placement at the Earth's surface, expressed as a dimensionless factor and in decibels.

2.4 Picking threshold

At the KNMI, continuous data is flowing in from sensors of the NL network and nearby countries. An automatic detection algorithm is running on data streams from sensors on the detection list (Section 2.1). First it is identified whether there is a sudden rise in amplitudes. If this is found, the precise onset of this rise is found with another algorithm. The timing of this onset is called a pick. An automatic event-location module is only started up when picks are made at multiple sensors within a certain time window. In this section, we determine the picking signal-to-noise ratio (SNR) threshold. This is the average minimum factor needed between signal and noise for an automatic detection to be made. We use the following definition of SNR:

$$SNR = 20 \log_{10} \frac{\max(|s(t)|)}{\max(|n(t)|)}, \quad (2.6)$$

where $\max(|s(t)|)$ is the maximum absolute amplitude of the signal (within a certain time window). In the current application that would be synonymous to the vertical-component PGV (Section 2.2.2). $\max(|n(t)|)$ is the maximum absolute amplitude of the background noise. In Section 2.3.2, we expressed the measured noise as root-mean-square values v_{rms} . With the assumption that the noise is sinusoidal in nature so that $v_{rms} = 0.707 \max(|n(t)|)$, the minimum amplitude needed to make a pick can be expressed as

$$PGV = \frac{10^{\Phi/20}}{0.707} v_{rms}, \quad (2.7)$$

where Φ is the automatic picking SNR threshold in decibel.

An automatic detection is made with a trigger algorithm. The STA/LTA algorithm is used for the detection of induced earthquakes in the Netherlands. STA/LTA is an abbreviation of short-time-average over long-time-average (e.g., *Havskov and Alguacil*, 2016). With the SeisComP (*GEMPA*, 2022) implementation at KNMI, the STA is the average absolute amplitude over the last 0.5 s of waveform data flowing in, and the LTA is the average absolute amplitude over the last 10 s. An automatic detection is made when the STA/LTA reaches a level of 3.

The STA/LTA algorithm detects a rise of amplitudes with respect to the background-noise amplitude levels. The threshold of 3 cannot straightforwardly be translated to a SNR because the corresponding SNR depends on the mix of waveforms (signal amplitudes) that is measured within the short-time window. We determine Φ empirically by using a large collection of seismograms with different waveform distributions.

Fig. 2.16 shows the different steps in the derivation of the picking threshold. 77 seismograms of the Westerstjwerd M=3.4 event are used, recorded at 200 m depth. Due to the different distances, azimuths, near-surface conditions, etc., the seismograms show a wide variation in recorded waveforms. All seismograms are bandpass filtered in the detection band and inspected. Seismograms with noise levels visible without zooming in, are disregarded. This leaves 75 seismograms in the analysis. 15 levels of uniform noise are added to each remaining seismogram (Fig. 2.16a), corresponding to SNR levels varying from 1 to 15 dB. For each noisy seismogram, subsequently, the STA/LTA is computed and the first SNR level is determined for which the STA/LTA reaches the threshold of 3 (Fig. 2.16b). Repeating this for the 75 seismograms recorded at different stations, a distribution of picking thresholds is obtained (Fig. 2.16c). This distribution is by approximation a Gaussian distribution with mean 9.33 and standard deviation 2.33. This corresponds to the PGV being, on average, 2.92 times higher than the (peak) noise and 4.14 times higher than the v_{rms} (Eq. 2.7).

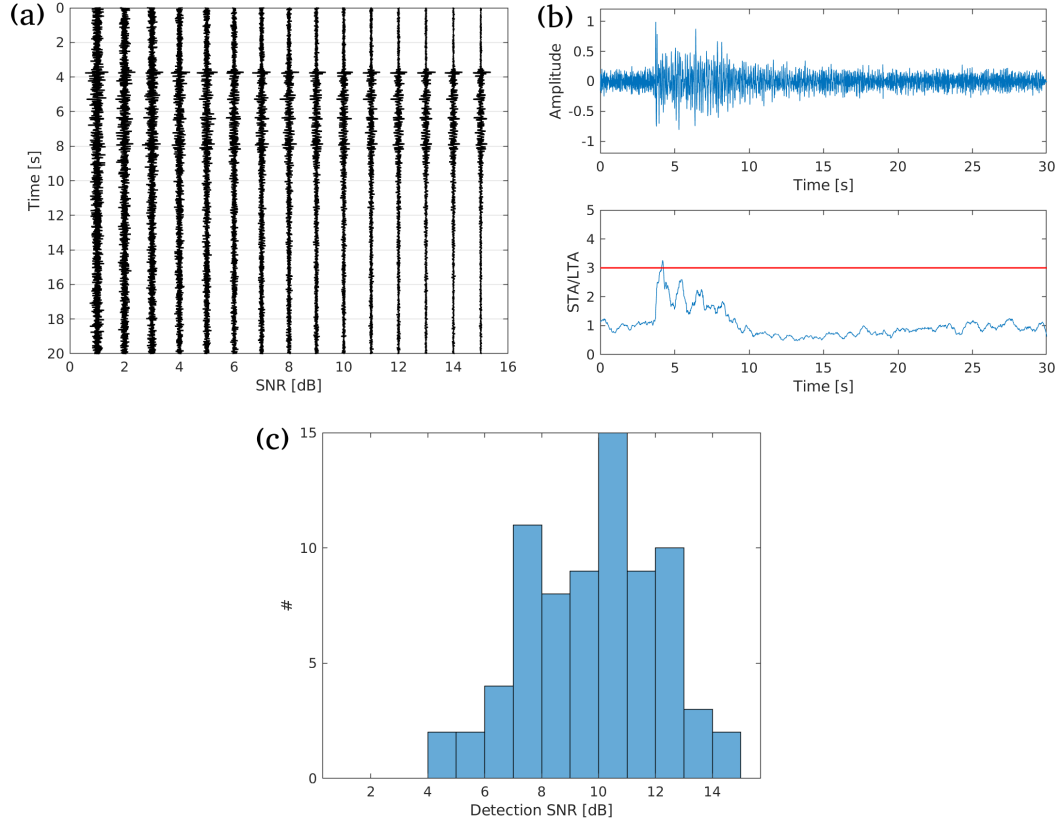


Figure 2.16: Workflow illustration for deriving an automatic picking signal-to-noise threshold. (a) G054 recording of the Westerijtwerd M=3.4 event (May 22, 2019) with different levels of uniform noise in the detection frequency band, yielding signal-to-noise ratios (SNRs) ranging from 1 to 15 dB. (b) Seismogram with lowest SNR level (10 dB) for which an automatic detection (upper panel) is obtained, meaning that the STA/LTA (lower panel) reaches a level of 3 (for details, see main text). (c) Histogram of minimum detection SNR levels at 75 different recordings of the Westerwijtwerd event.

Fig. 2.17 shows, for each magnitude, the distance until which automatic picks are likely feasible. As noise conditions vary from station to station, this distance range is station dependent. Moreover, for each station also the noise conditions vary over time. Per station, 5 lines are shown to visualise the impact of different noise conditions.

With low noise conditions (v_{rms}^{P05} resulting in the dark green lines) the epicentral distance at which a pick can be made is considerably larger than for high-noise conditions (v_{rms}^{P95} resulting in the yellow lines). Making a pick at a distance larger than indicated by the P05 line is very unlikely, because noise conditions lower than P05 are very unlikely. On the other hand, having a non-pick at distances smaller than indicated by the P95 line is very unlikely, as noise conditions worse than P95 are very unlikely.

The most important line for our analysis is the P90 model as we are using this model to decide (in the next chapter) whether a pick can, or cannot be made. To find out whether the model predictions make sense, we compare them with actual picks. Fig.2.17 shows, additionally, distances and magnitudes combinations where detections could be made (black crosses) and where not (grey circles), using the STA/LTA algorithm. When checking the picks that could actually be made (black crosses) there is indeed no pick at distances larger than forecasted with the P05 lines. Most picks are present below the P50 lines. Below the distances indicated by the P90 lines, almost all picks can be made (i.e. many crosses and little circles). This shows that the P90 model yields little false positives. That is, it is only rarely the case that this model predicts that a pick can be made, while in reality it cannot be made.

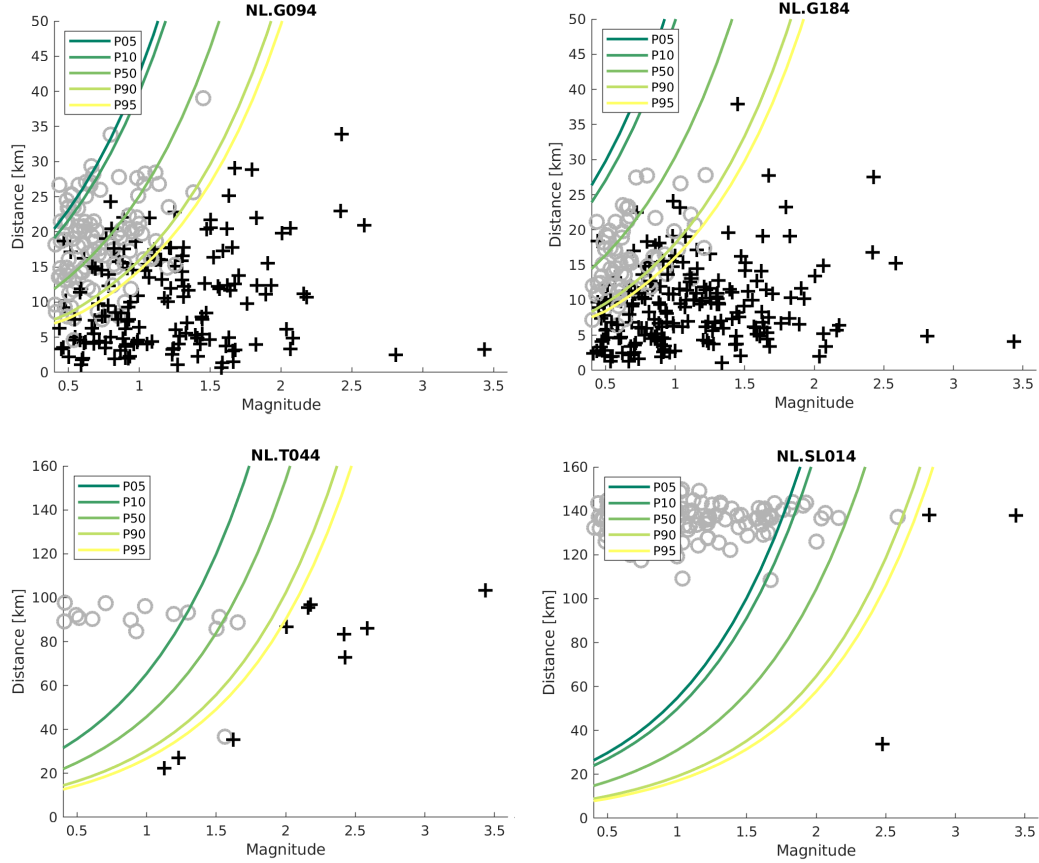


Figure 2.17: Distance-versus-magnitude picking models, together with actual picks (black crosses) and non-picks (grey circles). The models are shown for five different noise levels, represented by the percentile levels of the noise distribution: P05, P10, P50, P90 and P95. At distances larger than indicated by the models (colored lines), likely no pick can be made, given the specific noise condition. At smaller distances, a pick is likely feasible. The model predictions are compared with distance-magnitude combinations at which a pick could actually be made, with a STA/LTA algorithm, on recorded Z-component waveforms (black crosses) or not (grey circles). The comparison is shown for four recording sites at different distances from the centre of gravity of induced events (which lies in Groningen). Models and picks at 200 m depth geophones are used as they have less false-positive picks due to transient noise. Picks and non-picks are shown for waveform data from the years 2016, 2017 and 2018. At sensor T044, the P05 and P10 lines are on top of each other.

2.5 Hard-rock correction factors

In Section 2.2 two P-wave GMPEs were developed, one for motions at the surface (surface-level or SL model), and one for motions at about 200 m depth (depth-level or DL model). Both models have been calibrated with measurements in the north of the Netherlands. All the used sensors are situated in, or on top of, unconsolidated sediments. Hence, these models are not valid for hard-rock sites. Factors need to be found that express how much weaker ground motions are on typical hard-rock sites in the Netherlands than on soft-rock sites (in the 5-40 Hz frequency band). Here we determine two hard-rock correction factors, one for the deep- and one for the surface-level motions. Fig. 2.18 illustrates the rationale behind these factors.

The average P-wave velocity at 200 m depth in Groningen is $v_{sd}=1900$ m/s (*Hofman et al.*, 2017) and density is $\rho_{sd}=1700$ kg/m³ (*Kruiver et al.*, 2017). The subscript sd is used to denote it are values for the typical unconsolidated sediment. For non-Groningen north-Netherlands buried sensors, similar Neogene sediments are present and hence similar elastic conditions exist at 200 m depth (*van Ginkel et al.*, 2022a). In the Netherlands, hard-rock conditions are only found within 200 m depth in the Achterhoek (Fig. 2.14, area around station WTSB) and in the south of Limburg (area around station VKB and further south from this station). For hard-rock conditions at depth (sensors buried more than 40 m deep) we assume a sandstone with a P-wave velocity $v_{hr}=3700$ m/s and a density $\rho_{hr}=2200$ kg/m³ (e.g., *Romijn*, 2017). Assuming further that the impedance difference is reached over a gradient, the hard-rock correction factor Ψ can be found with

$$\Psi = \sqrt{\frac{\rho_{hr}v_{hr}}{\rho_{sd}v_{sd}}}. \quad (2.8)$$

Filling in the aforementioned values results in $\Psi_{DL} = 1.6$, where the subscript DL shows that this is for deep-level recordings. This hard-rock correction factor is applied to all deeply buried Belgium stations in the detection list (BOST, OPTB, RCHB and UCC) and TERZ.01 in the Netherlands. Hence, the modeled motion (Y_{mod} in Eq. 2.1) is divided by Ψ_{DL} to obtain an equivalent PGV at a deep hard-rock site.

In Section 2.2.2 it was derived that the SL model has amplitudes 4.06 times larger than the DL model. At hard-rock sites, the amplitude difference between DL and SL motions is different. The free-surface effect remains present, yielding a factor of two. The amplification caused by impedance reduction towards the surface and resonances of near-surface layers, is mostly less at hard-rock sites (without a severely weathered layer near the surface) than at soft-rock sites. Therefore, a total factor of 2.5 is assumed between the hard-rock depth-level motion and the hard-rock surface-level motion. As a consequence, there is a factor of 2.6 difference in amplitudes between the SL sedimentary sites and the hard-rock sites: $\Psi_{SL} = 2.6$. This factor is implemented for stations at, or close to, the surface in Belgium (BEBN and MEM), the German broadband stations in the detection list (IBBN and BUG) and a few NL broadband stations (WTSB, VKB, MAME and HGN).

For station NL.HGN we assess the impact of applying the hard-rock correction factor. Fig. 2.19 shows distance-magnitude combinations for which automatic picks could, or could not, be made. The station is a few hundred kilometers south from induced seismicity in the north of the Netherlands. In Fig. 2.19(a) the distance-magnitude picking models (Section 2.4) are plotted without applying $\Psi_{SL} = 2.6$. Beyond the P90 curve there are 16 picks and 12 non-picks. This means that this model is much too optimistic in forecasting that a pick can be made. Fig. 2.19(b) shows the models after applying the hard-rock correction factor. Beyond the P90 curve, there are now 10 picks that could be made, and two that could not be made. Hence, this model yields a more realistic forecast of the likelihood that a detection can be made.

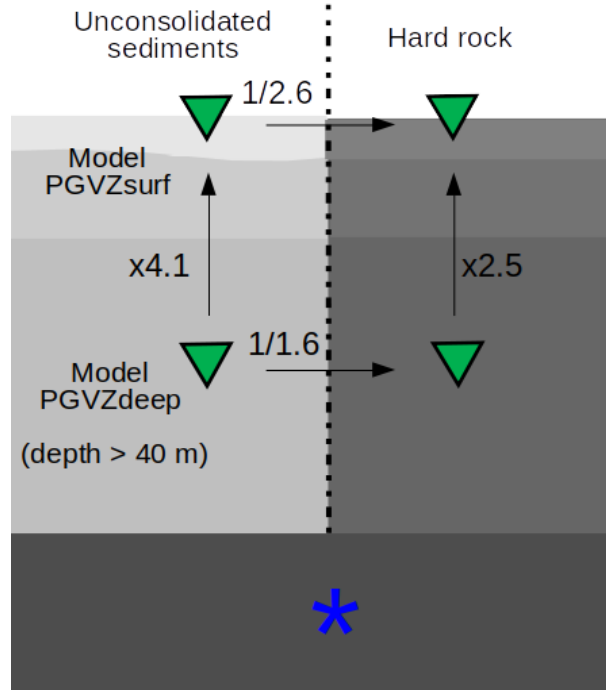


Figure 2.18: Sensors (green triangles) at hard-rock sites experience smaller ground motions than at unconsolidated sedimentary sites, for the same earthquake (blue star). This is due to larger seismic impedance (darker grey colour) at the hard-rock sites together with smaller near-surface amplification. The figure shows factors used to translate the derived models for sedimentary sites (Section 2.2), to hard-rock sites (at horizontal arrows) and the mean amplification factors in both settings (at vertical arrows).

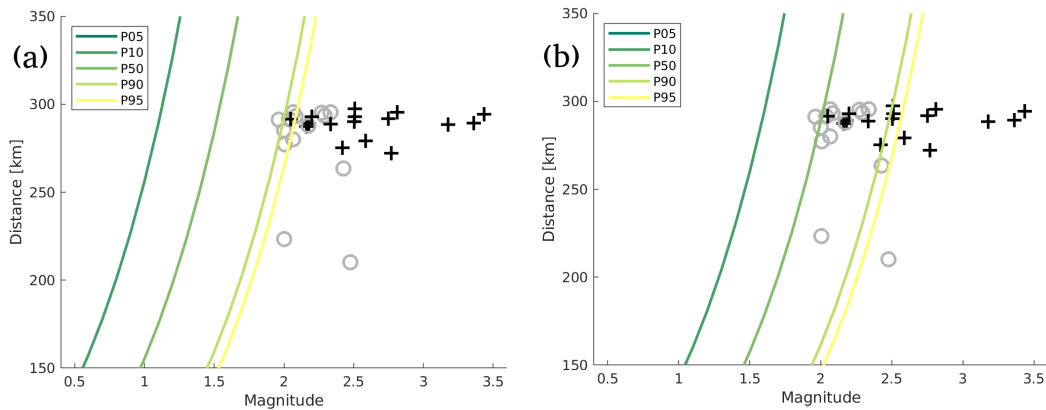


Figure 2.19: Distance-versus-magnitude picking models, together with actual picks (black crosses) and non-picks (grey circles) for (a) station HGN and (b) station HGN after applying the hard-rock correction factor. Induced events from the years 2016—2021 are used with $M \geq 2.0$. An STA/LTA algorithm is used for detection. The P05 and P10 models are visually identical for station HGN.

2.6 Travel-time models

In this section we discuss travel-time models that well fit the observations from induced seismicity in the Netherlands. Like for the P-wave GMPEs in Section 2.2, both mean models and their variability (or uncertainty) can be derived. The mean models are useful for different applications. E.g, for defining time windows in which the direct P- or S-wave arrival is expected. In this report, however, we are especially interested in the uncertainty of the travel-time model. It is the uncertainty in travel time that is mapped to uncertainty in source location. Hence, the travel-time uncertainty is one of the main ingredients for deriving the earthquake uncertainty maps,

For locating earthquakes, P- and/or S-wave velocity models are needed. In the operational practise at KNMI, 1D velocity profiles are used and —for the Groningen region— hypocenters are fixed at 3 km depth. After picking arrival times of P- and S-waves (t_{obs}) from an induced event, the actual location algorithm is started up. The algorithm uses pre-computed travel times (t_{mod}). E.g, raytracing is used to derive travel-time models from the velocity models. Subsequently, it is assessed which source location yields the best fit between the observed and modeled travel times. Even for the best-fitting source-location, there is a misfit between the modeled travel times and the picks. The remaining residuals ($t_{obs}-t_{mod}$) are a measure of the uncertainty in travel times. The travel-time model cannot perfectly explain the observations. Possible causes are related to:

- **Lateral heterogeneity.** In the assumed model there are no lateral variations in velocity structure. In reality there are, of course.
- **Picking errors.** The picked arrival times are not necessarily the true arrival times of first P- or S-waves. In the presence of noise, it could be difficult to pick the precise onset of the wave. Furthermore, it could be that a secondary arrival was picked instead of the assumed first arrival.

In this section we derive empirical travel-time models, and their uncertainty. We take advantage of the rich database of induced events in the Netherlands. The advantage of these empirical relations is that they are intrinsically based on arrival times that can be well picked. Since the models are directly expressed in travel times, instead of velocity, no forward modelling is needed to derive travel times. Hence, numerical modeling error¹ is left out of the equation. We estimate travel-time uncertainties for two situations; for the situation that there are not sufficient local picks and one needs to resort to a background travel-time model (Section 2.6.1) and for the situation that there is sufficient data to extract an event-specific travel-time model (Section 2.6.2). In the last section, a travel-time model for the near range is discussed.

2.6.1 Background polynomial model

For three years of data (2016-2018) we select all induced events with magnitudes of 1.5 and higher. This includes events from Groningen, Friesland, Drenthe and North Holland. We use the operationally hand-picked P-wave arrival times (3048 picks). This gives an overview of actually observed arrival times as function of epicentral distance (Fig. 2.20). For all these events, the approximate source depth is 3 km. Subsequently, these arrival times are least-squares fitted with a polynomial of degree 2:

$$T_P = 0.725 + 0.204R - 2.55 \cdot 10^{-4}R^2, \quad (2.9)$$

¹Numerical techniques to forward model arrival times have their limitations. E.g., raytracing cannot forward model all possible phases. With a more advanced technique, phases might be forward modeled that in reality are too weak to be picked.

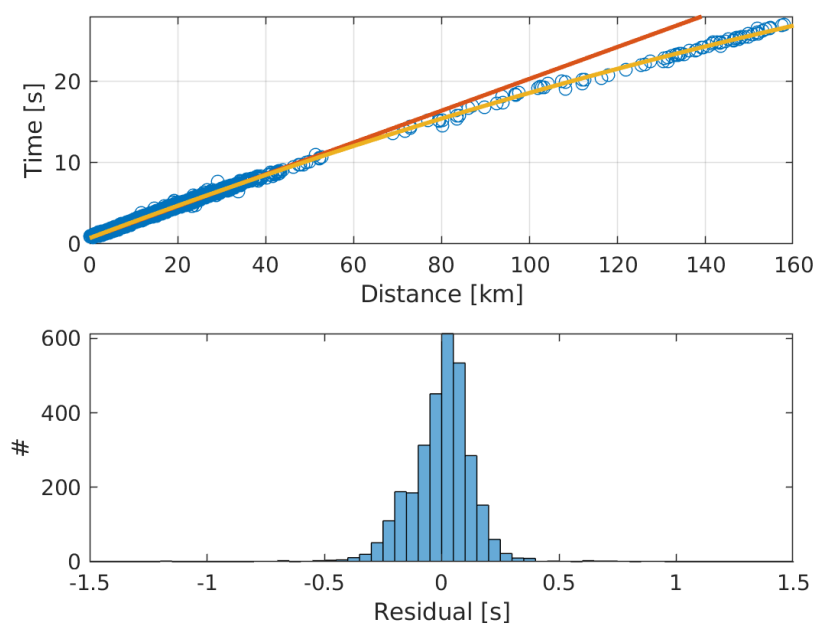


Figure 2.20: Ppick distributions for induced events in 2016-2018. The upper panel shows the Ppicks (blue circles) as function of epicentral distance and time, together with a linear travel-time model (straight red line) and a degree-2 polynomial least-squares fit (curved yellow line). Until about 40 km the travel times can be well explained with a linear model. The lower panel shows the residual distribution with respect to the degree-2 polynomial model.

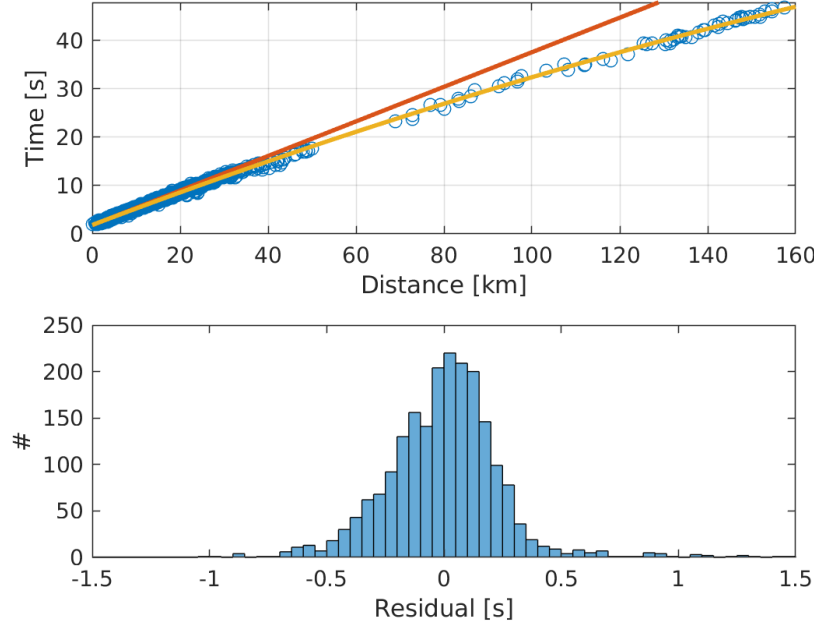


Figure 2.21: Spick distributions for induced events in 2016–2018. The upper panel shows the Spicks (blue circles) as function of epicentral distance and time, together with a linear travel-time model (straight red line) and a degree-2 polynomial least-squares fit (curved yellow line). Until about 20 km distance the travel times can well be explained with a linear model. The lower panel shows the residual distribution with respect to the degree-2 polynomial model.

where R denotes epicentral distance. The lower panel in Fig. 2.20 shows the residual distribution, which has a standard deviation of 0.137 s and a mean that is practically zero ($6.23 \cdot 10^{-11}$ s).

For the same collection of events, we add 2066 hand-picked S-wave travel times. Also the S-wave picks are fitted with a degree-2 polynomial function (Fig. 2.21):

$$T_S = 1.855 + 0.344R - 3.89 \cdot 10^{-4}R^2, \quad (2.10)$$

The lower panel in Fig. 2.21 shows the residual distribution, which has a standard deviation of 0.248 s.

In the preceding we estimated average travel-time uncertainty for a large epicentral distance range, from 0 to 160 km. In practise, however, travel-time uncertainty is a function of distance. At small distances, the travel-time residuals remain small as the travel times are small. With larger distances, a few percent velocity deviation results in larger residuals. For this reason, smaller location uncertainty can be reached when a local network is used instead of a regional or national network.

For the quantification of the distant-dependent travel-time uncertainty we split up the distance range in three bins: local [0 20], regional [20 60] and national [60 160] km. Subsequently, using only picks that fall in the respective bins and using the derived P- and S-wave travel-time models (Eqs. 2.9&2.10), we compute the residuals and their standard deviations. The results are listed in Table 2.4, together with the previous result for the entire distance range. As expected, the uncertainty increases with distance.

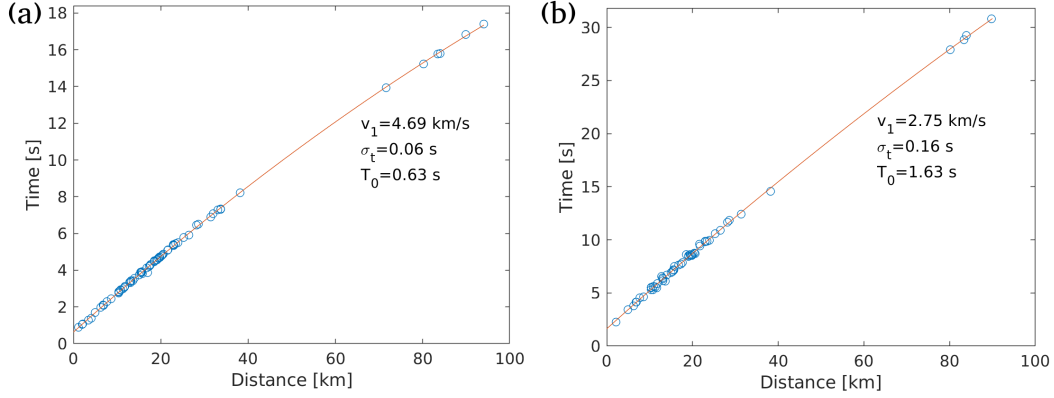


Figure 2.22: Picks of (a) P-wave arrival times and (b) S-wave arrival times for the Froombosch M=2.4 event on February 25, 2016. The picks are fitted with a degree-2 polynomial function, from which the intercept time T_0 , the apparent velocity at small offset v_1 and the standard deviation of the residuals σ_t are shown in the panels.

Travel-time uncertainty in Table 2.4 is determined with picks until 160 km distance. Picks are sometimes possible for even larger distances. For example, the larger events in the Groningen gas field can be detected at stations in the south of Limburg. These picks are not used, however, for location. There is a sufficient number of picks available at smaller distances with lower travel-time uncertainty.

Naming	Distance range [km]	σ_P [s]	σ_S [s]
Local	0 – 20	0.115	0.186
Regional	20 – 60	0.162	0.322
National	60 – 160	0.295	0.568
All	0 – 160	0.137	0.248

Table 2.4: P-wave travel-time uncertainty σ_P and S-wave travel-time uncertainty σ_S for four distance bins.

2.6.2 Event-specific polynomial model

The uncertainty of the travel-time models (Eqs.2.9&2.10) is largely caused by lateral heterogeneity in the subsurface velocity structure. This uncertainty is larger when using an average travel-time model for different regions in the Netherlands, like in the previous section. If a sufficient number of local or regional picks is available, then an event-specific travel-time model can be derived. A good sampling of the regional velocity structure, and its uncertainty, is achieved when the source is well surrounded by stations. As a measure of 'well-surroundedness' the largest azimuthal gap ψ is used. Only when $\psi < 120^\circ$ (Bondár *et al.*, 2004) and at least 12 picks are available at distances smaller than 40 km, an event-specific model is derived.

Fig. 2.22 shows examples of event-tailored travel-time models, both a P-wave and S-wave model. For the years 2016–2018, 53 induced $M \geq 1.5$ events are available of which 50 satisfy the pick-number and $\psi < 120^\circ$ condition. For each of these events, the standard deviation of the residuals is determined. The average P uncertainty (average over the standard deviations) is 0.0899 s. The S uncertainty is again almost twice as large, with a value of 0.170 s.

2.6.3 Hyperbolic travel times

The above travel-time models were derived for sources at approximately 3 km depth. For the entire distance range, it was assumed that the travel-time models can be described by degree-2 polynomials. These polynomial models would be fair to describe refracted and diving waves. Polynomial models are less suitable, however, to describe up-going waves near the source. The travel-times for these waves have a hyperbolic move-out and can be approximately described with

$$T = \sqrt{(D/v_0)^2 + (R/v_0)^2}, \quad (2.11)$$

where D is the depth of the event, R the epicentral distance and v_0 an average overburden velocity near the source. The hyperbolic move-out describes a strong curvature of the travel-time function in the first few kilometers epicentral distance. In the following we show both with actual travel-time picks and modelling that a hyperbolic behavior is a fair description at near offset.

Fig. 2.23 shows picks for the near-offset range (same picks as in Figs. 2.20&2.21). The picks show a somewhat curved line from zero to approximately 3 km epicentral distance. At further distances ([3 20] km) the travel times are well approximated with the nearly linear trajectory of a degree-2 polynomial travel-time model.

If the hyperbolic model (Eq. 2.11) is used to fit the picked travel times, then for the [0 20] km distance range the obtained uncertainty is similar to the polynomial fit (Table 2.4). Only for the first few kilometers distance, the hyperbolic model yields a better description. The variability in observed travel times in the first few kilometers (the hyperbolic part) epicentral distance is still similar to the variability at somewhat larger distances (e.g., around 8 km distance in the linear part, Fig. 2.23).

The picks in Fig. 2.23 are compared with modelling results generated with a 1D velocity profile at Huizinge, a small village in Groningen. This profile is extracted from a detailed velocity model of the region (NAM 3D velocity model; *Romijn (2017)*). Subsequently, for sources at varying depths and receivers at 200 m depth, the wavefield is forward modelled with a visco-elastic finite-difference scheme (*Robertsson et al., 1994*). The direct P- and S-waves are then picked with a kurtosis-based picker (*Langet et al., 2014*) and stored in a travel-time table. Fig. 2.23 shows the resulting travel-time curves for sources between 1 and 5 km depth. These travel-time curves are shifted with 0.07 seconds so that the 3 km depth curve fits quite well with the mean of the travel-time picks. Thus, the used Huizinge region profile results in a somewhat biased travel-time model for the collection of picks from varying regions.

For depths other than 3 km, we do not have a good empirical model. There may be enough data points to estimate the polynomial part of the travel-time model, as in the previous sections. However, it is unlikely that the near-offset range is well covered with seismic stations. Generally, one needs to resort to a local 1D profile to estimate the travel times at short offsets, and specifically to find the influence of source depth on these travel times.

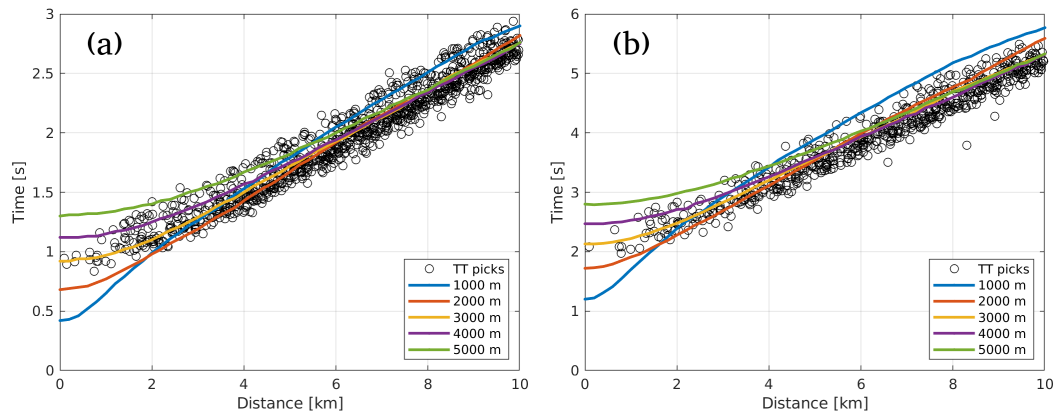


Figure 2.23: (a) Ppick and (b) Spick distribution for induced events in 2016–2018, together with travel-time curves. The curves have been generated for a realistic 1D velocity profile in the Groningen region (Huizinge), for source depth varying between 1 and 5 km. By approximation, the curves are hyperbola.

3 Location-uncertainty method

3.1 Bayesian framework

Different data attributes can be used to locate a source: arrival times of different phases (P, S, Pg, PmP, etc.), polarization of P waves and relative amplitudes over the network for different phases. In this report we restrict ourselves to timing data attributes, and specifically to the timing of the first arriving P and S waves. It is also this timing that is used in an operational setting, and for which the uncertainty has been quantified. In Section 2.6 we derived mean travel-time models and quantified the variability thereof. Since the residual distribution was approximately a normal distribution, this variability could be quantified with a standard deviation, one for P-wave timing σ_p and one for S-wave timing σ_s . These uncertainties are used in the inversion and are placed in a data covariance matrix \mathbf{C}_d .

The location of an event is found by taking a misfit between observed and forward modeled travel times, or travel-time differences. When doing this in a grid search over 3D space, a most likely location is found where the misfit reaches a minimum value. Moreover, a cloud of points is found where the misfit is still reasonably low. Because the timing is uncertain, locations around the most likely location cannot be discarded; they remain likely given the chance that, e.g., the actual velocity structure is a bit different from the model that was assumed. In a Bayesian framework, the misfit together with a data-error model are used to compute the probability of each point being the actual source. After knowing these probabilities and normalizing the sum of all probabilities to one, source-confidence volumes can be extracted. The assumption of normally distributed data errors allows using the following formulation from *Tarantola* (2005) to map errors in timing to the 'a posteriori' location probability density function (PDF, $\sigma_M(\mathbf{x})$):

$$\sigma_M(\mathbf{x}) = k \exp\left(\frac{-1}{2}[\mathbf{d}^{calc}(\mathbf{x}) - \mathbf{d}^{obs}]^T \mathbf{C}_d^{-1} [\mathbf{d}^{calc}(\mathbf{x}) - \mathbf{d}^{obs}]\right), \quad (3.1)$$

where k is a normalization constant, $\mathbf{x} = (x_1, x_2, x_3)$ are the Cartesian coordinates spanning up the model space and T denotes the transpose. \mathbf{d}^{obs} is a vector that contains the data attributes for the different receivers, or receiver combinations. $\mathbf{d}^{calc}(\mathbf{x})$ is the forward modeled data attribute for a source at position \mathbf{x} in the model. \mathbf{C}_d is the data covariance matrix and \mathbf{C}_d^{-1} is the inverse thereof. $\sigma_M(\mathbf{x})$ is computed numerically by doing a grid search over the three space variables.

In the following sections we work out Eq. 3.1 for two different data attributes, and a combination of these two. Similar approaches of computing location error ellipsoids are worked out in *Lomax et al.* (2009), *Freudenreich et al.* (2012) and *Kinnaert et al.* (2016).

3.1.1 P delay

From the travel-time measured at one receiver, an equidistant curve can be drawn. This is a surface in 3D with possible source locations. With three receivers, three such equidistant

curves can be drawn and their intersection (a single point in case there is no noise) is the hypocenter location in 3D space (x_1, x_2, x_3) . What is measured, though, are no (P-wave) travel times, but arrival times. The arrival time t_p is the addition of the earthquake origin time t_0 and the travel-time tt_p . Thus there are four unknowns, (t_0, x_1, x_2, x_3) for which a 4D PDF could be constructed. However, location and time are linked variables. Given a certain velocity model, there is one t_0 that corresponds to each \mathbf{x} . Thus, there is no need to compute a 4D PDF. t_0 can be obtained in a post-processing step.

In this report, we constrain ourselves to the location problem; we do not invert for t_0 . Taking arrival-time differences is equivalent to taking travel-time differences and it is an effective approach to take t_0 out of the equation, e.g., $\Delta t_{p21} = t_{p2} - t_{p1} = (t_0 + tt_{p2}) - (t_0 + tt_{p1}) = tt_{p2} - tt_{p1}$. *Zhou* (1994) used these travel-time differences over a network of stations to find a hypocenter location and coined it the equal differential time (EDT) method. The same data attribute was used by *Spetzler and Dost* (2017) for location of events in the Groningen region. To put the EDT method in a Bayesian framework in which multiple data attributes can be combined, in this report we refer to it as location using the P-delay data attribute.

In the following we illustrate the terms in Eq. 3.1 with the P-delay data attribute. Using a minimum of three receivers to find a PDF in \mathbf{x} , the data attribute vector is composed as

$$\mathbf{d} = \begin{bmatrix} t_{p2} - t_{p1} \\ t_{p3} - t_{p1} \end{bmatrix}, \quad (3.2)$$

Containing only the travel-time differences of the P-wave arriving at the second and first receiver (top row) and third and first receiver (bottom row).

For location, we use the travel-time models as derived in Section 2.6. It is assumed that the arrival times can be modeled using a Gaussian probability density distribution with a mean travel time and standard deviation σ_p . For three P-wave detections we have three measurements including noise, yielding the following noise model:

$$\mathbf{N} = \begin{bmatrix} t_{p1} \\ t_{p2} \\ t_{p3} \end{bmatrix} \quad (3.3)$$

and corresponding noise covariance matrix

$$\mathbf{C}_n = \begin{bmatrix} \sigma_p^2 & 0 & 0 \\ 0 & \sigma_p^2 & 0 \\ 0 & 0 & \sigma_p^2 \end{bmatrix}. \quad (3.4)$$

The matrix \mathbf{A} that relates the data-attribute vector to the noise vector ($\mathbf{d}=\mathbf{AN}$) is

$$\mathbf{A} = \begin{bmatrix} -1 & 1 & 0 \\ -1 & 0 & 1 \end{bmatrix}. \quad (3.5)$$

Using that $\mathbf{C}_d = \mathbf{AC}_n\mathbf{A}^T$, the data covariance matrix reads

$$\mathbf{C}_d = \sigma_p^2 \begin{bmatrix} 2 & 1 \\ 1 & 2 \end{bmatrix}. \quad (3.6)$$

In the data-attribute vector (Eq. 3.2) only $n - 1$ linearly independent receiver combinations are taken. Instead, one could choose to use all $n(n - 1)/2$ unique receiver combinations. This results in larger matrices, and thus longer computation times. Moreover, in this case the inverse of \mathbf{C}_d does not exist; different rows of \mathbf{d} can be written as a linear combination of other rows. And hence one would need to resort to a pseudo inverse. Given the used noise model, using $n(n - 1)/2$ unique station combinations yields the same hypocenter PDF as with the $n - 1$ linearly independent station combinations. So, there is no advantage in extending the data-attribute vector with linearly dependent receivers combinations and thus we use the $n - 1$ implementation as outlined above.

3.1.2 P-S delay

An equidistant curve can also be constructed from the arrival-time difference of the first P- and S-wave arrival at a single receiver. With P-S delay measurements at three stations, a location can be found in 3D which would be —if there was no noise— the intersection of the three equidistant curves. We use the minimum amount of three receivers to illustrate the terms in Eq. 3.1 when the P-S delay data attribute is used. The data-attribute vector looks as follows

$$\mathbf{d} = \begin{bmatrix} t_{s1} - t_{p1} \\ t_{s2} - t_{p2} \\ t_{s3} - t_{p3} \end{bmatrix}, \quad (3.7)$$

where t_s and t_p are S-wave and P-wave arrival times and the subscript number refers to the receiver number.

We use the travel-time models as derived in Section 2.6. It is assumed that the arrival times can be modeled using Gaussian probability density functions with mean arrival times and standard deviation thereof, σ_p for P-waves and σ_s for S-waves. For three P-wave detections and three S-wave detections we have six measurements including noise, yielding the following noise model:

$$\mathbf{N} = \begin{bmatrix} t_{s1} \\ t_{s2} \\ t_{s3} \\ t_{p1} \\ t_{p2} \\ t_{p3} \end{bmatrix} \quad (3.8)$$

and corresponding noise covariance matrix

$$\mathbf{C}_n = \begin{bmatrix} \sigma_s^2 & 0 & 0 & 0 & 0 & 0 \\ 0 & \sigma_s^2 & 0 & 0 & 0 & 0 \\ 0 & 0 & \sigma_s^2 & 0 & 0 & 0 \\ 0 & 0 & 0 & \sigma_p^2 & 0 & 0 \\ 0 & 0 & 0 & 0 & \sigma_p^2 & 0 \\ 0 & 0 & 0 & 0 & 0 & \sigma_p^2 \end{bmatrix}. \quad (3.9)$$

The matrix \mathbf{A} that relates the data-attribute vector to the noise vector ($\mathbf{d}=\mathbf{AN}$) is

$$\mathbf{A} = \begin{bmatrix} 1 & 0 & 0 & -1 & 0 & 0 \\ 0 & 1 & 0 & 0 & -1 & 0 \\ 0 & 0 & 1 & 0 & 0 & -1 \end{bmatrix}. \quad (3.10)$$

Using that $\mathbf{C}_d = \mathbf{A}\mathbf{C}_n\mathbf{A}^T$, the data covariance matrix reads

$$\mathbf{C}_d = \begin{bmatrix} \sigma_s^2 + \sigma_p^2 & 0 & 0 \\ 0 & \sigma_s^2 + \sigma_p^2 & 0 \\ 0 & 0 & \sigma_s^2 + \sigma_p^2 \end{bmatrix}. \quad (3.11)$$

3.1.3 Joint inversion

When errors may be assumed to be normally distributed, the use of more measurements reduces the location uncertainty. Likewise, also the usage of multiple data attributes per station, or station-combination, may be advantageous. Many different combinations are possible. Below we work out the combination of the P-S delay and P-delay data attribute. These two data attributes yield location PDFs with very different shapes. That is, both data attributes have complimentary sensitivity to the location problem. The joint inversion thus results in a significantly smaller PDF, as will be shown in Section 3.3.

We use again the minimum amount of three receivers to illustrate the terms in Eq. 3.1. When using both P-S delays and P delays, we have the following data-attribute vector:

$$\mathbf{d} = \begin{bmatrix} t_{s1} - t_{p1} \\ t_{s2} - t_{p2} \\ t_{s3} - t_{p3} \\ t_{p2} - t_{p1} \\ t_{p3} - t_{p1} \end{bmatrix}. \quad (3.12)$$

In the above vector, the same measurements are used as in the previous section. Hence, the same noise model (Eq. 3.8) and noise covariance matrix (Eq. 3.9) remain valid.

The matrix \mathbf{A} that relates the data-attribute vector to the noise vector ($\mathbf{d}=\mathbf{A}\mathbf{n}$) is

$$\mathbf{A} = \begin{bmatrix} 1 & 0 & 0 & -1 & 0 & 0 \\ 0 & 1 & 0 & 0 & -1 & 0 \\ 0 & 0 & 1 & 0 & 0 & -1 \\ 0 & 0 & 0 & -1 & 1 & 0 \\ 0 & 0 & 0 & -1 & 0 & 1 \end{bmatrix}. \quad (3.13)$$

Using that $\mathbf{C}_d = \mathbf{A}\mathbf{C}_n\mathbf{A}^T$, the data covariance matrix reads

$$\mathbf{C}_d = \begin{bmatrix} \sigma_s^2 + \sigma_p^2 & 0 & 0 & \sigma_p^2 & \sigma_p^2 \\ 0 & \sigma_s^2 + \sigma_p^2 & 0 & -\sigma_p^2 & 0 \\ 0 & 0 & \sigma_s^2 + \sigma_p^2 & 0 & -\sigma_p^2 \\ \sigma_p^2 & -\sigma_p^2 & 0 & \sigma_p^2 + \sigma_p^2 & \sigma_p^2 \\ \sigma_p^2 & 0 & -\sigma_p^2 & \sigma_p^2 & \sigma_p^2 + \sigma_p^2 \end{bmatrix}. \quad (3.14)$$

3.2 Uncertainty quantification

For finding a source location, Eq. 3.1 is implemented. The resulting solution is not just the most likely source location, but the entire location PDF ($\sigma_M(\mathbf{x})$) which gives the probability of the source as function of a range of locations in 3D space. This PDF is —by approximation— a multivariate normal distribution which can be described with its mean $\boldsymbol{\mu}=[\mu_x \mu_y \mu_z]$ and standard deviation $\boldsymbol{\sigma}=[\sigma'_1 \sigma'_2 \sigma'_3]$, which are the square root of the variances in the three principal directions. The mean of a normal distribution corresponds to the location in space with the highest probability density (the mode). This most likely location

is communicated as the source location, and the uncertainty is communicated through the standard deviations of the location PDF.

Fig. 3.1(a) shows an example of a 95% confidence ellipsoid. There is a probability of 95% that the source location lies within this volume. Generally, the principal directions of the uncertainty ellipsoid (x'_1, x'_2, x'_3) do not correspond to the directions of the grid (X, Y, Z). Hence, nine parameters are needed to describe the ellipsoid: three to position the center point of the ellipsoid in 3D space, three rotation angles to map from (X, Y, Z) to (x'_1, x'_2, x'_3) and three sigmas to denote the size of the ellipsoid. The ellipsoid can be constructed by

1. Taking the mode of the PDF as the center point of the ellipsoid.
2. Outlining the directions of the principal axes.
3. Drawing the ellipsoid with the correct size of the principal axes.

This size is found by multiplying 2 times 2.7955 to σ'_1 , σ'_2 and σ'_3 . That is, for a tri-variate normal distribution, 95% of the probability lies within 2.7955 standard deviations from the mean (*Wang et al.*, 2015). Conversely, e.g. σ'_1 can be obtained by dividing the distance span by the 95% confidence in the x'_1 direction by 5.5910.

With the above parameterisation, three parameters are used to denote the most likely location and 6 to describe the uncertainty. In seismology, however, we are especially interested in epicentral uncertainty (uncertainty in the horizontal plane) and depth uncertainty (uncertainty in depth direction). Instead of parameterising the uncertainty ellipsoid, we quantify the epicentre uncertainty and depth uncertainty separately. This yields a simplified description of the uncertainty with only four parameters: σ_Z (Fig. 3.1b), σ_1 , σ_2 and θ (Fig. 3.1c).

The depth uncertainty is found by —at the epicentre— computing the location probability in the vertical direction. This gives a 1D PDF as depicted in Fig. 3.1(b), which is a univariate normal distribution. Subsequently, the depth range is computed that contains 95% of the location probability, Z^{95} . In 1D, $Z^{95} = 2 \cdot 1.96\sigma_Z$ (e.g., *Wang et al.*, 2015), wherein σ_Z is the standard deviation in the vertical direction. The above formula is often simplified by stating the 95% of the probability is reached within two sigmas from the mean. We use the more exact

$$\sigma_Z = \frac{Z^{95}}{2 \cdot 1.96} = \frac{Z^{95}}{3.92}. \quad (3.15)$$

The epicentral uncertainty is found by computing the 95% confidence region in a X-Y (or East-North) plane through the assumed earthquake depth. This gives a 2D PDF as depicted in Fig. 3.1(c), which is a bi-variate normal distribution. Subsequently, the ellipse is computed that contains 95% of the location probability:

$$\frac{x_1^2}{\sigma_1^2} + \frac{x_2^2}{\sigma_2^2} = \kappa^2, \quad (3.16)$$

which ellipse has principal directions x_1 and x_2 . θ is the clockwise rotation angle of x_1 with respect to north (Fig. 3.1c). When the ellipse describes the bounds of the 95% confidence region computed in 2D, $k = 2.4477$ (*Wang et al.*, 2015). From this ellipse, the major axis and minor axis are X_1^{95} and X_2^{95} , respectively. Using the above formula, it follows that the uncertainty in the two principal directions in the X-Y plane can be found with

$$\sigma_{1,2} = \frac{X_{1,2}^{95}}{2 \cdot 2.4477} = \frac{X_{1,2}^{95}}{4.8954}. \quad (3.17)$$

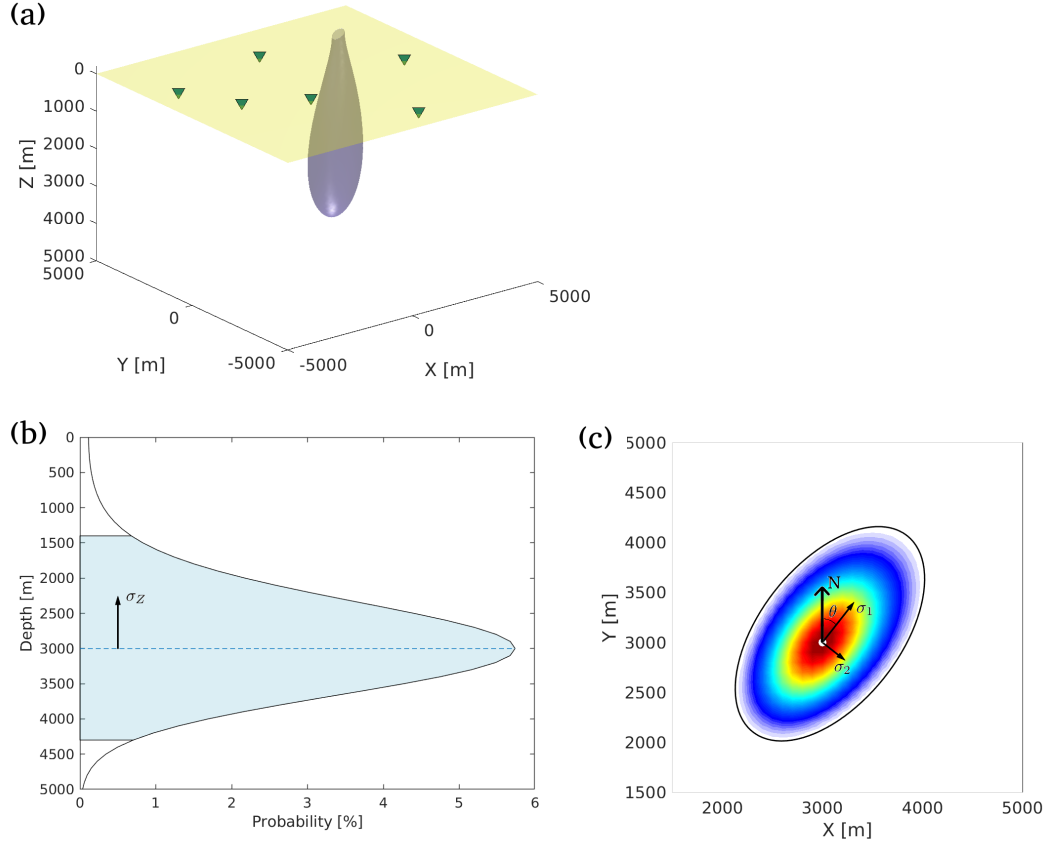


Figure 3.1: Location-uncertainty assessment in different dimensions. (a) 95% confidence ellipsoid (grey shaded) in 3D shown together with the acquisition geometry of receivers (green triangles) on the Earth's surface (yellow). (b) Event-depth PDF, of which the 95% confidence region is shaded blue. The arrow depicts the size of σ_z (Eq. 3.15). The PDF has been computed at the epicenter. (c) Event-epicenter PDF of which the 95% confidence region is bounded by the black ellipse. The PDF has been computed at the event depth. σ_1 and σ_2 are the largest and smallest uncertainties in the epicentral plane and the angle θ describes the directivity of σ_1 with respect to north. σ_2 is orthogonal to σ_1 .

The following steps are taken to quantify the epicentral uncertainty in terms of σ_1 , σ_2 and θ :

- Compute a PDF in the 2D plane through the (assumed) depth of the source.
- Normalize the total probability to 100%.
- Sort the probabilities from high to low.
- Find the grid point, and corresponding probability, at which 95% of the probability is reached (P95).
- Interpolate a contour through 2D space where the P95 value is reached.
- Fit the contour with an ellipse.
- Determine the major (X_1^{95}) and minor axes (X_2^{95}) of the ellipse and its rotation angle with north θ .
- Use Eq. 3.17 to compute σ_1 and σ_2 .

In *Wang et al. (2015)* the ratio between the size of the confidence region and σ are called magnification ratio's. In 2D, the magnification ratio corresponds to κ in Eq. 3.16. Note that these ratios are not fixed, but depend on the space dimensionality in which the PDF is computed. The higher the dimension, the larger the confidence zone and thus the larger the magnification ratio. This can be seen, e.g., when comparing Fig 3.1(a) and (b). In 3D the depth extend of the 95% confidence region extends to the free surface, whereas in 1D it ends about 1400 m below the surface.

At Fig. 3.1 it can be seen that the PDF is a (multivariate) normal distribution only by approximation. For instance, note the (small) asymmetry of the PDF in the depth extend. The confidence zone is larger at the shallow end of the most likely source location than at the deep end. Due to the presence of the free surface, the mean does not correspond to the mode. The mode still depicts the actual source depth, whereas the mean would give an underestimate of the source depth. Therefore, the mode is used to communicate the source location.

3.3 Comparison of data attributes

In this section we illustrate the location uncertainty that is obtained using the different data attributes as introduced in Section 3.1. For this illustration, we use a local network with six receivers (Fig. 3.2 left column). The source is placed at 3 km depth below the network. There is a good azimuthal coverage. We use the low travel-time uncertainties of $\sigma_p = 0.0893$ and $\sigma_s = 0.170$ s as estimated in Section 2.6.2. Fig. 3.2 shows the resulting PDFs.

The top row in Fig. 3.2 shows the PDFs (in 3D, epicenter and depth) when using P delays. This data attribute gives a quite confined epicentral 95% confidence zone. However, it yields a large uncertainty in depth. This is because shifting the source location in X-Y results in (significant) time shifts of the incoming P wave at the receivers. Changing the depth of the source, however, leads to much smaller differences in the P-wave front that is sampled with the local network.

The middle row in Fig. 3.2 shows the PDFs (in 3D, epicenter and depth) when using P-S delays. This data attribute gives a less confined epicentral 95% region than when using P delays. However, it yields a much better confined depth uncertainty. The travel-time differences between incoming P- and S-waves contain especially information of the distance

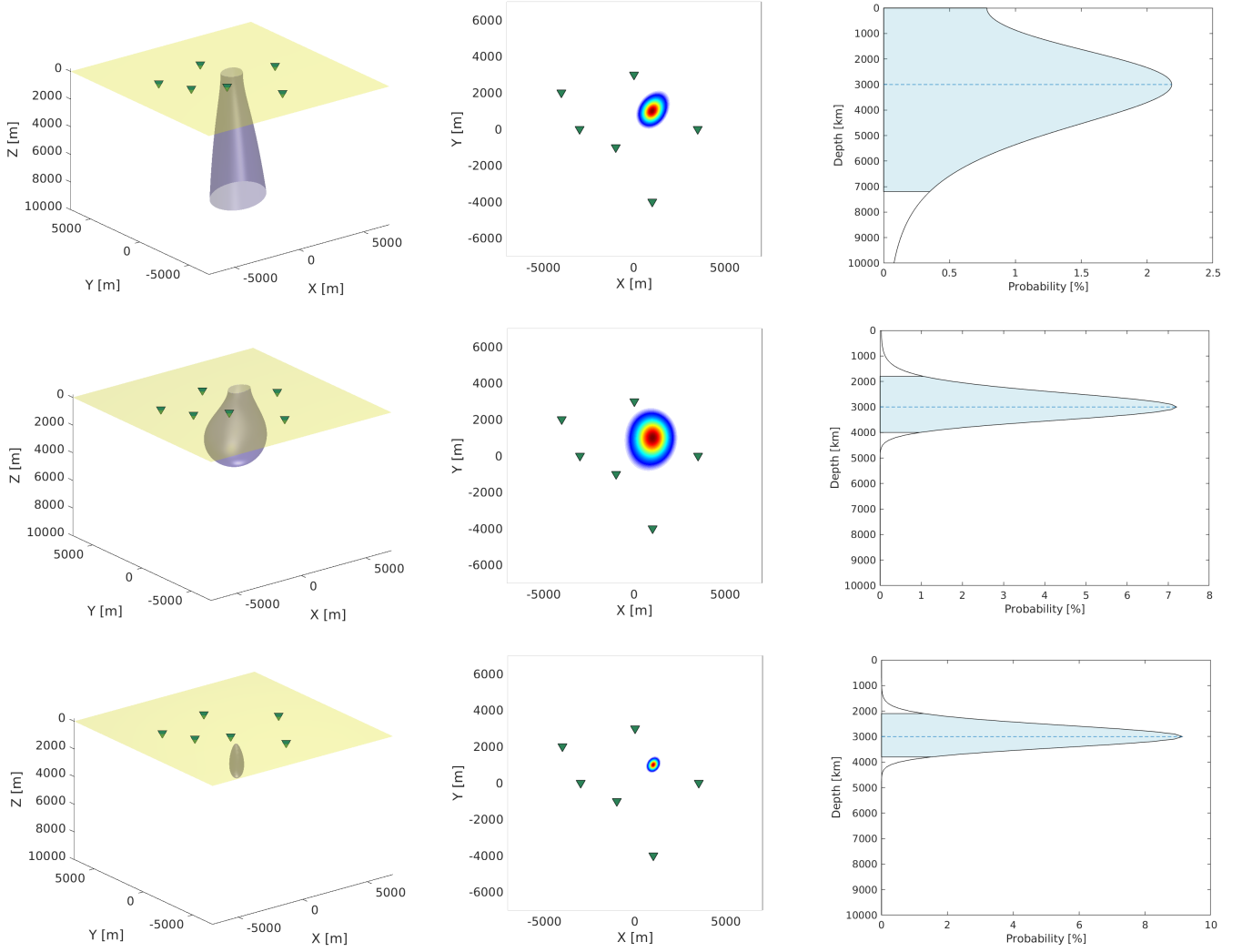


Figure 3.2: 95% location confidence regions for a setup with six receivers (green triangles) on the Earth's surface (yellow surface) in 3D (left column), in the epicentral plane (middle column) and in the depth direction (right column). The top row shows results when using the P-delay data attribute, yielding $(\sigma_1, \sigma_2, \sigma_Z) = (489, 350, 1837)$ m and $\theta = 33.7^\circ$. The middle row shows results when using the P-S delay data attribute, yielding $(\sigma_1, \sigma_2, \sigma_Z) = (746, 622, 561)$ m and $\theta = 7.1^\circ$. The lower row shows results when doing a joint inversion with both data attributes, yielding $(\sigma_1, \sigma_2, \sigma_Z) = (205, 153, 434)$ m and $\theta = 31.4^\circ$.

of the source with respect to the network. When a high depth resolution is desired, it is elementary to include this data attribute in an inversion.

The bottom row in Fig. 3.2 shows the PDFs (in 3D, epicenter and depth) when using a joint inversion with both P- and P-S delays. The joint inversion leads to smaller PDFs than with any of the two data attributes separately. The near-orthogonal sensitivity of both data attributes (depth versus epicentral) are combined to obtain the best overall hypocentre estimate. Its estimate has the smallest values for σ_1 , σ_2 and σ_Z .

4 Sensitivity analysis

In this chapter we assess how location uncertainty varies as a function of different variables. The variables we consider are uncertainty of the underlying measurements (σ_p and σ_s), systematic changes in the network layout (azimuthal gap, number of receivers, array aperture and distance to the nearest station) and source depth. Understanding sensitivity of location uncertainty to the different variables paves the way of finding effective means of reducing the location uncertainty. Also, the sensitivity analyses help in understanding the details of the Netherlands-wide location-uncertainty maps that follow in the next chapter and in Appendices B and D.

In the previous chapter, specifically in Fig. 3.2, it was shown that a joint inversion with P-S and P delays is most effective in constraining the source location, both in the depth extend (σ_Z) and in the epicentral plane (σ_1 and σ_2). All numerical analyses in this chapter are only performed for this joint inversion with multiple data attributes.

The local velocity structure, and the resulting travel-time function, has an influence on how the travel-time errors are propagated to event-location uncertainty.

On the one hand, errors are relative to an average propagation velocity. When the propagation velocity is 6 km/s instead of 3 km/s, a 0.1 s error leads to twice the location error. Hence, it is important to take realistic background velocities. In this section, the error propagation is done through a homogeneous background medium with P-wave and S-wave velocities of $v_P=4.9$ and $v_S=2.9$ km/s. These are typical apparent velocities as found for upper-crustal events in the Netherlands (Section 2.6.1 and *Jagt et al.* (2017)).

On the other hand, complex details in the velocity structure lead to perturbations of the location-uncertainty PDF. This last aspect is ignored in this report. The perturbations in the location PDF can only be computed well if detailed local velocity models are available. Moreover, these perturbations will be different in different regions, which make them hard to include in a NL-wide framework. When possible complexities of the location PDF are of interest, (additional) local or regional studies are needed. The modelling outcomes in the coming sections serve to show the general trend of how location uncertainty behaves as function of different variables.

4.1 Travel-time uncertainty

In Section 2.6 we estimated the uncertainty of the measurements that are used for location: P-wave and S-wave travel times. In specific cases, the travel-time uncertainty may in fact be lower. For instance, when travel times are forward modeled through a detailed local 3D velocity model. In other cases, the travel-time uncertainty may be higher. For instance, when the source is on a fault with a large offset, with very different velocity structure on either side of the fault. Here we assess how the location uncertainty changes when the travel-time uncertainty is varied.

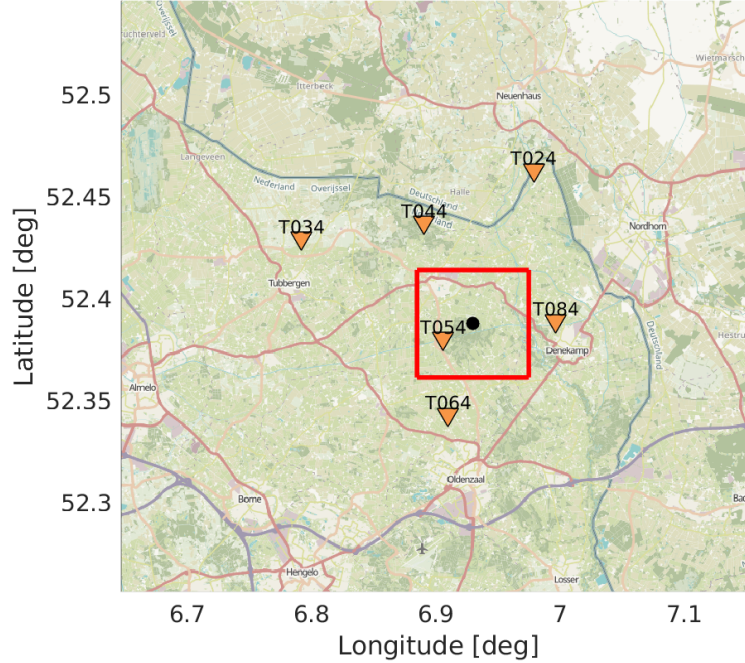


Figure 4.1: Configuration of stations (orange triangles) that likely detects a $M=0.5$ source placed at 3km depth (black dot). Within the red box, a grid search is done to find the epicentral PDF. Background map is from www.openstreetmap.org.

We use a source with $M=0.5$ at 3km depth within the Twente seismic network (Fig. 4.1) and forward model travel times through a homogeneous background medium. We use the tools developed in Chapter 2 to identify the stations where this scenario earthquake is likely detected: we use the P-wave GMPES to model the amplitude at each station of the detection list and compare the modeled amplitude with the P90 local noise conditions and account for the picking threshold and possibly a hard-rock correction factor. For this $M=0.5$ event this yields a pick list that contains the six nearest stations of the Twente array. From these stations we use the P-wave and S-wave picks.

We first assess the situation in which σ_s is kept fixed at 0.248s and σ_p is varied. The resulting location uncertainties (σ_1 , σ_2 and σ_Z) are shown in Fig. 4.2. For small σ_p , the model uncertainties depend nearly linearly on σ_p . For larger σ_p , the epicentral uncertainties are reaching a plateau. The plateau starts at $\sigma_p = \sigma_s = 0.0248$ s. Beyond that point, the epicentral location is predominantly constrained by S-wave travel-time information, and since the uncertainty thereof is fixed, the epicentral uncertainty nearly stops increasing.

In Fig. 4.2 it is also shown how the epicentral (lower panels) and depth (higher panels) PDFs look like for $\sigma_p = 0.12$ and $\sigma_p = 0.36$ s. The epicentral PDFs do not notably change shape when σ_p increases. The depth PDF shows different behavior: for small σ_p there is limited interaction of the PDF with the free surface. For larger σ_p , the PDF is bounded on the upper end by the free surface, which limits its growth. Thus, the σ_Z function (yellow line) has a higher derivative ($d\sigma_Z/d\sigma_p$) for $\sigma_p < 0.18$ than for $\sigma_p > 0.18$.

We have derived in Section 2.6 that σ_s is about twice as large as σ_p for induced seismicity in the Netherlands. We assess how the location uncertainty changes as function of simultane-

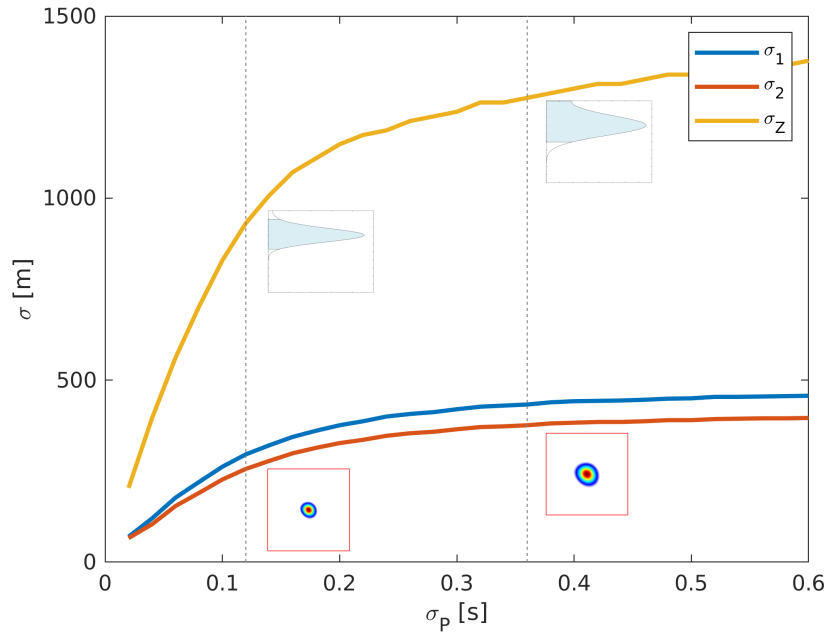


Figure 4.2: Location uncertainty as a function of P-wave travel-time uncertainty σ_p . The model uncertainty in the epicentral plane is expressed by σ_1 and σ_2 . The uncertainty in the depth direction is given by σ_z . The location uncertainty is computed using the station configuration in Fig 4.1 and a source at 3 km depth. The S-wave travel-time uncertainty is fixed at 0.248 s. The insets show the shape of the epicentral and depth PDFs for $\sigma_p = 0.12$ s (left-hand panels) and for $\sigma_p = 0.36$ s (right-hand panels).

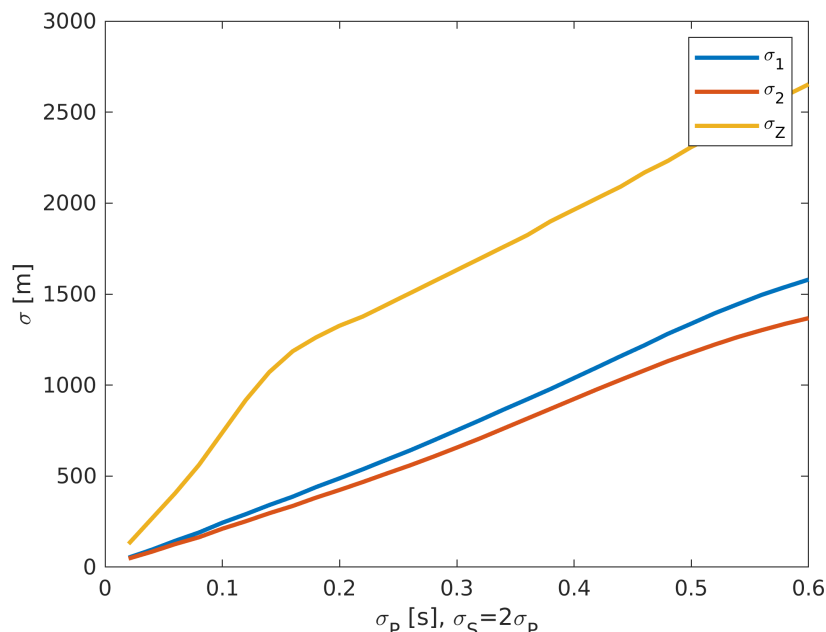


Figure 4.3: Location uncertainty (σ_1 , σ_2 and σ_Z) as function of travel-time uncertainty (σ_p and σ_s). σ_1 and σ_2 are varied at the same time, with σ_p annotated on the horizontal axis and $\sigma_s = 2\sigma_p$. The location uncertainty is computed using the station configuration in Fig 4.1 and a source at 3 km depth.

ous changes in both σ_p and σ_s , when $\sigma_s = 2\sigma_p$. Fig. 4.3 shows the resulting curves. In this case, the epicentral uncertainty (orange and blue lines) varies approximately linearly with travel-time uncertainty. The depth uncertainty (yellow line) is significantly larger than the epicentral uncertainty. σ_Z shows a similar character as in Fig. 4.2, with a large $d\sigma_Z/d\sigma_p$ derivative at small σ_p and a smaller derivative as soon as the PDF becomes constrained by the free surface.

4.2 Source depth

In this section we numerically compute the influence of source nucleation depth on location uncertainty. We use the same six stations as in the previous section (Fig. 4.1), with the same epicentral source location and vary the source depth. The smallest (epicentral) source-receiver distance within the configuration is 1.82 km and the largest is 10.48 km. We use low travel-time uncertainties of $\sigma_p = 0.0893$ and $\sigma_s = 0.170$ s as estimated in Section 2.6.2. Fig. 4.4 shows the variation of σ_1 , σ_2 and σ_Z as function of source depth.

Fig. 4.4 shows that epicentral uncertainty (σ_1 , σ_2) increases nearly linearly with source depth. As we have seen in Section 3.3, the epicentral uncertainty is predominantly constrained by P-delays over the network. The deeper the source, the smaller the P-delay times (that is, the more the wavefront approximates a plane wave over the network) and hence the larger the epicentral uncertainty.

The depth uncertainty σ_Z shows a different character (yellow line in Fig. 4.4). As function of depth, it peaks at shallow depths and becomes stable at larger depths. When the source is shallow, the PDF is bounded by the free surface. This results in a reduction of the PDF

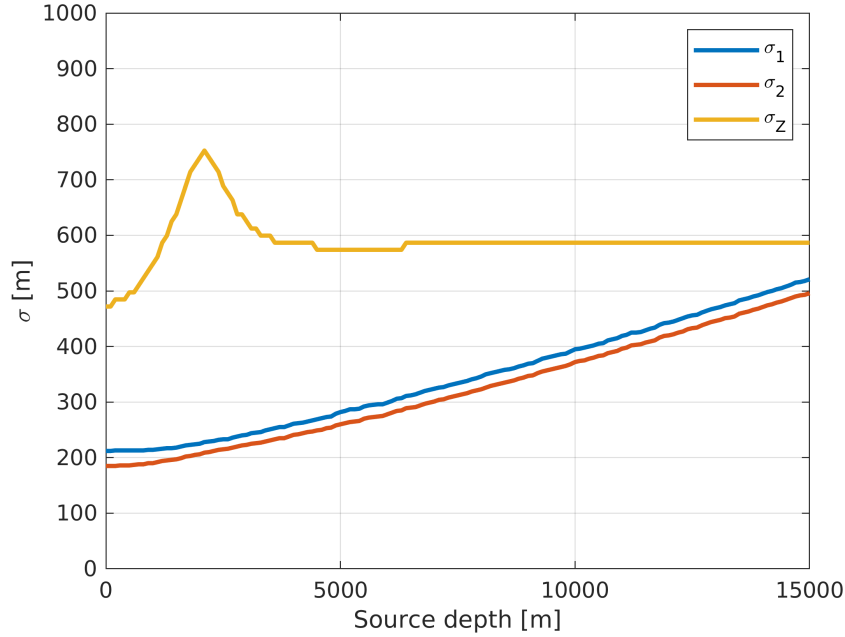


Figure 4.4: Location uncertainty (σ_1 , σ_2 and σ_Z) as function of source depth for the source-receiver configuration as shown on Fig. 4.1.

and its 95% confidence zone when the source is shallower than 1200 m. It results in an extended 95% confidence zone when the source depth is between 1200 and 3600 m. At larger depths, there is no interaction of the PDF with the free surface and σ_Z remains stable at approximately 590 m.

Generally, velocities increase with depth, which results in an increase of the location error with depth (or likewise a decrease of the error at shallow depth). This additional effect is not reflected in Fig. 4.4) because propagation velocities were kept stable.

4.3 Source-receiver distance

In this section we numerically illustrate the influence of source-receiver distance on location uncertainty. We use a setup of five receivers that surround a source at 3 km depth. We use the travel-time uncertainties of $\sigma_p = 0.0893$ and $\sigma_s = 0.170$ s as estimated in Section 2.6.2. We consider two different configurations.

Fig. 4.5(a) shows the setup for the first experiment. All five receivers are placed with a source-receiver distance of 0.4 km and the location uncertainty is computed. Successively, the source-receiver distances are increased to 20 km, with steps of 0.4 km, and for each configuration the location uncertainty is computed. Fig. 4.6 shows the resulting variation of location uncertainty as function of array aperture.

When the aperture becomes infinitely small, it would be impossible to estimate a location based on travel-time differences. In this case, the PDF would be a half-sphere with many epicenter-depth combinations with equal likelihood. Hence, the numerical illustration starts for source-receiver distances of 0.4 km.

Fig. 4.6 shows that σ_1 and σ_2 decrease from 0.4 until 1.2 km distance and remain stable

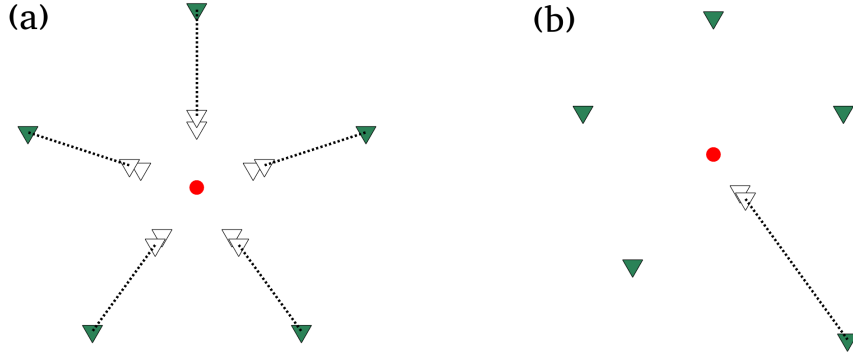


Figure 4.5: Configurations for a numerical experiments in which (a) the aperture of a local network is varied and (b) only the source-receiver distance for a single receiver is varied. For each experiment, the same source location (red dot) is used, whereas positions of the five receivers (triangles) are varied in experiment (a) and only the position of one receiver is varied in experiment (b).

for larger distances. The uncertainties are relatively high when the local network has an aperture that is considerably smaller than the source depth. From 1.2 km distance, the epicentral location precision is not affected by an increase in aperture. The epicentral uncertainty is primarily dictated by the azimuthal coverage of receivers. Since this coverage remains identical, also the uncertainty remains identical.

In practise, the epicentral uncertainty would grow with larger aperture as the travel-time uncertainty does grow with epicentral distance (Section 2.6.1). This effect is not included in this section. Here we merely model the geometric influence of the source-receiver configuration.

The depth uncertainty σ_Z behaves differently. For small array aperture it approximates 600 m. For larger array aperture it increases first with a large derivative, and for distances beyond ~ 3.5 km with a smaller derivative. As with the previous experiments, the trajectory with smaller growth in σ_Z with aperture occurs when the PDF is bounded on the upper side by the free surface.

Fig. 4.5(b) shows the setup for the second experiment. Four receivers are placed with a source-receiver distance of 10 km. A fifth receiver is placed at a distance of only 0.4 km and the location precision is computed. Successively, the source-receiver distances is increased to 20 km, with steps of 0.4 km, whereas the other four receivers remain at 10 km distance. Fig. 4.7 shows the resulting variation of location uncertainty.

Fig. 4.7 shows that σ_1 and σ_2 are approximately stable as function of source-receiver distance. As with the previous experiment, the azimuthal coverage remains identical and hence there is no variation in epicentral uncertainty. Since four of the receivers are at 10 km distance, the case with a network aperture smaller than the source depth, does not occur.

The depth uncertainty σ_Z in Fig. 4.7 behaves similarly as σ_Z in Fig. 4.6 for the first 10 km source-receiver distance. This means that placing only a single receiver close to the source zone is nearly as effective as placing all receivers near the source, in reducing the depth uncertainty.

The depth uncertainty on Fig. 4.7 reaches a maximum when all receivers have a distance of 10 km from the source. The depth uncertainty decreases again when the single receiver

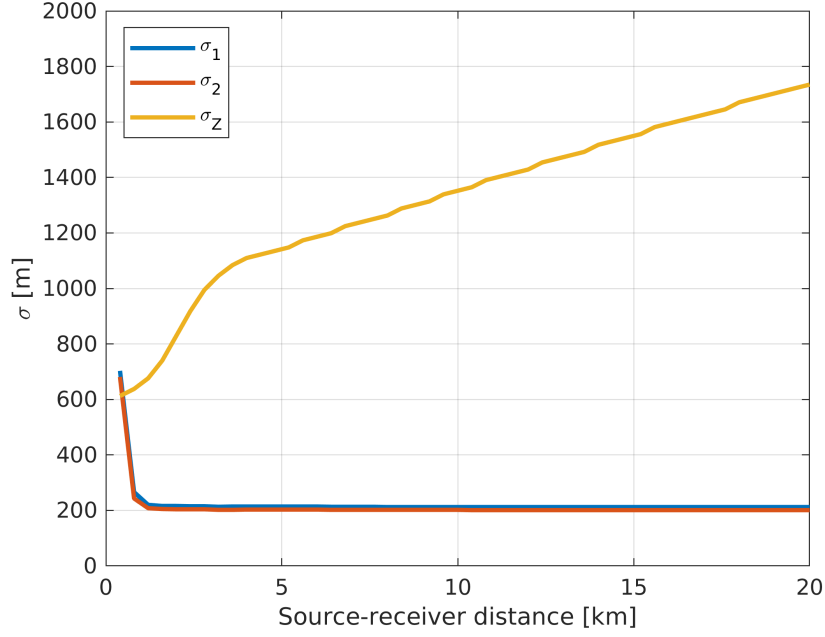


Figure 4.6: Location uncertainty (σ_1 , σ_2 and σ_z) as function of network aperture. The corresponding source-receiver configuration is shown in Fig. 4.5(a) for a source at 3 km depth.

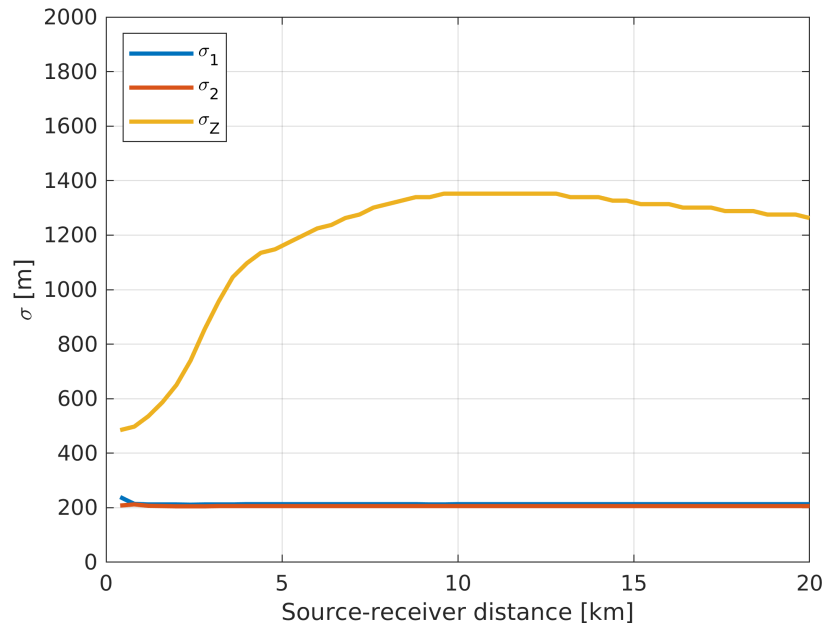


Figure 4.7: Location uncertainty (σ_1 , σ_2 and σ_z) as function of source-receiver distance variation of one (out of five) receivers. The other four receivers are fixed at a source-receiver distance of 10 km. The corresponding source-receiver configuration is shown in Fig. 4.5(b) for a source at 3 km depth.

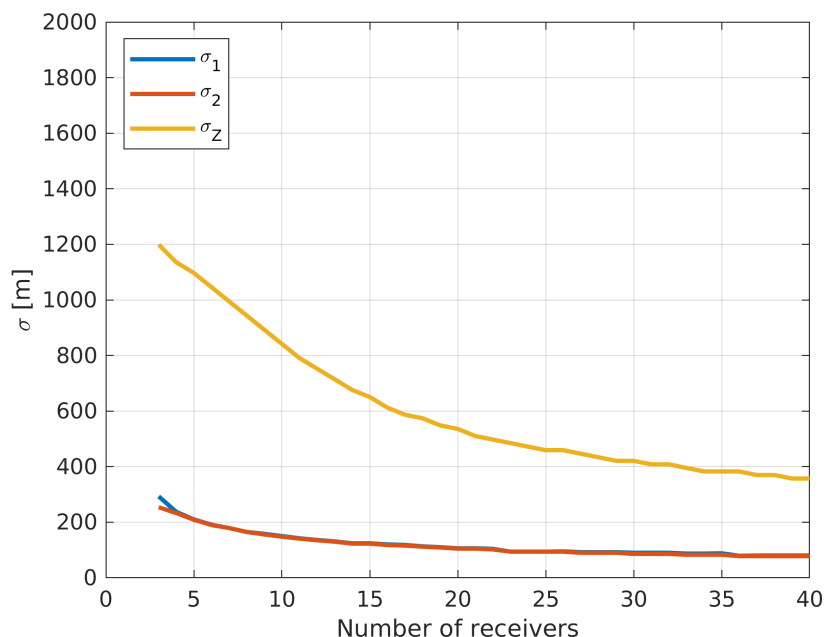


Figure 4.8: Location uncertainty (σ_1 , σ_2 and σ_Z) as function of number of receivers surrounding a 3 km deep source at 5 km epicentral distance.

is placed at larger distances. This shows that it is favourable to have receivers at varying distances from the source, to better sample the P-S-delay-versus-distance function and therewith to have a better estimate of source depth.

With a point-symmetric configuration, the 95% epicentral confidence zone is a circle, meaning that the ellipse minor axis is equal to the major axis, or $\sigma_1 = \sigma_2$. In order to show non-overlapping lines for σ_1 and σ_2 we have slightly perturbed the configurations in Fig. 4.5, resulting in σ_1 and σ_2 functions that can just be distinguished from each other in Figs. 4.6&4.7.

4.4 Number of receivers

In this section we numerically illustrate the influence of the amount of receivers on location uncertainty. We use the low travel-time uncertainties of $\sigma_p = 0.0893$ and $\sigma_s = 0.170$ s as estimated in Section 2.6.2. We use a setup with a source at the origin, at 3 km depth, which is surrounded with receivers at a distance of 5 km. The receivers are placed at equal azimuths. The minimum amount of receivers is three, in which case the azimuthal difference (with respect to the source) between any two receivers is 120° . For the next case with four receivers, the azimuthal difference is 90° , etc. The location uncertainty is continued until the case with 40 surrounding receivers. The results are shown in Fig. 4.8.

Fig. 4.8 shows that both the epicentral (σ_1 , σ_2) and depth uncertainty σ_Z keep reducing with an increasing amount of receivers. However, the returns are diminishing as the reduction in uncertainty becomes smaller at larger receivers numbers. In this example, there is a limited reduction in epicentral uncertainty beyond ~ 10 receivers and a limited reduction in depth uncertainty beyond ~ 20 receivers.

For this experiment $\sigma_1 = \sigma_2$, because point-symmetric configurations of receivers are chosen.

In reality, receiver configurations would be more chaotic, resulting in larger azimuthal gaps (next section) and, as a consequence, $\sigma_1 > \sigma_2$. The uncertainties in Fig. 4.8 are a lower bound of what can be achieved.

4.5 Azimuthal gap

In this section we numerically illustrate the influence of azimuthal gap on location uncertainty. The azimuthal gap is defined as the largest angle between two stations as seen from the epicenter (*Scudero et al.*, 2021). Thus, it is the largest azimuthal difference between any two stations. We use the travel-time uncertainties of $\sigma_p = 0.0893$ and $\sigma_s = 0.170$ s as estimated in Section 2.6.2. We use a setup with a source at the origin, at 3 km depth and receivers at 5 km distance.

Fig. 4.9 shows the receiver configurations and modelling results for three different azimuthal gaps. Fig. 4.10 shows the location uncertainties as function of all considered azimuthal gaps. The epicentral uncertainties (σ_1, σ_2) increase with azimuthal gap. The depth uncertainty σ_Z remains stable as function of azimuthal gap. σ_Z is primarily a function of the nearest source-receiver distance (Fig. 4.7). Because this distance remains stable, also the uncertainty remains stable.

Fig. 4.9 shows in the left column the 3D location uncertainty 95% confidence regions. For the upper two configurations (azimuthal gaps of 140 and 220°, respectively) the 3D PDF can still be approximated as multivariate normal distributions. That is, the 95% confidence regions are still by approximation ellipsoids which size can be approximated with σ_1, σ_2 and σ_Z . For the lower configuration (azimuthal gap of 300°) the 95% confidence region does not look like an ellipsoid. It is a twisted volume in 3D which cannot be easily parameterized. With these complicated PDFs, location may be ambiguous. By inspection of the PDFs for all azimuthal gaps, we find that complicated PDFs start to occur from an azimuthal gap of about 250°. This corresponds to the point Fig. 4.10 where σ_1 ceases to grow further. We mark this as the azimuthal gap beyond which the assumptions made with the PDF parameterization are increasingly invalid, and location may be problematic.

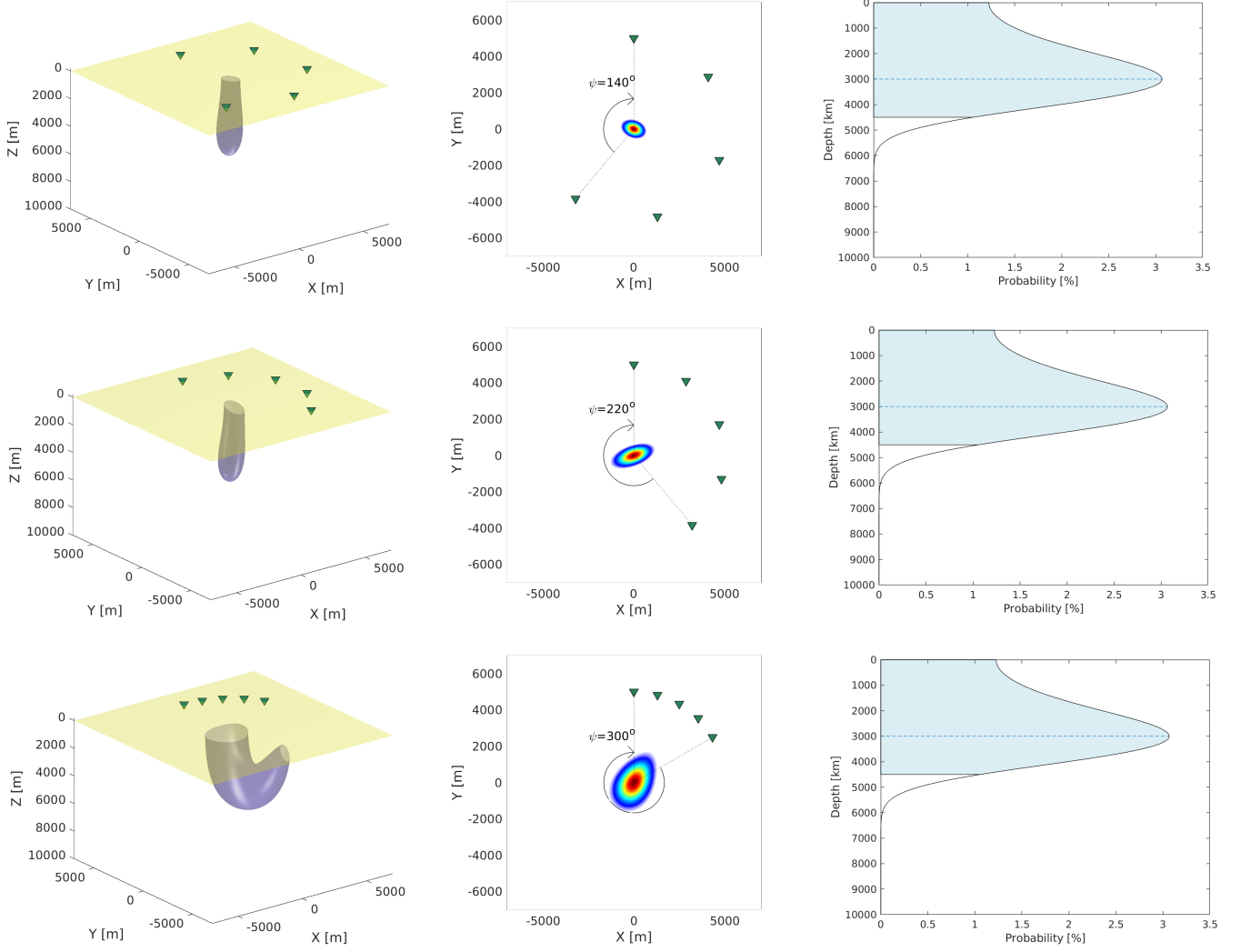


Figure 4.9: 95% location confidence regions for a setup with 5 receivers (green triangles) on the Earth's surface (yellow surface) in 3D (left column), in the epicentral plane (middle column) and in the depth direction (right column). The top row shows results when the receivers are placed such that there is an azimuthal gap ψ of 140° . The middle row shows results for an azimuthal gap of 220° . The lower row shows results for an azimuthal gap of 300° .

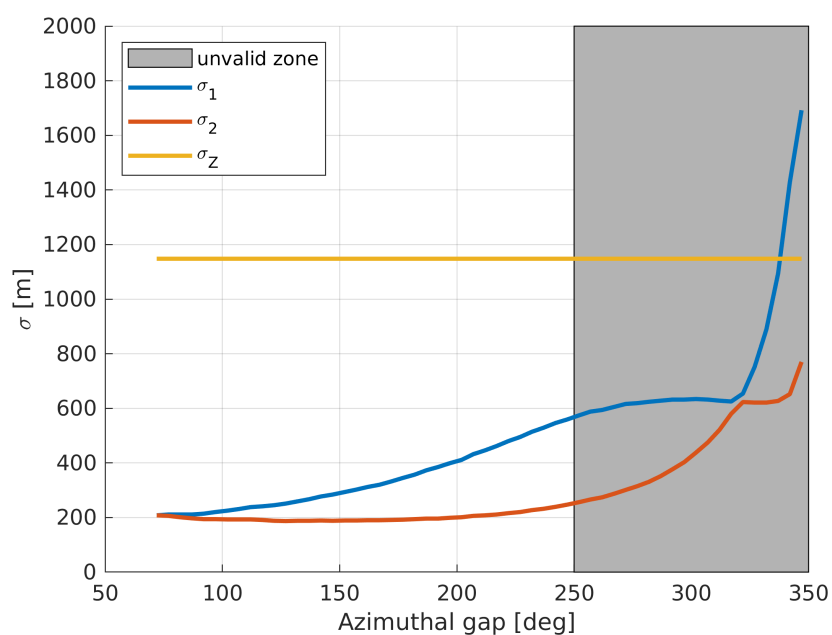


Figure 4.10: Location uncertainty (σ_1 , σ_2 and σ_Z) as function of azimuthal gap, for 5 receivers at 5 km epicentral distance from a 3 km deep source. The grey zone shows the area where the parameterization of the location PDF with σ_1 , σ_2 and σ_Z is increasingly invalid. In this zone, the found uncertainties are likely underestimates.

5 Location-uncertainty maps

In this chapter, the location uncertainty maps are presented. These maps have been computed for the Netherlands and border regions, using the station distribution as shown in Fig. 2.2.

The computation is done using a grid of (scenario) events with a grid-cell spacing of 1 km. A local rectangular grid (Rijksdriehoekstelsel, RD) is used, which has as origin a location close to Paris. This ensures that the whole of the Netherlands is in the first quadrant. The computation grid spans $RDx=[-20\ 300]$ and $RDy=[270\ 670]$ km. $(RDx, RDy)=(-20, 270)$ is the grid center of the grid cell furthest to the southwest. In wgs84, this point corresponds to $(lat, lon)=(50.3940, 2.9262)$. $(RDx, RDy)=(300, 670)$ is the grid center of the grid cell furthest to the northeast. In wgs84, this point corresponds to $(lat, lon)=(53.9951, 7.5980)$.

In this chapter, uncertainty maps are shown for parts of the Netherlands. In the appendix, one-page maps are included that show location uncertainty for the whole of the Netherlands. Even when these maps are printed on a single page, not all details are clearly visible. Therefore, also the underlying data with the computed values for each grid cell are distributed via *KNMI* (2023b).

The azimuthal gap ψ is often used as a proxy for epicentral uncertainty. In this research we explicitly compute epicentral uncertainty, so there is no need to show a proxy. Nevertheless, we also derive ψ maps in this chapter. The reason to include ψ maps is that we found that the used parametrization of location uncertainty ceases to hold for $\psi \geq 250^\circ$ (Section 4.5). The maps in Appendices B.3&D.3 serve to indicate where this may occur. Moreover, ψ could be used to assess the possibility to implement a well-constrained moment-tensor inversion, which is possible when multiple quadrants of the radiation pattern are sampled. At least two different quadrants are sampled when $\psi \leq 120^\circ$.

This chapter starts with a description of the detailed settings that have been chosen in the implementation. The province of North Holland has been used for quality checking. In the following two sections, the epicentral- and depth-uncertainty maps are presented. In the last section, the updated magnitude-of-completeness map is published.

Examples in this report have been made with a station distribution that was in operation in September 2021 (Table A.1). By September 2022 there have been several changes (Table C.1). These changes are described in the magnitude-of-completeness section, which contains both the 2021 and 2022 map. The 2021 and 2022 versions of MoC maps can be found at *KNMI* (2023a) whereas the location-uncertainty maps are published at *KNMI* (2023b) and can be found in the appendix. Yearly updates of the maps are foreseen. Overall, the station coverage has been improved from 2021 to 2022, resulting in reductions in uncertainty.

5.1 Settings and quality checks

5.1.1 Travel-time uncertainty

When a distance-independent uncertainty can be assumed, the more picks are included, the better (Fig. 4.8). There are diminishing returns, however. Beyond 20 picks there is hardly an improvement. We restrict the amount of picks that is used to the 40 nearest detections to limit computation time.

In Section 2.6 we found the travel-time uncertainty for different cases. When the azimuthal coverage is good ($\psi < 120^\circ$) and there are many picks nearby (at least 12 at distances smaller than 40 km) an event-specific travel-time model can be derived from these picks. This yields the smallest uncertainties as estimated in Section 2.6.2.

For all other cases, a background travel-time model is used, which is an average over different regions in the Netherlands. This leads to larger uncertainties as estimated in Section 2.6.1. We found that the travel-time uncertainty is distance dependent. The uncertainty was estimated for three different distance bins (local, regional and national) and listed in Table 2.4. The effective uncertainty is implemented as a weighted average of picks available in different distance bins. For P-waves we use

$$\sigma_p = (0.115n_1 + 0.162n_2 + 0.295n_3)/n_T, \quad (5.1)$$

where σ_p is the effective P-wave travel-time uncertainty for the collection of picks, n_1 , n_2 and n_3 are the number of picks that fall within the first, second and third distance bin, respectively, and n_T is the total number of picks. Similarly, for effective S-wave travel-time uncertainty we use

$$\sigma_s = (0.186n_1 + 0.322n_2 + 0.568n_3)/n_T. \quad (5.2)$$

Eqs. 5.1&5.2 entail that the effective travel-time uncertainty increases when more distant picks are used in the location assessment. In these cases, a more constrained source location might be achieved when leaving out more distant picks. Whether it is favourable to include more distant picks depends on the station coverage that is already achieved with local or regional picks.

We test one approach of removing more distant picks. We assume that a good coverage is achieved when there are at least 9 picks available in smaller distance bins. Thus, when at least 9 picks are available in the local distance bin, picks are removed from the regional and national bin. Likewise, if at least 9 picks are available in the local and regional bin combined, picks are removed from the national bin. If say 20 picks are available in the local distance bin, they are all used, because the larger amount of picks (with the same intrinsic uncertainty) only helps to constrain the source location.

Fig. 5.1 shows a case in which the pick-restriction approach is favourable. A source is chosen in the province North Holland ($[RD_x, RD_y, \text{depth}] = [110, 520, 3]$ km) with $M=2.0$. Picks can be made at 12 stations, 8 of which are local, 2 regional and 2 national (Fig. 5.1a). The last two picks are at distant low-noise stations: T064 (Twente) and TERZ.01 (South Limburg). Using all the available picks leads to the epicentral uncertainty ellipse as shown in Fig. 5.1(b) in green. After removing the two most distant picks, a smaller epicentral location uncertainty is obtained (purple ellipse in Fig. 5.1b).

Also for both cases, the depth uncertainty is computed. When all picks are used, $\sigma_Z=472$ m. When the two most distant picks are omitted, $\sigma_Z=446$ m.

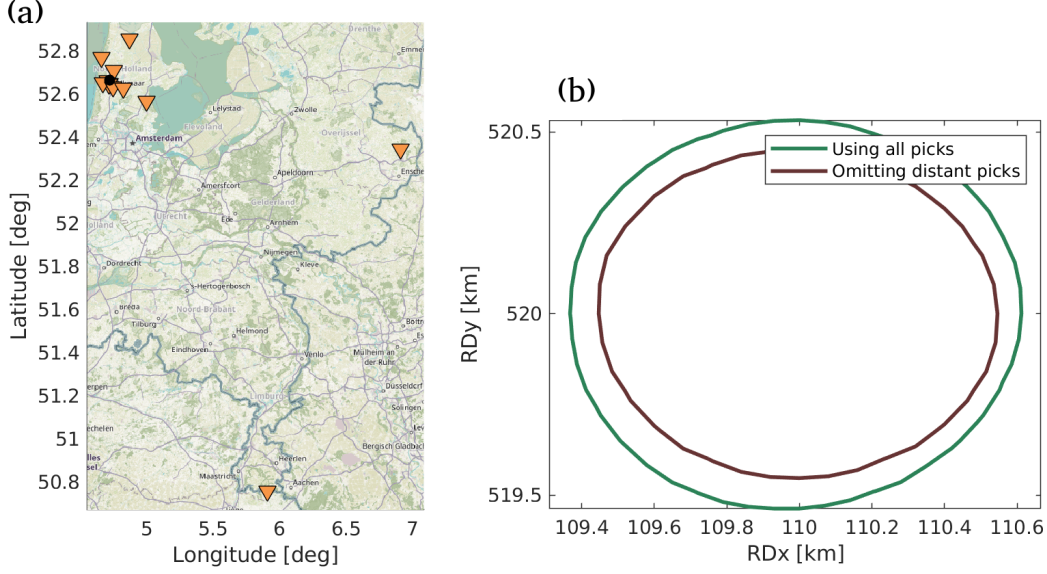


Figure 5.1: (a) Stations (orange triangles) where likely picks can be made for a $M=2.0$ source in North Holland (black dot) at 3 km depth. (b) Size of the 95% confidence area when all stations indicated in (a) are used (green ellipse) or when the two most-distant stations are omitted (purple ellipse). This figure is for illustration purposes only. It was made with an obsolete version of the v_{rms}^{P90} values.

The above pick restriction is tested for all locations in the Netherlands. Fig. 5.2 shows σ_1 results for a part of North Holland, when (a) the pick restriction is implemented and (b) when the pick restriction is not implemented. It can be seen that small gains are indeed achieved at some place. That is, Fig. 5.2(a) contains areas with colder colours than Fig. 5.2(b). On the other hand, Fig. 5.2(a) also contains areas where the epicentral location deteriorates when the pick restriction is implemented, leading to warmer colours than in Fig. 5.2(b). Overall, the effect of the picking-restriction approach (i.e., difference between Fig. 5.2(a) and Fig. 5.2(b)) is small.

For implementing an ideal pick restriction, for each source location a grid search would need to be done on many possible omissions of more distant picks. This additional grid search could lead to a hundred-fold computation increase, while bringing only modest gains (Fig. 5.1). For that reason, a more refined pick-restriction approach is omitted and simply all 40 closest picks (when available) are included for location.

5.1.2 Grid-search settings

The epicentral search is done within a square that is defined around the event location. For magnitudes 0.5 until 1.5, the square is defined with parameter $a = 2(11 - 9/27p)$, where p is the amount of P-wave detections for the event. The square has sides with size a in km and the actual event location is always in the middle of the square. One side is at least 2 km. Hence, in the limiting case when location is done with only 3 detections, the size of search grid is 20x20 km. With 30 or more detection, the area is reduced to 2x2 km. Irrespective of the size of the area, the zone is sampled with 10000 equally spaced points to obtain the epicentral PDF.

For magnitudes 2.0 until 4.0, the square is defined with parameter $a = 2(31 - p)$. For

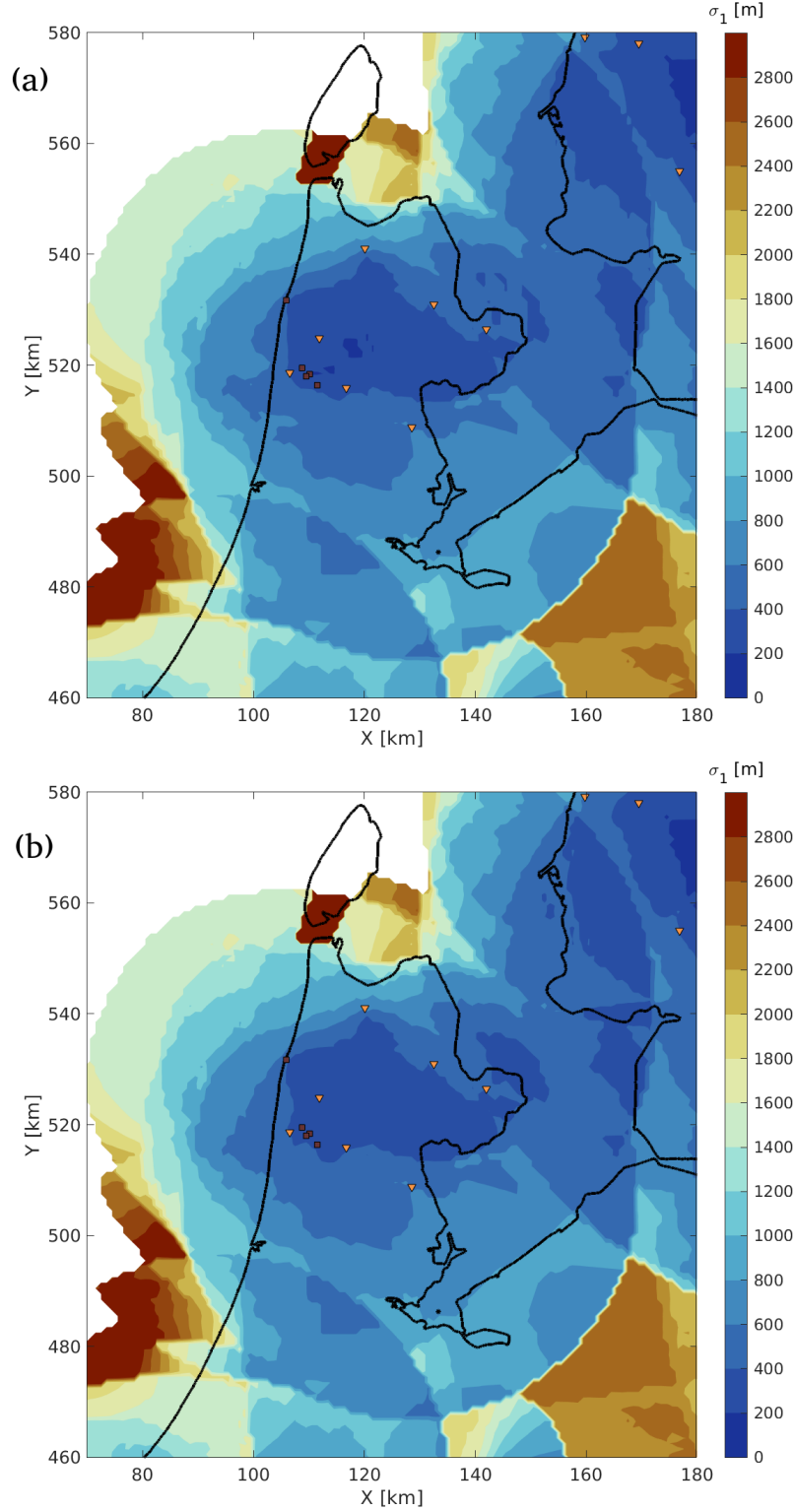


Figure 5.2: σ_1 location uncertainty in North Holland and surroundings when (a) removing picks from large-distance bins if nine local or regional picks are available and when (b) leaving out the most distant picks if more than 40 picks are available. The epicentral uncertainty has been computed for $M=2.0$. Thick black lines denote the coastline. Dark red depicts uncertainties of 2800 m and higher. This figure is for illustration purposes only. It was made with an obsolete version of the v_{rms}^{P90} values.

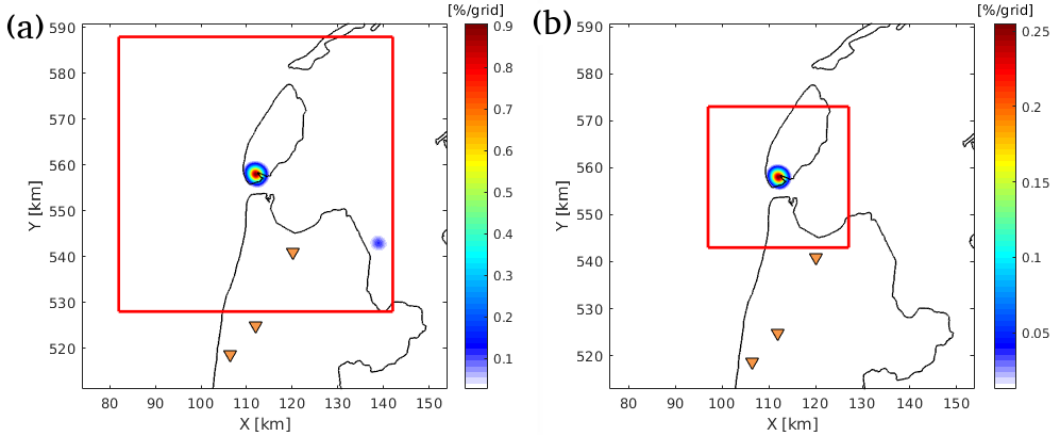


Figure 5.3: Grid search within (a) 60 km rectangular search grid (as depicted with the red square) and (b) 30 km search grid, yielding $\sigma_1=8552$ and $\sigma_1=1516$ m, respectively. Note that in (a) the 95% epicentral confidence zone includes a small area in the southeastern edge of the grid-search area, which is responsible for the large extend of the fitted σ_1 value. Examples are for $M=2.0$ with $[rdx, rdy]=[112, 558]$ km. This figure is for illustration purposes only. It was made with an obsolete version of the v_{rms}^{P90} values.

these magnitudes, more extreme cases occur with detections at only a few distant stations. This gives the need to extend the grid-search area, to cover potentially very large PDFs. For these magnitudes, when there are only 3 detections, the size of the grid-search area is 56x56 km. When there are 30 or more detections, the area is reduced to 2x2 km.

The above settings have been tested for many different sensor configurations, to make sure that the grid search contains the area where nearly all location chances lie. In the testing, one case was found where these settings may not be desired. This is a case with a near-linear array of receivers as shown in Fig. 5.3(a). In this example there is an ambiguity. A large part of the probability lies around the actual source location, but some part of the 95% confidence zone lies in a separated blob that is a blurred mirror image (with the line extending from the receivers acting as the mirror) from the actual event location. Through these separated areas, an ellipse is fitted, which results in an ellipse with a large major axis and hence a large apparent σ_1 value. When using an outdated v_{rms}^{P90} list, this occurred for $M=2.0$ in a zone at the southern side of Texel (Fig. 5.2). In these cases that the confidence area is split up in two zones that are well separated, it would be possible to find the actual source area by using the polarization of the incoming P waves. In order to show only the parameters of the main blob, for this case the grid-search area could be reduced to a square of 30x30 km yielding the results as shown in Fig. 5.3(b).

To keep a fixed procedure, we have not implemented the above fix. This means that for some areas with very large modeled σ_1 values, the uncertainty could be reduced in reality by including a polarization data attribute.

The discretization settings for computing the depth PDF are as follows. A search in depth is done from 0 to 20 km with steps of 50 m. The actual source is placed at 3 km depth. As such, the PDF is well sampled for any receiver-distribution scenario. As only 401 points are evaluated in the computation, no further optimizations are needed to save computational time.

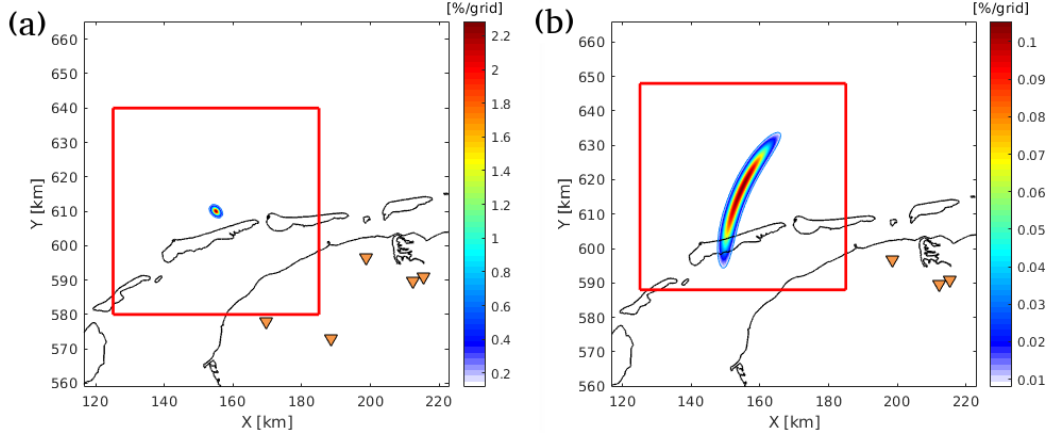


Figure 5.4: Large difference in uncertainty of nearby source locations (a) $[rdx, rdy]=[155, 610]$ and (b) $[rdx, rdy]=[155, 618]$ because of large difference in receiver (orange triangles) azimuthal coverage. σ_1 is (a) 1132 and (b) 11314 m, respectively. Examples are for a $M=2.0$ at 3 km depth. This figure is for illustration purposes only. It was made with an obsolete version of the v_{rms}^{P90} values.

5.1.3 Discontinuity in uncertainty

When generating test maps, cases were found where the uncertainty varies largely from one to the other grid point. These discontinuities were seen as possible implementation errors. However, after inspection, it turned out that these discontinuities would actually occur. One of the cases is shown in this section.

In Fig. 5.4 an example is shown of σ_1 increasing about ten fold by moving the scenario source location just a few kilometers. In Fig. 5.4(a) the uncertainty is relatively small with 5 receivers picking up the signal from this event. When the event is moved a few kilometers to the north (Fig. 5.4b) only 3 receivers, in a more or less linear configuration, pick up the event. Like in Fig. 5.3(a), the probabilities lie in two mirrored confidence zones, but in this case these two zones have merged, which makes it harder to disentangle them. This merging of two confidence zones occurs when the source location lies in the extension of a linear configuration of receivers and the actual source location is also close to the line that can be fitted through the receivers. In other words, these extended confidence zones occur if the azimuthal gap is very large.

5.2 Epicentral uncertainty

In this section, epicentral uncertainty is computed for hypocentres at 3 km depth. For this depth, the epicentral probability density function (PDF) is computed. This PDF is approximately a bi-variate normal distribution, of which a confidence zone typically looks like an ellipse. This PDF is parameterized with σ_1 , σ_2 and θ . σ_1 and σ_2 describe the uncertainty in the major-axis and minor-axis directions of the ellipse, respectively (Fig 3.1c). θ describes the orientation of the ellipse. It is defined as the smallest clockwise angle between north and the major axis.

In this section we show the computed values for specific areas in the Netherlands for an event depth of 3 km. The effect of event depth on epicenter uncertainty has been studied in Section 4.2. σ_1 -maps for the Netherlands can be found in Appendices B.1&D.1 .

Fig. 5.5 introduces the plotting style that has been chosen. Fig. 5.5(a) shows the grid cells for an area around North Holland. The colour of the grid cell denotes the modeled σ_1 value for a scenario M=1.5 event in the centre of the grid cell. A white grid cell means that for the center point there is an insufficient number of detections (< 3) and hence no location, nor location uncertainty, can be computed.

In Fig. 5.5(b) the visualisation is simplified by removing the grid lines and discretizing σ_1 with 200 m levels. Hence, the lowermost uncertainty level, denoted by dark blue, contains all values from 0 to 200 m. Higher uncertainties are shown with warmer colours. Dark orange denotes uncertainties between 2600 and 2800 m. Red denotes values of 2800 m and higher. In practise, there are little places where $\sigma_1 > 3000$ is reached.

In Fig. 5.5(c) the visualisation is made more eye-pleasing by interpolating the borders between different uncertainty levels. This is the plotting style that is chosen in this report. The exact place where borders are drawn depends on the interpolation scheme that is used. This post-processing leads to small perturbations. Thus, this map gives a good overview of the uncertainties, but for the actual value at a position, one needs to resort to the value of the underlying grid cell.

Fig. 5.6 gives an overview of the different epicentral uncertainty parameters. Again the same part of North Holland is shown as in Fig. 5.5 for scenario events with M=1.5. Fig. 5.6(a)&(b) show the σ_1 and σ_2 distribution, respectively. For the entire area where a location can be made, σ_2 remains small to modest (< 1200 m). Outside of the area covered by the local network, however, there are areas where σ_1 becomes large (≈ 3000 m). Fig. 5.6(c) shows the aspect ratio σ_1/σ_2 . In the same areas where σ_1 is large, also the aspect ratio is significant. This indicates that the uncertainty ellipse is elongated. Also Fig. 5.6(d) shows the size of the aspect ratio, by the size of the stripes. At the same time, the orientation θ of the uncertainty ellipse is shown through the orientation of the stripe. In an area surrounded by geophones, the aspect ratio is very small, resulting in points instead of stripes. At some places, the orientation of the ellipse, and the aspect ratio, vary smoothly as function of distance from a station. At places where there is a discontinuity in detectability (addition or removal of stations that can be used for location) there is also a discontinuity in aspect ratio and/or ellipse orientation. Altogether, the interplay of irregularly placed stations with different detection capabilities (noise levels) results in a complicated spatial pattern of epicentral uncertainty.

From the three parameters describing epicentral uncertainty, one would generally be interested in the largest uncertainty σ_1 . Overview maps of σ_1 for the complete zone of investigation (the Netherlands and border areas) can be found in Appendices B.1&D.1. They have been computed for M=0.5 to M=4.0 with steps of 0.5. For each magnitude and each grid point, the values of θ , σ_1 , σ_2 and ψ are available at KNMI (2023b).

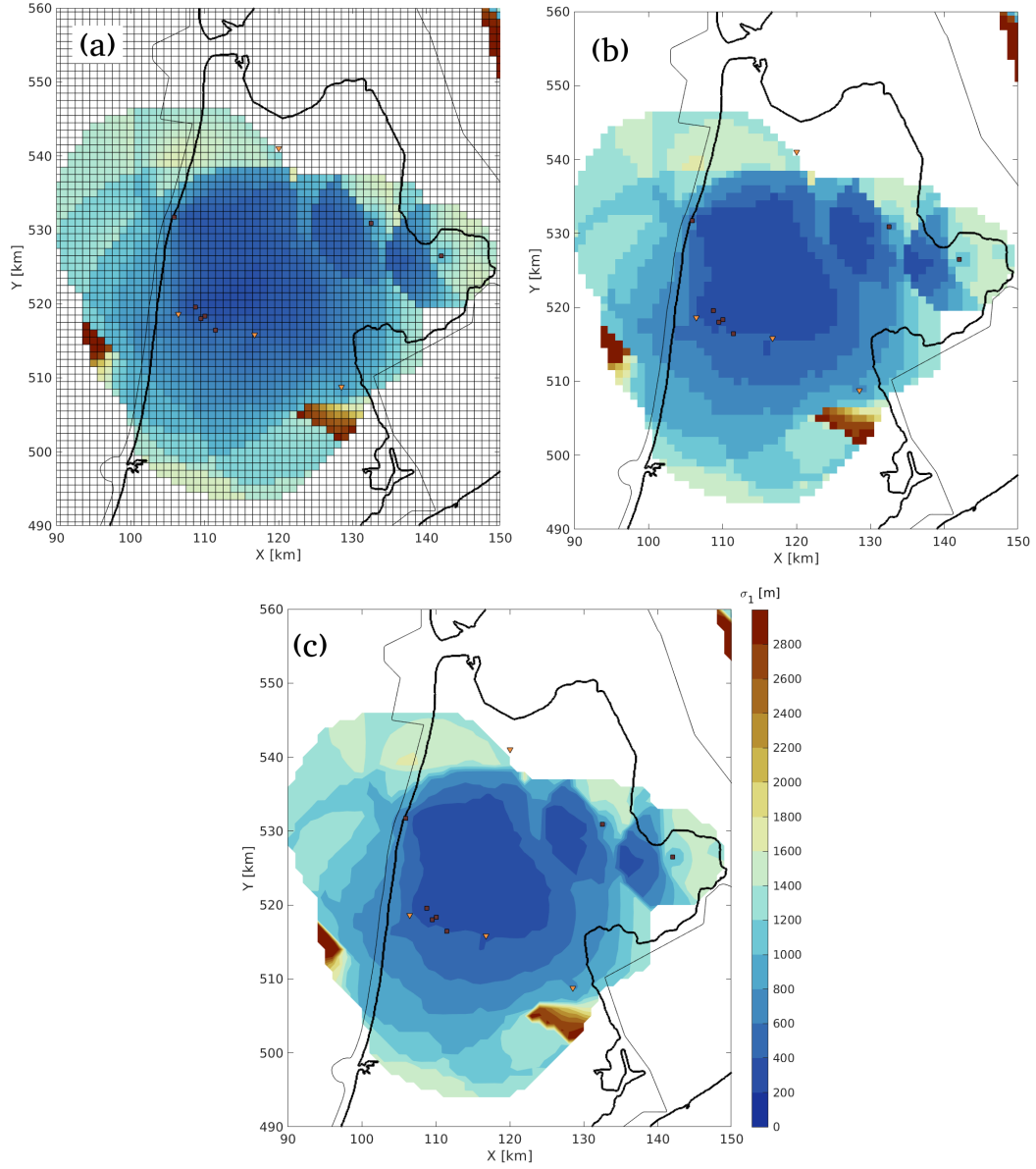


Figure 5.5: Comparison of plotting styles for showing the σ_1 distribution over North Holland for scenario events with $M=1.5$; (a) With color gradient and showing the underlying grid, (b) with one colour for each 200 m bin and (c) with interpolated contours between the different σ_1 zones. The coast is depicted with a thick black line, province boundaries with a thin black line, surface sensors are shown with squares and deep geophones with inverted orange triangles. Coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

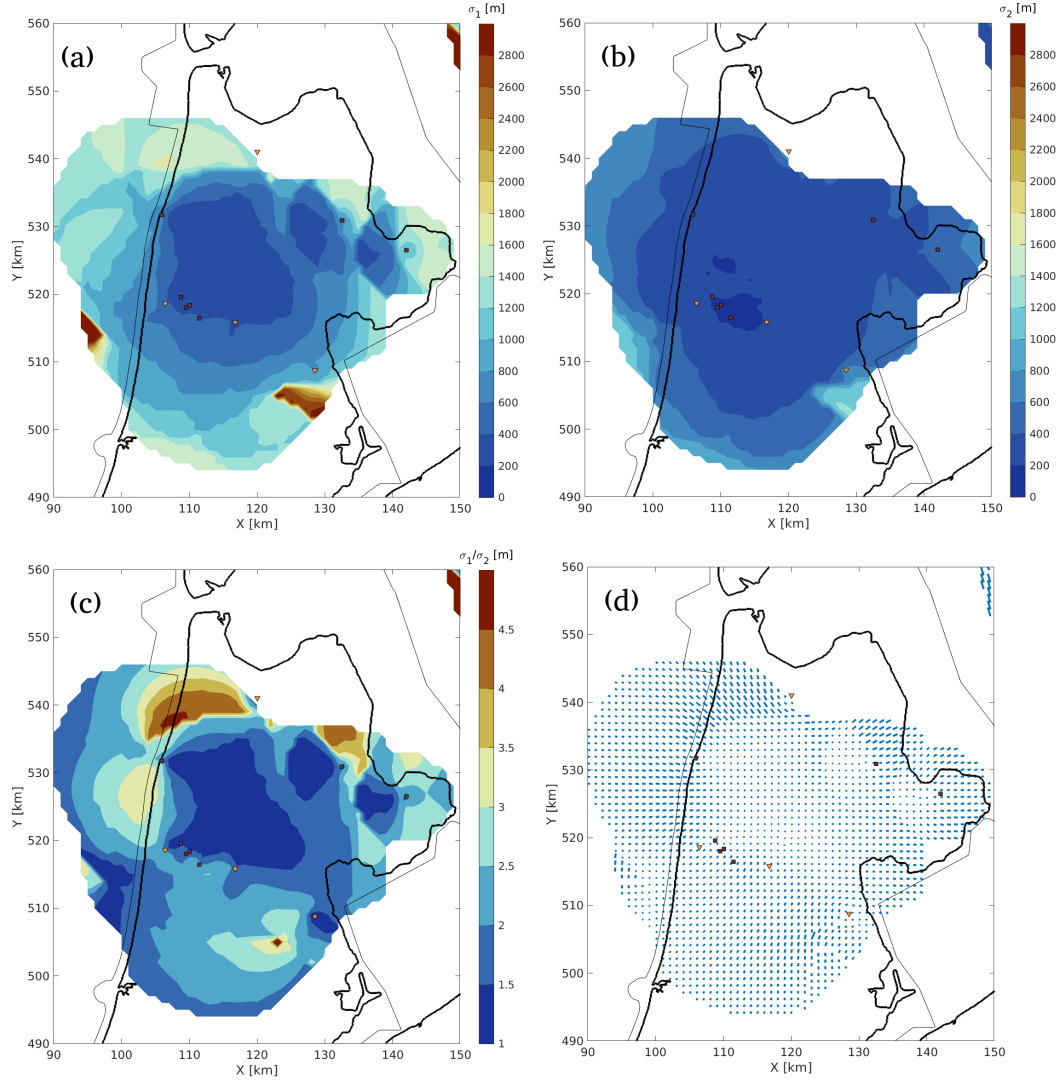


Figure 5.6: Epicentral uncertainty plots for North Holland, for scenario events with $M=1.5$. (a) Uncertainty in direction with largest uncertainty σ_1 , (b) uncertainty in direction with smallest uncertainty σ_2 , (c) aspect ratio between the two orthogonal uncertainty directions in the horizontal plane: σ_1/σ_2 . (d) a quiver plot showing the aspect ratio (through size of the stripes) and orientation of the uncertainty ellipses (through orientation of the stripes), for a 1x1 km grid of scenario events. Coordinates are in the local rectangular grid (Rijks-driehoekstelsel) in kilometers.

When inspecting the σ_1 figures for the different magnitudes (Fig. B.1–B.8) one can see that, first of all, the area where a location can be made grows with magnitude. Also, generally, the uncertainty reduces with magnitude. This is to be expected, since higher magnitudes lead to higher amplitudes, more detections and thus a better receiver coverage. Within the onshore part of the Netherlands, the location uncertainty keeps reducing until $M=3.0$. For larger magnitudes, the location accuracy remains practically stable.

Higher magnitudes lead to inclusion of picks from further distances, which results in an increase in the effective travel-time uncertainty. At some places, the increased station coverage is undone by the increased travel-time uncertainty, leading to a (small) increase in location uncertainty with magnitude. These places can, e.g. be noted by comparing Fig. B.7 and Fig. B.8. In Section 5.1.1 we discussed this effect. It would be possible to undo the (small) increase in σ_1 by finding an optimal combination of picks to include for the location. Because the gain is small, and the additional computational expense large, this fix has not been implemented.

5.3 Depth uncertainty

In this section it is shown how the expected depth uncertainty varies over the Netherlands. The depth PDF is computed at the epicentral location for sources at 3 km depth. The last panel in Fig. 1.1 shows an example. The depth PDF is parameterized with σ_z which is approximately one-fourth of the depth interval that contains 95% of the location probabilities.

σ_z -maps for the Netherlands can be found in Appendices B.2&D.2. The underlying values can be obtained at *KNMI* (2023b). When inspecting the σ_z figures for the different magnitudes (Figs. B.9–B.16) one can see that, first of all, the area where a location can be made grows with magnitude. Also, generally, the uncertainty reduces with magnitude. However, the uncertainty reduction with magnitude is more limited than for σ_1 , because σ_z is especially dependent on the sampling of the travel-time curve at near offsets, which sampling does not further improve for higher magnitudes. The σ_z maps for $M=3.5$ and $M=4.0$ are nearly identical.

The general pattern of σ_z can be studied on, e.g., Fig. B.12. The further an event location is from the nearest station, the larger the depth uncertainty. The smallest location uncertainties (< 500 m) are reached in areas with the densest station distributions. This occurs for the dense Groningen network with receiver spacings of around 4 km, but also for dense accelerometer distributions in Friesland and Drenthe. For the area covered by the networks in Twente, South-Holland and North-Holland there is a larger depth uncertainty ($500 < \sigma_z \leq 1000$) than in Groningen. In the former areas there would, overall, be less travel-time picks available to constrain the location than in Groningen and surroundings.

Fig. 5.7 shows $M=1.5$ σ_z maps for two areas in the Netherlands. σ_z is shown in more detail than in the appendix; in Fig. 5.7 the σ_z values are discretized with steps of 200 m, in Fig. B.11 with steps of 500 m. In Groningen (Fig. 5.7a), the quite uniform receiver layout results in a quite uniform $\sigma_z < 400$ m over most of the network. In the southeast of the area, there is an additional infill network to monitor gas storage in salt caverns: the Zuidwending network. In the area covered with this very dense network, $\sigma_z < 200$ m is achieved.

In Fig. 5.7(b) the station distribution is more patchy. Near two geophones in Twente $\sigma_z < 600$ m is achieved. The smallest uncertainties ($\sigma_z < 400$ m) are obtained here near the accelerometers in South Drenthe. Because these accelerometers are still relatively close to the G-network, one obtains more travel-time picks for event locations in this area, than for locations in Twente.

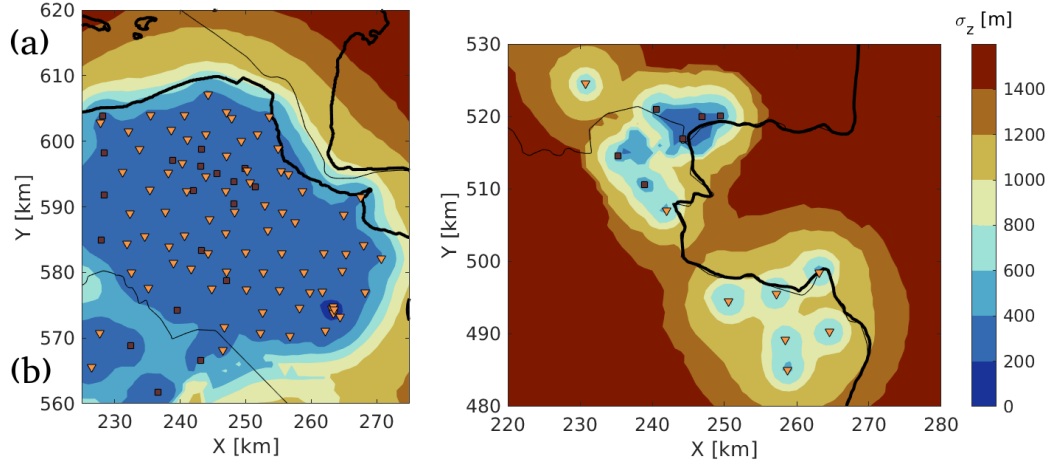


Figure 5.7: $M=1.5$ depth-uncertainty σ_z distribution for two areas in the Netherlands: (a) Groningen and (b) an area covering the south of Drenthe and Twente. Accelerometers are shown with purple squares, deep geophones with orange triangles. Uncertainties of 1400 m and higher are shown in red.

5.4 Magnitude of completeness

In this section we use the tools developed in Chapter 2 to update the magnitude-of-completeness (MoC) map for the Netherlands. This new map supersedes the map published in *Kruiver et al.* (2021). The MoC at a certain point is the minimum magnitude for which an earthquake can be located in severe noise conditions. This corresponds to a minimum of three detections that are likely to occur in the presence of P90 noise conditions. The map is computed for a typical induced-seismicity focal depth of 3 km. The map is made using a 1x1 km grid. A contoured version of this map is published in this report. The computed values for each grid point can be found at *KNMI* (2023a).

Computation of a MoC (for one point) entails the following steps:

- Compute the expected amplitude at all receivers in the detection list, using the P-wave GMPEs (Section 2.2).
- Repeat this computation for a range of magnitudes.
- Implement the picking threshold (Section 2.4).
- For hard-rock sites, implement a hard-rock correction factor (Section 2.5).
- Compare the resulting amplitudes with the station P90 noise levels (Section 2.3) to find the minimum detection magnitudes at each receiver.
- Order this detection-magnitude-per-receiver list from low to high. The third-lowest magnitude is the MoC for the evaluated source location.
- MoC values lower than 0.4 are replaced with 0.4. Clipping at MoC=0.4 is done to make sure that the modeled outcome falls within the range for which the underlying P-wave GMPE has been calibrated.

The term magnitude of completeness implicates that the catalogue is complete from the stated MoC. With other words, if the MoC for a certain region is 0.5, it implicates that all events with $M=0.5$ and higher are detectable and locatable. However, one needs to take into

account that also the modeled MoC is a probabilistic value. It is computed using the 90th percentile of the noise distribution, meaning that it yields MoC values that can very often, but not always, be reached in practise. If the P90 MoC value is 0.5, then 90 percent of the time a $M=0.5$ event would be locatable. However, there remains a 10 percent chance (e.g., during stormy days) that noise levels are so high that a $M=0.5$ event would be missed.

5.4.1 Sensitivity tests

Fig. 5.8 shows the impact of noise-percentile choice on the estimation of MoC. Using the P10 from the noise, a very optimistic magnitude distribution is obtained as shown on Fig. 5.8(a). When noise is very low (during quiet summer nights) a magnitude of 0.8 (cyan contour) can be detected and located over a large part of North Holland. For average noise conditions (P50, Fig. 5.8b) the area where a magnitude of 0.8 can be located is significantly smaller. When the 90th percentile of the noise is used (P90, Fig. 5.8c) an 0.8 event can only be located in an area around Alkmaar. This last map shows the magnitude distribution we refer to as MoC. For the zone indicated with $\text{MoC}=0.8$ on this map, a nearly complete catalogue could be reached for $M \geq 0.8$. So, it is the P90 level we further use when making MoC maps.

Another point of attention is the choice of three detections. The number of three makes sense from a location perspective, since it is the minimum number for which a 3D location of the source can be found. It is not the amount of receivers, however, that is currently used in an operational sense at KNMI. At KNMI the software package SeisComP (*GEMPA*, 2022) is used, which only starts up a location algorithm after detection at six stations. Fig. 5.9 shows, for the Groningen region, the MoC that is obtained when requiring (a) detection at six receivers and (b) at three receivers. Clearly, lower MoC levels can be reached when only three detections are needed. Especially for monitoring the gas-storage plants it was deemed important to reach these lower MoC levels also in an operation sense. This was achieved by placing also the second-deepest geophones in the detection list, for the geophone strings at the Norg and Grijpskerk networks. For the Groningen gas field, Fig. 5.9(b) shows the MoC that could be reached by post-processing the automated picks.

5.4.2 2021 Map

Fig. 5.10 shows the MoC map for the Netherlands, with a contouring of 0.5 magnitude levels. At regions where dense distributions of borehole geophones exist, $\text{MoC} \leq 0.5$ is reached (dark-blue areas). An exception is for the network in South Holland, primarily due to high noise levels (Fig. 2.14). In the south of Limburg, $\text{MoC} \leq 0.5$ is reached with broadband stations, with some of them having very low noise conditions. Around the regional networks, there are zones in which $0.5 < \text{MoC} \leq 1.0$ is reached (blue), and at further distance from the dense station distributions, a zone in which $1.0 < \text{MoC} \leq 1.5$ is reached (light blue). In areas far from station clusters, like most of Utrecht and Gelderland, $1.5 < \text{MoC} \leq 2.0$ (yellow). In the onshore part of the Netherlands $\text{MoC} \leq 2.0$ with the exception of a tiny part of Zeeland. Most of the offshore region on the map has $2.0 < \text{MoC} \leq 2.5$ (orange).

Note that the map does show parts of Germany and Belgium, but would not be correct for these surrounding countries. The seismological authorities across the border use additional stations of their own networks. Also take into consideration that we did not include British stations (nor Danish and Norwegian stations) in this study. Hence, further offshore the MoC could be lower than modeled in this study.

Within the dense borehole network in Groningen, a MoC of 0.4 is achieved (Figs. 5.9

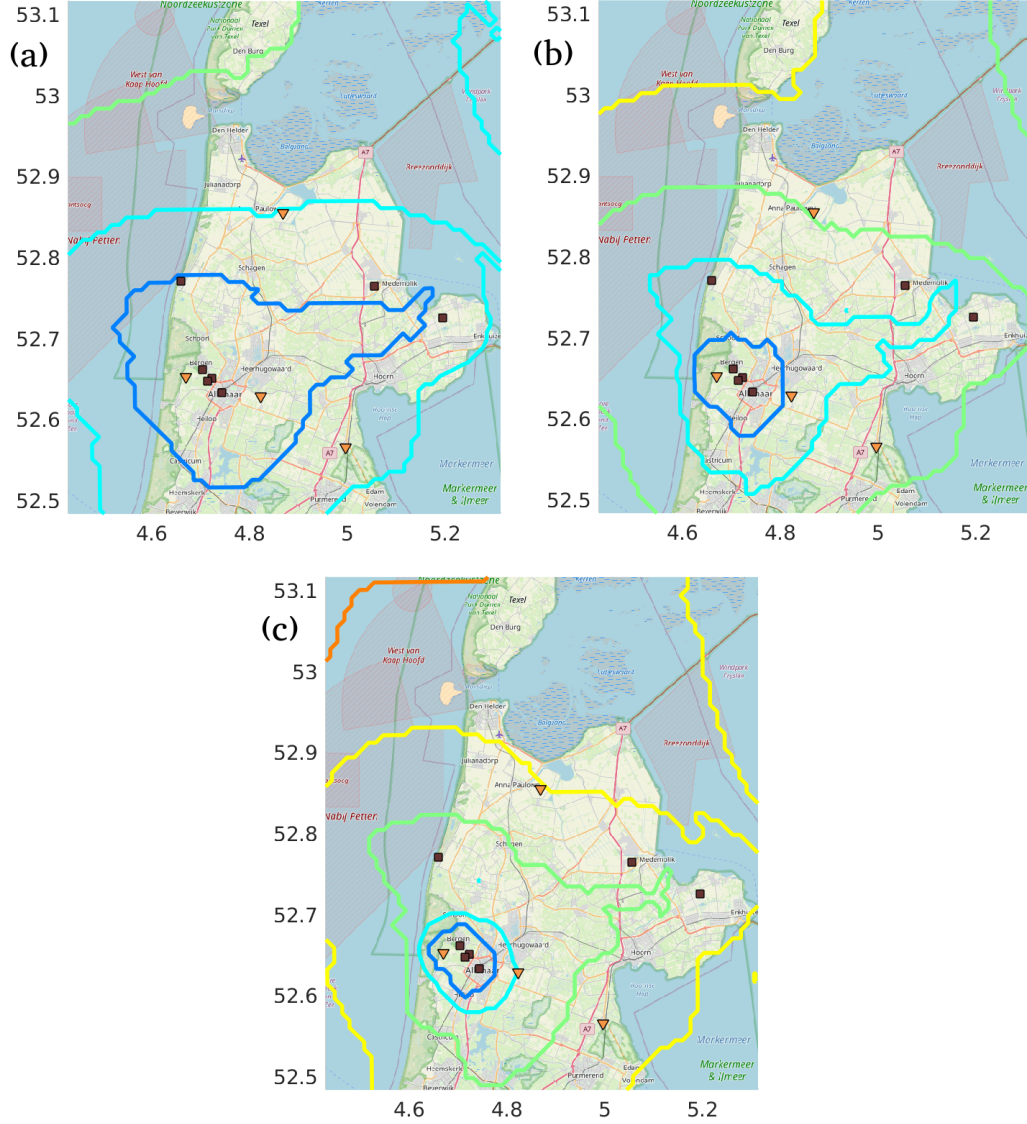


Figure 5.8: Magnitude-of-completeness (MoC) estimation at the northwestern part of the Netherlands, using (a) the 10th percentile, (b) the 50th percentile and (c) the 90th percentile of the noise distribution (Section 2.3). The actual MoC is obtained using the P90 noise distribution. Contours are shown for MoC values of 0.4 (blue), 0.8 (cyan), 1.2 (light green), 1.6 (yellow) and 2.0 (orange). Surface sensors are shown with purple squares, deep geophones with orange triangles. Background map is from www.openstreetmap.org.

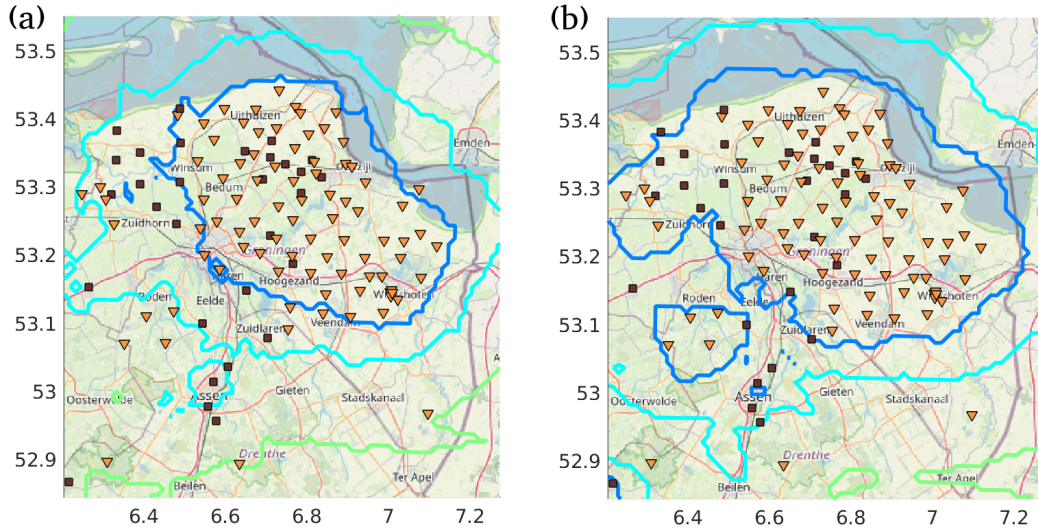


Figure 5.9: Magnitude of completeness (MoC) for the Groningen area using a minimum of (a) six detections and (b) three detections. Contours are shown for MoC values of 0.4 (blue), 0.8 (cyan) and 1.2 (light green). Background map is from www.openstreetmap.org.

and 5.10). When MoC is computed with historical events from the KNMI induced event catalogue, a higher MoC would be obtained. There are different reasons for this:

- The modeled MoC would always be slightly lower since it is based on P90 noise levels. This means that events with magnitudes close to MoC would just be missed in severe noise conditions.
- For determination of a local magnitude, the records are first convolved with a Wood-Anderson response function. This convolution acts as a high-pass filter for frequencies larger than 1.0 (*Chung and Bernreuter, 1981*). In the Groningen area there is severe low-frequency noise (0.4 to 2.0 Hz) due to the nearness of the Wadden Sea and North Sea. As a result, a part of the low-magnitude events has sufficient SNR to be detected and located using the 5-to-40 Hz frequency band, but cannot be assigned a local magnitude due to insufficient SNR at frequencies between 0.9 and 5 Hz. Currently, an analogue local-magnitude relation is derived that is especially suited for the small-magnitude events. For the time being, the KNMI induced data catalogue misses some of the locatable events for approximately $M \leq 0.6$.
- The modeled MoC is for the network layout of September 2021, whereas measured MoC is based on events in the past in which the network was less well developed.

From the above list, it is expected that the second item is the most important; for the Groningen region, over the years a long list was built up with events that were too small to assign a local magnitude. For other regions, there are no such lists. It is hence expected that for other regions, the modeled MoC would come quite close to an MoC as derived from the published KNMI catalogue. For example, an M0.3 event in Twente —if it were to happen— can always be assigned a magnitude since there is no strong noise source between 0.9 and 5 Hz. In South-Holland there is a similarly strong low-frequency noise source as in Groningen (i.e., the active coast). However, in South-Holland the $\text{MoC} \geq 0.8$. An M0.8 would yield sufficiently strong low-frequency amplitudes to assign a local magnitude.

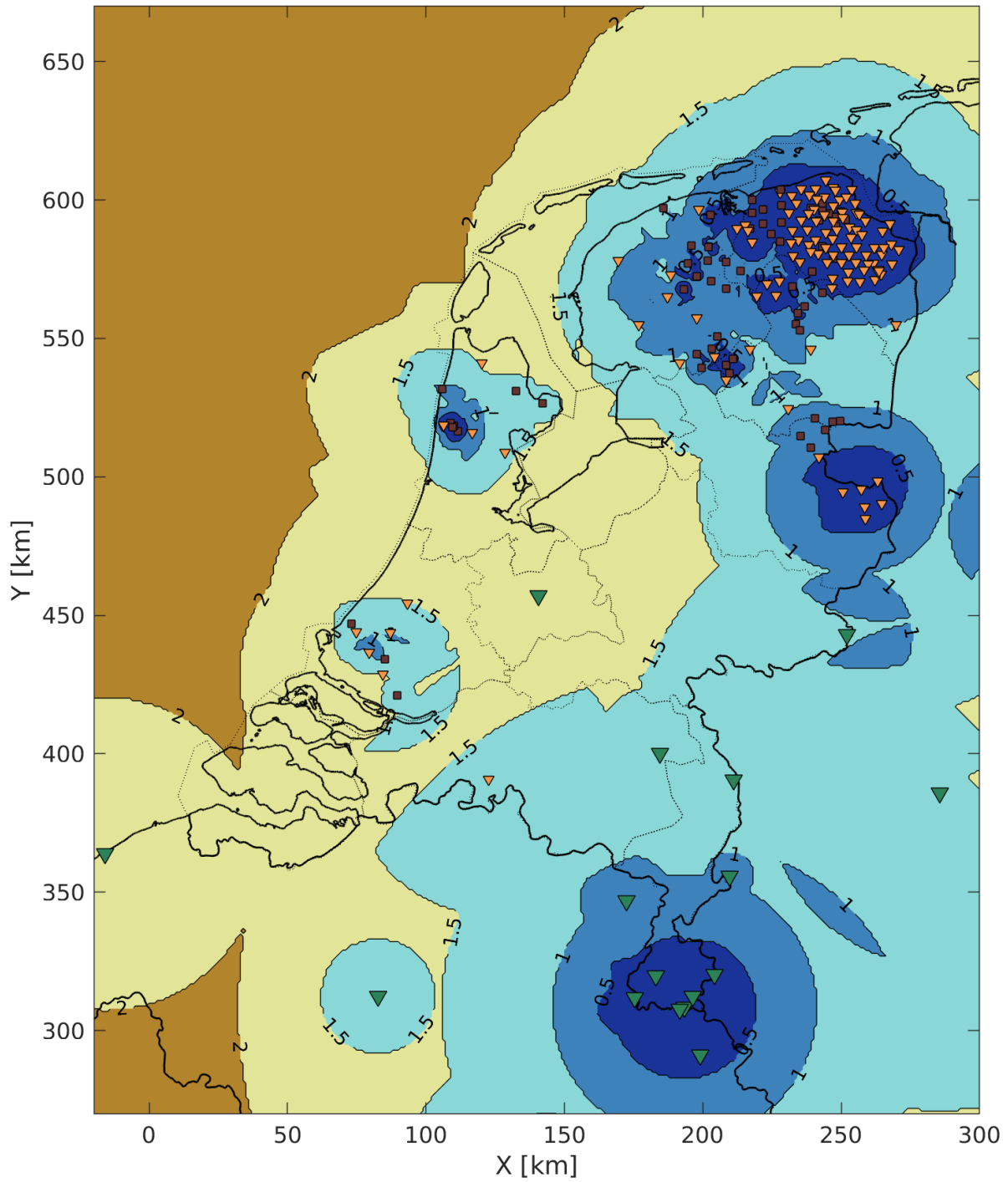


Figure 5.10: Magnitude of completeness over the Netherlands, computed using P90 noise levels in 2021, detections at three stations, and the station list of September 2021 (Fig. 2.2).

5.4.3 2022 Map

Also for the network state of September 2022 the MoC is computed. For the 2022 map, data from September 1 2021 until August 31 is used for computing the v_{rms} values. The generally small changes in noise conditions result in small changes in the MoC. Larger changes in MoC are caused by changes in the effective network layout. The following changes exist from the September 2021 to September 2022 detection list:

- Addition of 9 stations that had insufficient data over 2021, but sufficient data over 2022 (G024, G314, STR, SUH4, T014, WDB4, WMH4, WYN4, AHRW).
- Addition of 2 new borehole geophones (ZH074 and VDM24).
- Addition of 8 new accelerometers (DOK, KMP, OOTH, SNB, EVK, PMR, BRTL and OLK).
- Removal of 1 accelerometer that was closed in 2022 (BHAR).
- Removal of geophone ZWE2 because of bad data quality.
- Replacement of borehole geophone CHA5 with surface broadband station NE116.

The locations of new stations can be found at, e.g., <http://rdsa.knmi.nl/network/NL/>.

Fig. 5.11 shows the MoC for 2022. Almost everywhere the MoC is either stable or improves with respect to the 2021 version (Fig. 5.10). Only in north of Friesland, the MoC somewhat deteriorates due to the removal of geophone ZWE2 from the detection list. It can be seen that both the MoC=1.5 area (cyan) and MoC=1.0 area (blue) become somewhat smaller here.

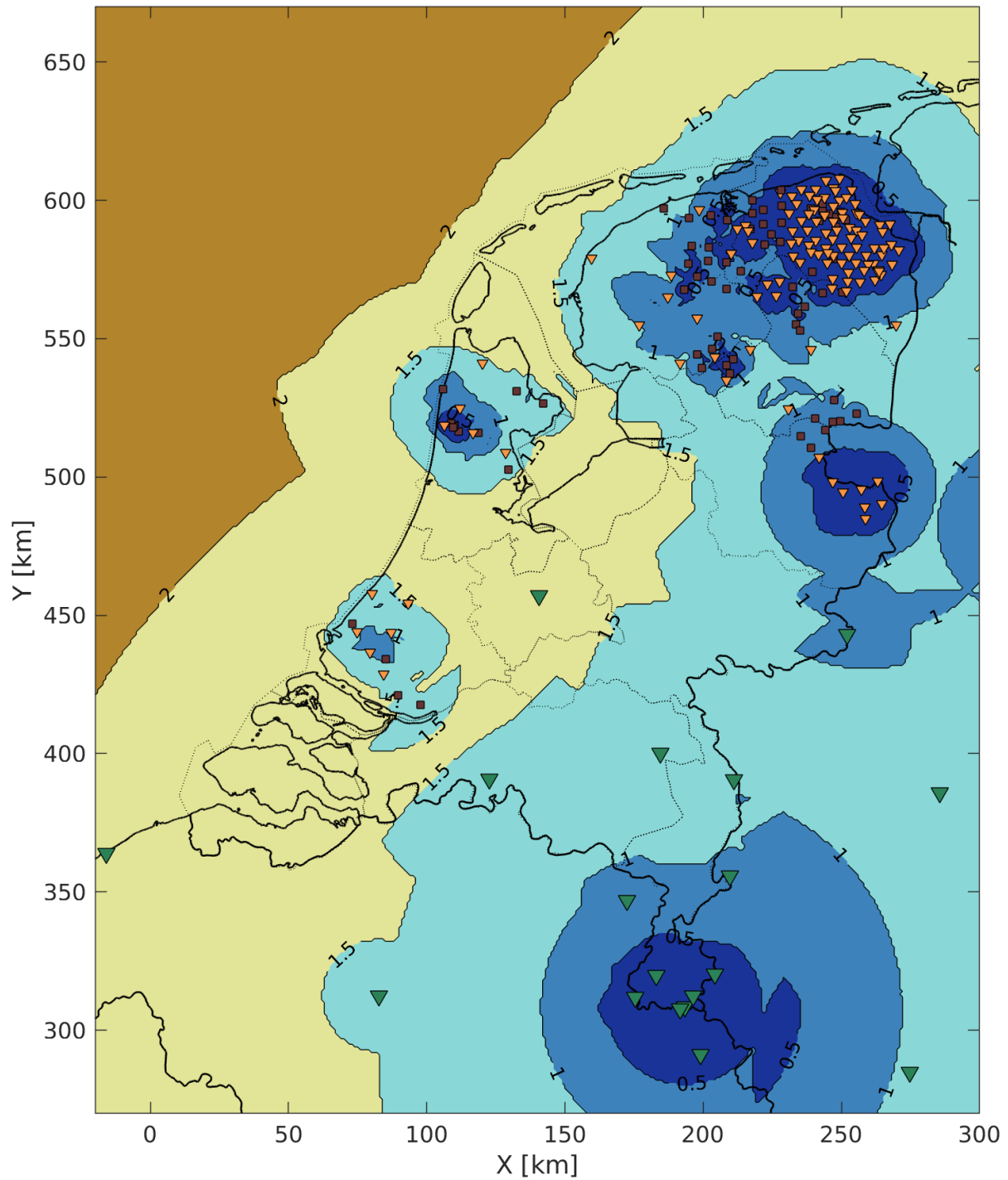


Figure 5.11: Magnitude of completeness over the Netherlands, computed using P90 noise levels in 2022, detections at three stations, and the station list of September 2022.

6 Discussion and Conclusions

In this report, location-uncertainty maps were derived for the Netherlands, using the network configuration as existed in September 2021 (Fig. 2.2 and Table A.1) and September 2022 (Table C.1). The resulting maps can be found in Appendices B&D and the underlying values can be obtained at *KNMI* (2023b). Yearly future updates are anticipated.

As a side-product, magnitude-of-completeness maps were updated. The 2021 and 2022 versions of these maps can be found in this report (Figs. 5.10&5.11) and the values for each grid point can be obtained at *KNMI* (2023a).

In this chapter, the produced maps, and the underlying results from the previous chapters, are summarized and discussed.

6.1 Detection

For modelling whether a detection can likely be made, P-wave ground-motion prediction equations (GMPEs) were derived (Section 2.2). These equations were specifically derived for seismicity in the upper crust, for a frequency band between 5 and 40 Hz, which is used for detecting and locating induced seismicity. They are empirical models that are constrained with a database of more than 20,000 PGV recordings from induced seismicity in the Netherlands. The models have an empirical base between local magnitudes of 0.4 and 3.6, epicentral distances between 0 and 120 km and earthquake depths between 2 and 3 km (or hypocentral distances between about 2 and 120 km). A model was derived for motions at the free surface (SL model) and at around 200 m depth (DL model). Both models are the same, except for the static-offset fitting term (c_1 in Eq. 2.1). This term is 4.06 times larger for the SL model than for the DL model (Table 2.1).

The derived GMPEs are specifically for sensors at soft-sediment settings. For calibration, only sensors in or at unconsolidated sediments were used. Most sensors at depth are in a P-wave velocity setting between 1.6 and 2.2 km/s (*Hofman et al.*, 2017). Near the Earth's surface, P-wave velocities can be significantly lower (*van Ginkel et al.*, 2022b). For sensors installed at bedrock conditions, a correction factor was derived (Section 2.5) to account for the higher seismic impedance and hence lower recorded amplitudes.

S-wave GMPEs were not used in this project. We assumed that, when a P-wave can be detected and picked, also an S-wave can be picked and used for locating the event. As nearly always S-waves are stronger than P-waves, this is mostly a valid assumption. Still, it can be argued that this assumption is both optimistic and pessimistic. We think that the net effect remains that it is a good assumption to make.

When a station happens to be in or close to a nodal point of the S-wave radiation pattern, the S-wave amplitude would be smaller than the P-wave amplitude. So, in a few percent

of the cases it could be that we assume that an S-wave can also be picked, whereas this is not possible. On the other hand, it can also happen that the station is in the nodal plane for the direct P-wave. This situation is implicitly included in the derivation of the P-wave GMPE. The implication is that, according to the model, both a P-wave and S-wave pick cannot be made. In reality, it might still be possible to pick S and include it for the location.

An important choice we made when modelling detection, is the noise percentile. Throughout the report, we assumed the 90th percentile, meaning a noise level that is only exceeded 10% of the time. Therewith, the generated maps show location accuracy that can be achieved in quite severe noise conditions. In low or average noise conditions, there would be more detections, leading to a better location.

The somewhat conservative noise-condition choice is partly offset by another factor that have not been included. It has been assumed that all stations in the detection list provide data. In practise, about 10% of the stations has issues that result in no real-time data being available at KNMI, and hence no (automatic) detection and pick being made.

In Section 2.3.3 we found that the average noise reduction with depth is a factor of 26, when sensors are placed at 200 m instead of at the Earth's surface. In Section 2.2.2 we found that P-wave amplitudes at the Earth's surface are, on average, a factor of 4 higher than at 200 m depth. Hence, effectively a factor of 6.5 gain in detection is reached by placing sensors at 200 m depth. This better detectability also resulted in a much larger empirical base of the derived GMPE for motions at 200 m depth: 15,078 vs 7,592 recordings that passed the (automated) QC criteria (Fig. 2.9). Note that these numbers are only valid for the investigated picking frequency band (from 5 to 40 Hz).

6.2 Location PDF

We used a Bayesian framework (Section 3.1) to map uncertainties in travel time (Section 2.6) to uncertainties in location. We considered location using differential P-wave arrival times over a network (P-delays). Also, we considered location using the travel-time difference between first P and S waves at each receiver (P-S delays). We found that a joint inversion with both P- and P-S delays yields smallest location uncertainty (Section 3.3). Both data attributes have complimentary information on the location problem. With sensors at or close to the Earth's surface, P-S delays are more sensitive to source depth, whereas P-delays are more sensitive to source epicenter.

The location probability density function (PDF) can in many cases be approximated well as a multivariate normal distribution (Section 3.2). We expressed the uncertainty in the different directions as the standard deviations of the PDF in these directions. Most of the probability lies within one standard deviation from the most likely source location (i.e., the mode of the PDF). Hence, this standard deviation gives a good measure of the uncertainty to take into account. However, bear in mind that the PDF describes a wide distribution of locations and that also less likely locations remain possible.

The 3D PDF may be very complicated and many parameters are needed to describe its shape. We have simplified its description to 4 parameters only (Fig. 3.1):

- σ_1 standard deviation describing the largest uncertainty in the horizontal plane, at the depth of the source.
- σ_2 standard deviation describing the smallest uncertainty in the horizontal plane, at the depth of the source.

- θ (smallest clockwise) angle of the σ_1 direction with respect to north.
- σ_Z standard deviation describing the depth uncertainty, at the epicenter of the source.

σ_1 and σ_2 are computed at the most likely depth of the source, σ_Z is computed at the most likely epicenter. Note that for an actual event, at least one iteration is needed to obtain these parameters. In the KNMI operational flow, first a solution is found with the Hypocenter method (*Lienert et al.*, 1986) using an average northern Netherlands or southern Netherlands (Lower Rhine Graben) velocity model. This first solution yields an approximate depth and epicenter. Only in a second iteration, an optimized travel-time model is used and location uncertainties are computed. If the second solution is quite different from the first, another iteration is needed.

Depth uncertainty behaves differently from epicentral uncertainty. In general, the depth is less well constrained than the epicenter (e.g., Fig. 4.7). For constraining source depth very well, receivers would be needed both above and below the source, which situation does not occur in the Netherlands seismic network. Because depth uncertainty is larger, also the dynamic range for displaying σ_Z (Appendices B.2&D.2) has been chosen larger than for σ_1 (Appendices B.1&D.1). For constraining the depth, it is most important to have receivers nearby the epicenter, so that the curved part of the travel-time function is sampled well (Section 4.3). For constraining epicenter it is more important to have a good azimuthal distribution of receivers (Section 4.5).

In seismology, different institutes use different ways to communicate uncertainty. With the assumption that a multivariate normal distribution is described, the most-common uncertainty measures can be translated to the standard deviations used in this report:

- The European-Mediterranean Seismological Centre (EMSC) uses the 95% confidence zone when communicating epicentral uncertainty. For earthquakes in their catalogue they list the semi-major and semi-minor axis of the 95% confidence ellipse. This corresponds to 2.45 times the σ_1 and σ_2 values used in this report, respectively (Eq. 3.17).
- The EMSC also communicates depth uncertainty. However, it is not clear how this parameter is defined. We assume that their depth-uncertainty parameter is analogous to their epicentral-uncertainty parameters; half of the depth-location 95% confidence interval. In that case, σ_Z needs to be multiplied by a factor of 1.96 to obtain their uncertainty parameter.
- In the IASPEI Seismic Format (ISF; <http://www.isc.ac.uk/standards/isf>) also epicentral and depth uncertainty are treated separately. An epicentral ellipse is used that describes the 90% confidence area. In ISF, the semi-major and semi-minor axis of that ellipse are reported. In *Wang et al.* (2015) it is derived that for a bi-variate normal distribution 90% of the probabilities lie within 2.1460 times the standard deviation from the mean. Hence, a factor of 2.15 needs to be applied to translate the σ_1 and σ_2 values in this report to the ISF format.
- In the ISF format, the depth error is communicated as the extend of the 90% confidence interval. Following *Wang et al.* (2015), for a (1D) normal distribution, 90% of the probabilities lie within 1.6449 times the standard deviation from the mean. Hence, a factor of 3.29 needs to be applied to translate σ_Z from this report to the ISF format.

We found that the simplified parameterization of the location PDF (in σ_1 , σ_2 , σ_Z and θ) starts to break down when the azimuthal gap is larger than 250° (Section 4.5). For this reason we have also included maps displaying the azimuthal gaps for the various magnitudes (Appendices B.3&D.3). For $\psi \geq 250^\circ$ location uncertainty is likely underestimated with the

values shown on the maps in Appendices B.1&D.1 and B.2&D.2. When $\psi \geq 250^\circ$, it would be important to find an additional constraint on the location, e.g., the polarisation of the direct P-wave.

Generally, ψ (and σ_1 , σ_2 , σ_Z) become smaller with increasing magnitude. The larger the magnitude, the more stations detect the event, the better the azimuthal coverage and hence the smaller the azimuthal gap. This is also reflected on the maps in Appendices B.3&D.3. Note, however, that we built in a pick restriction. Only the 40 nearest picks are included in the location. For that reason, at some places ψ (slightly) increases with magnitude. At such places, ψ does remain quite low; with already 40 picks, chances are very slim that a large ψ occurs.

The maps published in this report show contoured location-uncertainty levels. The levels at which these contours are drawn are somewhat subjective. Moreover, for drawing contours, an interpolation is done between the computed values on a grid of 1x1 km. This interpolation can result in small deviations from the actual computed values. Hence, if local values are needed with high precision, the values at the underlying grid are to be retrieved from *KNMI* (2023b).

The location-uncertainty maps were specifically derived for upper-crustal seismicity. For calibration, events were used that were caused by depleting gas reservoirs. These events have been found to be no pure slip events, but also to contain a negative isotropic component (*Kühn et al.*, 2020; *Dost et al.*, 2020). As a consequence, their P-wave content is somewhat larger than for tectonic events. Given the large variability in ground motions however (Table 2.2) also tectonic and other induced events are probably still quite well described with the derived GMPEs and hence the uncertainty maps can also be used for other upper-crustal events.

We have computed the uncertainty maps for a typical induced-seismicity case with a source depth at 3 km. In Section 4.2 we assessed what the effect would be if the source were to occur at a different depth. For larger depths, the epicentral uncertainty would be larger, whereas the depth uncertainty would be similar. For that modelling, however, we did not include that we have less knowledge of the velocities at larger depths, and thus a larger travel-time uncertainty and location uncertainty would result (Section 4.1). Combining the insights from Sections 4.1 and 4.2 both the epicentral and depth uncertainty are expected to grow for depths larger than 3 km. Likewise, the epicentral uncertainty would be somewhat smaller when the source is at, e.g., 2 km depth instead of 3 km. The depth uncertainty somewhat undulates at depths smaller than 3 km due to interaction of the PDF with the free surface (Fig. 4.4).

In this report we made the case that P-S delays need to be included in a joint inversion in order to improve the depth estimation of the source (Fig. 3.2). In the operational workflow for Groningen seismicity, however, location is done using only P-delays. To maintain an induced-earthquake catalogue with a consistent workflow from the nineteen eighties onwards, also for the Groningen region an average North-Netherlands (NN) 1D velocity model is used (*Kraaijpoel and Dost*, 2013). The NN S-wave model is not a good average for the Groningen region. The North-Sea Group (top 800 m) S-wave velocity in the NN model is 1.16 km/s whereas average velocities in Groningen are 0.53 km/s (*van Ginkel et al.*, 2020). Using the P-S data attribute together with the NN model would result in placing sources further from the network than they actually are. Thus, in the first iteration, only P-delays are used and the depth is fixed to 3 km. This is the location solution that is published in the online catalogue (<https://www.knmi.nl/kennis-en-datacentrum/dataset/aardbevingscatalogus>). In following iterations, P-S delays are included together with better local velocity or travel-time models,

like local 1D profiles from the detailed 3D velocity model (Romijn, 2017). The operator of the Groningen gas field (NAM) uses the latter 3D model to locate events in Groningen (Willacy *et al.*, 2019). The use of a 3D model has the potential to further reduce the travel-time residuals and therewith the location uncertainty.

6.3 Knowledge of travel-time uncertainty

The maps generated in this research are conditioned on the travel-time uncertainty model as estimated in Section 2.6. The estimated travel-time uncertainty models are based on deviations from a degree-2 polynomial model. These polynomial models act well to describe mean travel times from a few kilometers epicentral distance onwards. Travel-time picks at these distances are used especially to constrain the epicenter. Hence, the modeled epicentral location uncertainty is likely reasonable. And also the depth uncertainty that is achieved with picks from a few kilometers epicentral distance, is likely reasonable.

For more accurate depth location one needs picks at small epicentral distances. At these distances, the wavefield travels directly upwards from the source, leading to a travel-time curve that is hyperbolic (Section 2.6.3). This hyperbolic part of the travel-time model cannot easily be estimated or checked from the picks. For accurately modelling the arrival times at the nearby stations, a good 1D velocity profile of the source region is needed. We assumed that also in the hyperbolic part of the travel-time curve, similar uncertainty exists as with polynomial part. In reality, there will be regional variation in the quality of the available velocity information, or travel-time curves. The produced maps (near station locations) show likely realistic σ_Z values for regions for which already quite good models exist for P- and S-wave velocities in the top few kilometers of the crust¹, like most parts of north of the Netherlands. The model would have less variability than assumed when locating a source in, e.g., the Zuidwending salt dome, for which the velocity is well known and local stations exist. Hence, in practise a smaller σ_Z could be achieved than indicated, e.g., on Fig. 5.7(a). On the other hand, in regions of, e.g., Limburg where there are still large unknowns in the velocity structure of the upper crust, the σ_Z error (near receivers) would be larger than indicated on the maps in Appendices B.2&D.2.

For the polynomial part of the travel-time model, one can well check whether it is a good average description of the travel-time picks. For the hyperbolic part of the model, one often has too little picks to do this check. This increases the chance that the used local 1D model is a biased model and that hence the actual depth of the hypocenter is an outlier to the modeled depth PDF.

For constraining the source depth, the P-S delay plays an important role (Section 3.2). Hence, the depth uncertainty can be reduced by reducing both the P-wave and S-wave travel-time uncertainties. S-wave travel-times, however, show a larger variability and hence largest gains can be achieved by better modelling S-wave travel times. Generally, local P-wave velocities in the overburden are quite well known from seismic exploration campaigns. Local S-wave velocity is much less well known. For reducing depth uncertainty, it is elementary to improve the local S-wave models, e.g., with tomographic methods or S-wave logging.

¹Note that most of the travel time, and hence also most of the travel-time error, occurs where the wave propagates slowest, which is in the upper crust.

6.4 Magnitude of completeness

The computed magnitude-of-completeness (MoC) map (Fig. 5.10) is probabilistic. In the strict definition, MoC is the minimum magnitude for which the catalogue is complete. There is no such thing as complete, however. Exceptional circumstances may exist (e.g., a large storm) that causes seismic events with magnitudes above the MoC to be undetected. We computed MoC as the lower magnitude that can very likely be detected and located. We used the 90th percentile of the noise distribution to model whether at least 3 detections can be made. When computing the MoC in hindsight, from a catalogue of events in a region, it is thus likely that one finds a slightly higher MoC than indicated on Figs. 5.10 and 5.11.

The MoC is computed for a quite high noise level (P90). This means that when the noise levels are average (P50) or low (P05), events can be detected and located below the MoC threshold (Fig. 5.8).

The MoC-map was computed for a source depth of 3 km. If sources are shallower, the amplitudes in the near field are higher, which would result in a somewhat smaller MoC. On the other hand, for deeper seismicity, the MoC would be somewhat larger than shown on Fig. 5.10. For regions with a small number of seismometers, the MoC would be almost independent of source depth since amplitudes at larger distances are nearly source-depth independent.

The previous versions of the MoC map that were published for the Netherlands (*Kruijer et al.*, 2021; *KNMI*, 2021) were based on different models and assumptions. Whether a detection could be made, was based on an S-wave attenuation relation, in combination with average noise levels that existed over the old seismic network (< 2004) in the Groningen region. The new map is based on newly derived P-wave GMPEs together with station-specific noise distributions. Also it has been determined how much higher, on average, the signal should be than the noise for an automatic pick to materialize (Section 2.4) and hard-rock correction factors have been implemented where relevant (Section 2.5).

The latest MoC map based on old assumptions (*KNMI*, 2021) and the one derived in this report (Fig. 5.10) are similar within the dense networks in Groningen, Twente and North-Holland. In most other places, the new map shows slightly higher values. There are two exceptions. First, the area encompassed by the network in South-Holland shows significantly larger MoC values on the new map, due to very high noise conditions that were found. Second, the southeast of the Netherlands shows smaller MoC values on the new MoC map, due to very low noise levels that were found for stations in south Limburg and across the border.

6.5 Future work

In Section 2.5 average hard-rock correction factors were estimated for sensors at, or near, the Earth's surface and at > 40 m depth. In this study, the same correction factors were used for various sites. In reality, there is likely quite some variability between the hard-rock sites. In a future study, site-specific correction factors could be established. E.g., station WTSB is likely on much weaker rock than HGN and hence a smaller correction factor would need to be used here.

At most sites with a vertical array of sensors, only the sensor with the best signal-to-noise ratio (SNR) is used for detection. In the current operations, no advantage is taken of noise-reduction capabilities of these vertical arrays. E.g., the arrays can be used to increase the SNR of incoming body waves by forming 'enhanced seismograms' (*Ruigrok et al.*, 2016). A topic of future research would be how much gain can be achieved by this processing and how

the MoC and location accuracy maps would improve if the enhanced seismogram were to be implemented. Similarly, also the gain of horizontal phased arrays, like within the LOFAR setup (*Drijkoningen et al.*, 2006) can be studied.

For seven of the broadband stations in the NL network, the sampling frequency has not been optimized for detecting induced seismicity. As a consequence, we could not have a complete picture of the noise distribution in the detection frequency band (Section 2.3.1). When the sampling rate is increased at these stations, the noise levels need to be re-assessed.

It could be assessed whether changing the detection frequency band is favorable for improving the detection capabilities of the network.

In this study, the travel-time uncertainty was found to be distance dependent (Section 2.6.1). This dependence was taken into account by computing an average travel-time uncertainty over the collection of available picks (Section 5.1.1). Instead, the uncertainty pertaining to each individual pick could be used within the covariance matrix (e.g., Eq. 3.14). This likely solves the issue that sometimes the location uncertainty grows if more distant picks are included (Fig. 5.1).

The noise conditions in the Netherlands change over time. The seismic noise levels are expected to increase, with a population still growing and more anthropogenic activities near the Earth’s surface. Especially the size of onshore wind turbines is expected to grow the coming years. This makes it necessary to re-assess the station-specific noise levels in the future. This will be taken into account with the yearly updates, as currently already exist for the years 2021 and 2022.

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A 2021 Station list

Table A.1: The year 2021 list of 200 stations that is used in this study. For details on the selection procedure, see Section 2.1. The locations and sensor types are shown on Fig. 2.2. In the last column it is indicated whether the sensor is situated in hardrock conditions.

Station name	Latitude [deg]	Longitude [deg]	Depth [m]	Sensor	Hardrock
NL.ALK1..HGZ	52.6338	4.7444	0	accelerometer	0
NL.ALK2..HGZ	52.6510	4.7234	0	accelerometer	0
NL.ARCN..BHZ	51.5013	6.1942	0	broadband	0
NL.ASS1..HGZ	52.9778	6.5574	0	accelerometer	0
NL.ASS2..HGZ	53.0143	6.5706	0	accelerometer	0
NL.BAPP..HGZ	53.3148	6.8354	0	accelerometer	0
NL.BER1..HGZ	52.6614	4.7037	0	accelerometer	0
NL.BER2..HGZ	52.6478	4.7150	0	accelerometer	0
NL.BFB2..HGZ	53.1875	6.7655	0	accelerometer	0
NL.BGAR..HGZ	53.3679	6.7136	0	accelerometer	0
NL.BHAR..HGZ	53.2292	6.7090	0	accelerometer	0
NL.BHKS..HGZ	53.2919	6.7850	0	accelerometer	0
NL.BKMZ..HGZ	53.2901	6.3189	0	accelerometer	0
NL.BLIJE..HGZ	53.3603	5.8492	0	accelerometer	0
NL.BLOP..HGZ	53.3339	6.7466	0	accelerometer	0
NL.BMD2..HGZ	53.3533	6.6472	0	accelerometer	0
NL.BOWW..HGZ	53.3405	6.8125	0	accelerometer	0
NL.BSTD.00.HGZ	53.3122	6.6921	0	accelerometer	0
NL.BWIR..HGZ	53.3226	6.7861	0	accelerometer	0
NL.BWSE..HGZ	53.3444	6.7099	0	accelerometer	0
NL.COE2..HGZ	52.6328	6.7057	0	accelerometer	0
NL.COE3..HGZ	52.6704	6.6525	0	accelerometer	0
NL.COE4..HGZ	52.6598	6.7451	0	accelerometer	0
NL.COE..HGZ	52.6606	6.7836	0	accelerometer	0
NL.DBLS..HGZ	52.8391	6.0497	0	accelerometer	0
NL.DBN.01.BHZ	52.1017	5.1767	0	broadband	0
NL.DON..HGZ	53.1005	6.5429	0	accelerometer	0
NL.DR014..HHZ	52.7987	6.1785	200	geophone	0
NL.DR024..HHZ	52.7034	6.5081	200	geophone	0
NL.DR034..HHZ	52.8976	6.3084	200	geophone	0
NL.DRA..HGZ	53.1202	6.1033	0	accelerometer	0
NL.EETW..HGZ	53.3348	6.1054	0	accelerometer	0
NL.ELE..HGZ	52.9566	6.5776	0	accelerometer	0
NL.ENM4.01.HHZ	53.4064	6.4817	200	geophone	0
NL.ENV2.01.HHZ	52.8944	6.6337	100	geophone	0
NL.FDKD..HGZ	52.8483	6.1788	0	accelerometer	0

NL.FR014..HHZ	53.1424	5.8868	200	geophone	0
NL.FR024..HHZ	52.8559	5.9304	200	geophone	0
NL.FR034..HHZ	52.9815	5.7126	200	geophone	0
NL.FR044..HHZ	52.8761	6.1187	200	geophone	0
NL.FR054..HHZ	53.0705	5.8685	200	geophone	0
NL.FR064..HHZ	53.0007	6.0237	200	geophone	0
NL.FSW5.01.HHZ	53.2135	7.1195	300	geophone	0
NL.G014..HHZ	53.4425	6.7312	200	geophone	0
NL.G034..HHZ	53.4159	6.5976	200	geophone	0
NL.G044..HHZ	53.4147	6.6747	200	geophone	0
NL.G054..HHZ	53.4184	6.7713	200	geophone	0
NL.G064..HHZ	53.4111	6.8709	200	geophone	0
NL.G074..HHZ	53.3946	6.5458	200	geophone	0
NL.G084..HHZ	53.3954	6.6439	200	geophone	0
NL.G094..HHZ	53.3878	6.7245	200	geophone	0
NL.G104..HHZ	53.3785	6.8044	200	geophone	0
NL.G114..HHZ	53.3669	6.8891	200	geophone	0
NL.G124..HHZ	53.3699	6.5714	200	geophone	0
NL.G134..HHZ	53.3489	6.6686	200	geophone	0
NL.G144..HHZ	53.3586	6.7708	200	geophone	0
NL.G164..HHZ	53.3385	6.5314	200	geophone	0
NL.G174..HHZ	53.3362	6.6363	200	geophone	0
NL.G184..HHZ	53.3306	6.7217	200	geophone	0
NL.G194..HHZ	53.3378	6.8170	200	geophone	0
NL.G204..HHZ	53.3353	6.8952	200	geophone	0
NL.G214..HHZ	53.3134	6.5937	200	geophone	0
NL.G224..HHZ	53.3103	6.6776	200	geophone	0
NL.G234..HHZ	53.3095	6.7675	200	geophone	0
NL.G244..HHZ	53.2900	6.8562	200	geophone	0
NL.G254..HHZ	53.3081	6.9433	200	geophone	0
NL.G264..HHZ	53.2821	6.5458	200	geophone	0
NL.G274..HHZ	53.2832	6.6281	200	geophone	0
NL.G284..HHZ	53.2722	6.7292	200	geophone	0
NL.G294..HHZ	53.2814	6.7870	200	geophone	0
NL.G304..HHZ	53.2792	6.8951	200	geophone	0
NL.G324..HHZ	53.2498	6.5792	200	geophone	0
NL.G334..HHZ	53.2499	6.6708	200	geophone	0
NL.G344..HHZ	53.2525	6.7651	200	geophone	0
NL.G354..HHZ	53.2549	6.8621	200	geophone	0
NL.G364..HHZ	53.2651	6.9245	200	geophone	0
NL.G374..HHZ	53.2740	7.0350	200	geophone	0
NL.G384..HHZ	53.2005	6.5476	200	geophone	0
NL.G394..HHZ	53.2253	6.7241	200	geophone	0
NL.G404..HHZ	53.2250	6.8099	200	geophone	0
NL.G414..HHZ	53.2234	6.8919	200	geophone	0
NL.G424..HHZ	53.2217	6.9886	200	geophone	0
NL.G434..HHZ	53.2319	7.0791	200	geophone	0
NL.G444..HHZ	53.2042	6.6840	200	geophone	0
NL.G454..HHZ	53.1995	6.7645	200	geophone	0
NL.G464..HHZ	53.1978	6.8498	200	geophone	0
NL.G474..HHZ	53.1964	6.9413	200	geophone	0
NL.G484..HHZ	53.1973	7.0290	200	geophone	0
NL.G494..HHZ	53.1765	6.7306	200	geophone	0

NL.G504..HHZ	53.1745	6.8093	200	geophone	0
NL.G514..HHZ	53.1732	6.8848	200	geophone	0
NL.G524..HHZ	53.1699	6.9837	200	geophone	0
NL.G534..HHZ	53.1674	7.0801	200	geophone	0
NL.G544..HHZ	53.1243	6.7580	200	geophone	0
NL.G554..HHZ	53.1433	6.8461	200	geophone	0
NL.G564..HHZ	53.1484	6.9309	200	geophone	0
NL.G574..HHZ	53.1351	7.0222	200	geophone	0
NL.G584..HHZ	53.1103	6.9071	200	geophone	0
NL.G594..HHZ	53.1160	6.9877	200	geophone	0
NL.G604..HHZ	53.2979	7.0752	200	geophone	0
NL.G614..HHZ	53.3819	6.6813	200	geophone	0
NL.G624..HHZ	53.3870	6.8419	200	geophone	0
NL.G634..HHZ	53.3314	6.9109	200	geophone	0
NL.G644..HHZ	53.2207	7.0380	200	geophone	0
NL.G654..HHZ	53.1692	6.9543	200	geophone	0
NL.G664..HHZ	53.1794	6.5852	200	geophone	0
NL.G674..HHZ	53.3213	6.8228	200	geophone	0
NL.G684..HHZ	53.2351	6.6347	200	geophone	0
NL.G694..HHZ	53.2128	6.6438	200	geophone	0
NL.G704..HHZ	53.2407	6.5379	200	geophone	0
NL.G710..HGZ	53.1483	6.6507	0	accelerometer	0
NL.G720..HGZ	53.2462	6.4792	0	accelerometer	0
NL.G730..HGZ	53.2719	6.4301	0	accelerometer	0
NL.G740..HGZ	53.3043	6.3894	0	accelerometer	0
NL.G750..HGZ	53.3078	6.4875	0	accelerometer	0
NL.G760..HGZ	53.3507	6.3901	0	accelerometer	0
NL.G770..HGZ	53.3403	6.3304	0	accelerometer	0
NL.G780..HGZ	53.3648	6.4894	0	accelerometer	0
NL.G790..HGZ	53.3836	6.3319	0	accelerometer	0
NL.G800..HGZ	53.4154	6.4876	0	accelerometer	0
NL.G81B..HHZ	53.3103	6.6776	100	geophone	0
NL.G82B..HHZ	53.2253	6.7241	100	geophone	0
NL.G83B..HHZ	53.1484	6.9309	100	geophone	0
NL.G84B..HHZ	53.2207	7.0380	100	geophone	0
NL.GK014..HHZ	53.2906	6.2471	200	geophone	0
NL.GK024..HHZ	53.3010	6.2927	200	geophone	0
NL.GK034..HHZ	53.2827	6.3056	200	geophone	0
NL.GK044..HHZ	53.2462	6.3261	200	geophone	0
NL.GRP..HGZ	53.1788	5.9779	0	accelerometer	0
NL.HGN.02.BHZ	50.7640	5.9317	4	broadband	1
NL.HLB.11.EPZ	53.1465	7.0036	60	geophone	0
NL.HLG.11.EPZ	53.1491	7.0071	58	geophone	0
NL.HLH.11.EPZ	53.1404	7.0065	60	geophone	0
NL.HLK.11.EPZ	53.1450	7.0096	60	geophone	0
NL.HOEV..HGZ	52.9025	6.1055	0	accelerometer	0
NL.HRDB..HGZ	52.5770	6.6259	0	accelerometer	0
NL.HRKB..BHZ	51.1879	6.1678	0	broadband	0
NL.HWF4.01.HHZ	53.0710	6.3512	200	geophone	0
NL.J01..HHZ	52.7650	5.0545	0	geophone	0
NL.J02.01.HHZ	52.7255	5.1948	0	geophone	0
NL.KLWL..HGZ	51.7755	4.4406	0	accelerometer	0
NL.LUTT..HGZ	52.6133	6.5715	0	accelerometer	0

NL.MAME..BHZ	50.8000	5.9727	0	broadband	1
NL.MAR..HGZ	53.1540	6.2631	0	accelerometer	0
NL.N014..HHZ	53.1185	6.4733	200	geophone	0
NL.N024..HHZ	53.0720	6.4530	200	geophone	0
NL.N034..HHZ	53.1110	6.4045	200	geophone	0
NL.NGA..HGZ	53.1364	6.0257	0	accelerometer	0
NL.NH014..HHZ	52.5657	4.9965	200	geophone	0
NL.NIW4.01.HHZ	53.3548	6.0430	120	geophone	0
NL.NLDW..HGZ	52.0050	4.1957	0	accelerometer	0
NL.ODBK..HGZ	52.9405	6.1339	0	accelerometer	0
NL.ODHP..HGZ	52.8852	6.0249	0	accelerometer	0
NL.OPLO.01.BHZ	51.5888	5.8121	0	broadband	0
NL.OSTM..HGZ	53.1875	6.0857	0	accelerometer	0
NL.OTL2.01.HHZ	52.6289	4.8227	100	geophone	0
NL.PET..HGZ	52.7706	4.6595	0	accelerometer	0
NL.PPB3.01.HHZ	52.6526	4.6700	165	geophone	0
NL.ROLD..BHZ	50.8694	6.0847	0	broadband	0
NL.SL014..HHZ	52.8551	4.8683	200	geophone	0
NL.SPY4.01.HHZ	53.4098	6.7838	120	geophone	0
NL.SRHV..HGZ	53.1833	6.1849	0	accelerometer	0
NL.T024..HHZ	52.4634	6.9787	200	geophone	0
NL.T034..HHZ	52.4297	6.7918	200	geophone	0
NL.T044..HHZ	52.4375	6.8905	200	geophone	0
NL.T054..HHZ	52.3806	6.9057	200	geophone	0
NL.T064..HHZ	52.3432	6.9097	200	geophone	0
NL.T084..HHZ	52.3893	6.9965	200	geophone	0
NL.TERZ.01.BHZ	50.7568	5.9061	250	broadband	1
NL.TWZL..HGZ	53.2325	6.0927	0	accelerometer	0
NL.UTRP..HGZ	53.0969	6.1842	0	accelerometer	0
NL.VBG4.01.HHZ	52.5440	6.6693	192	geophone	0
NL.VHP..HGZ	53.0957	5.9553	0	accelerometer	0
NL.VKB..BHZ	50.8669	5.7847	0	broadband	1
NL.VLD..HGZ	52.8668	6.2157	0	accelerometer	0
NL.VLW3.01.HHZ	52.9682	7.0972	150	geophone	0
NL.VNDM4..HHZ	53.1151	6.8390	200	geophone	0
NL.VNWD..HGZ	53.2368	5.9984	0	accelerometer	0
NL.VPR..HGZ	51.8915	4.3715	0	accelerometer	0
NL.VRS..HGZ	53.0360	6.6046	0	accelerometer	0
NL.WSVN..HGZ	52.8204	6.1983	0	accelerometer	0
NL.WTSB.01.BHZ	51.9663	6.7989	0	broadband	1
NL.ZDL..HGZ	53.0791	6.7026	0	accelerometer	0
NL.ZH014..HHZ	51.9137	4.2881	200	geophone	0
NL.ZH024..HHZ	52.0744	4.4884	200	geophone	0
NL.ZH034..HHZ	51.9783	4.4001	200	geophone	0
NL.ZH044..HHZ	51.8417	4.3626	200	geophone	0
NL.ZH064..HHZ	51.9781	4.2204	200	geophone	0
NL.ZLV4.01.HHZ	53.0921	6.7533	200	geophone	0
NL.ZWE2.00.HHZ	53.1885	5.6045	60	geophone	0
GE.IBBN..HHZ	52.3063	7.7592	1	broadband	1
GR.BUG..HHZ	51.4455	7.2643	0	broadband	1
GR.RAST..EHZ	52.9096	7.7164	0	geophone	0
BE.BOST..HHZ	51.2382	2.9387	304	broadband	1
BE.MEM..HHZ	50.6087	6.0096	0	broadband	1

BE.BEBN..HHZ	50.7970	5.6778	0	broadband	1
BE.OPTB..HHZ	51.1115	5.6360	379	broadband	1
BE.RCHB..HHZ	50.1552	5.2268	43	broadband	1
BE.UCC..HHZ	50.7972	4.3604	141	broadband	1
NL.CHA5.01.HHZ	51.5043	4.9212	141	geophone	0

B 2021 NL-wide maps with location-uncertainty parameters

B.1 Largest-horizontal-location-uncertainty (σ_1) maps

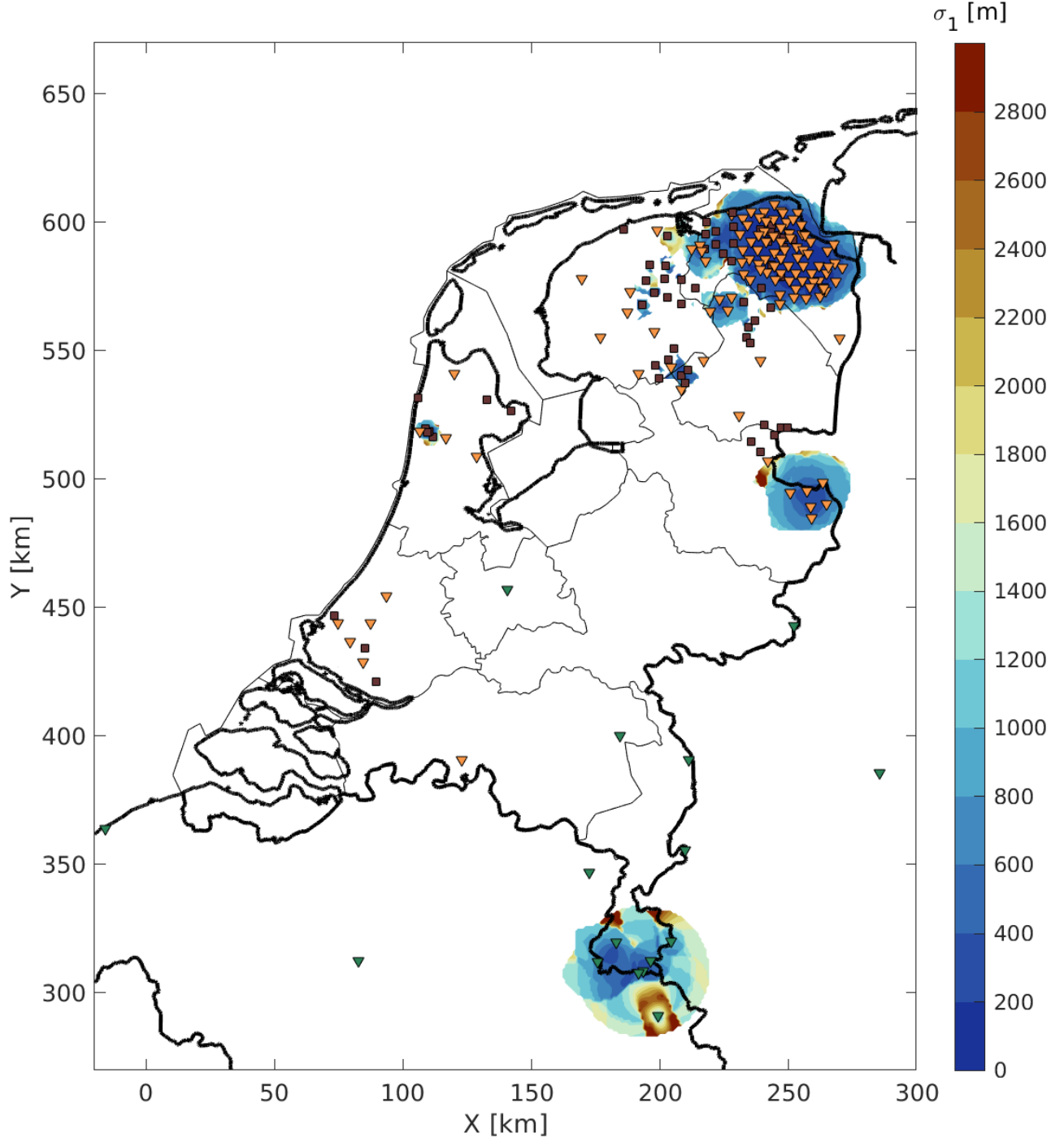


Figure B.1: Spatial distribution of σ_1 in 2021 for $M=0.5$ (color map). White means that events with $M=0.5$ can likely not be detected. Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers. Dark red depicts uncertainties of 2800 m and higher.

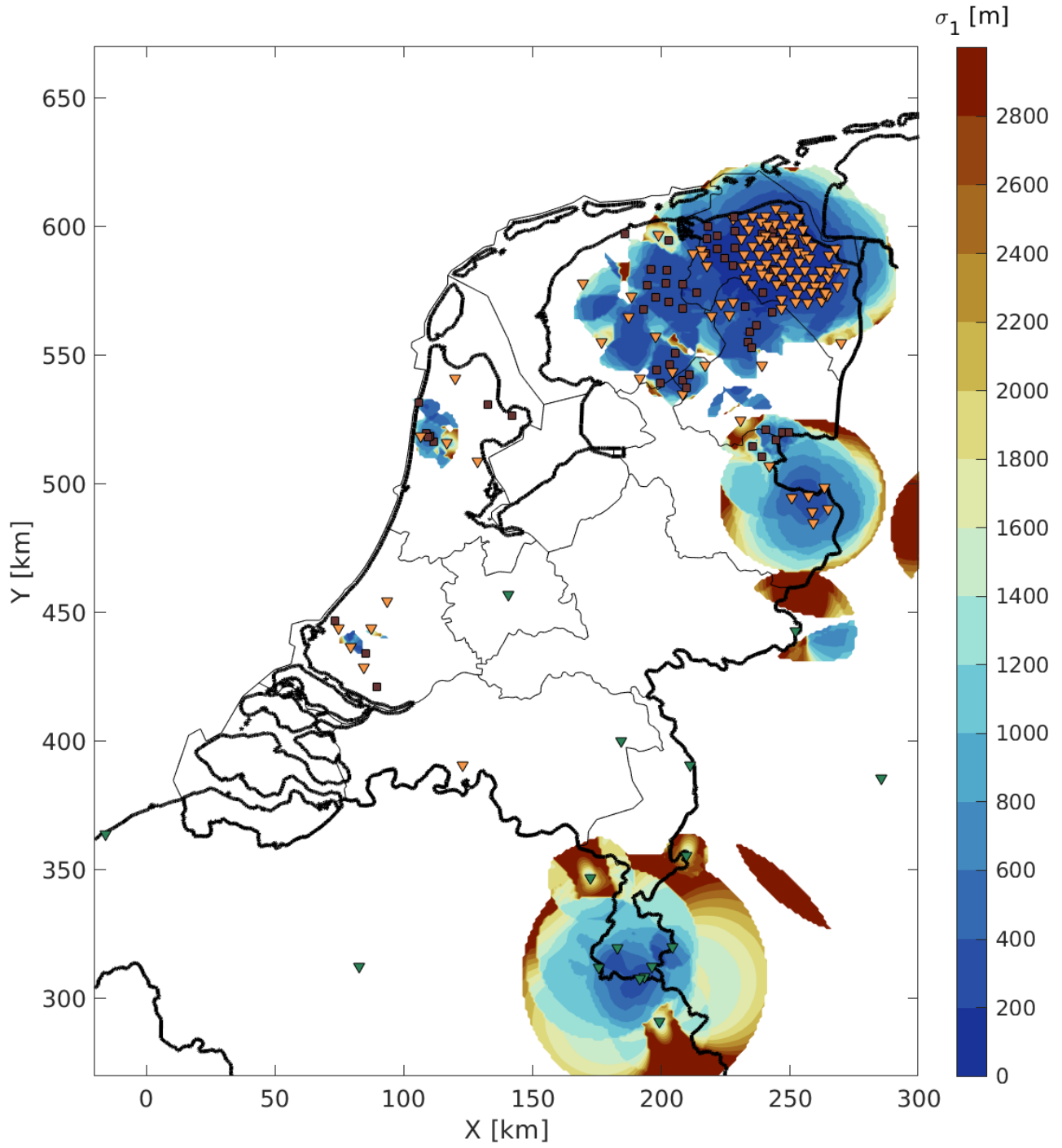


Figure B.2: Spatial distribution of σ_1 in 2021 for $M=1.0$ (color map). White means that events with $M=1.0$ can likely not be detected. Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

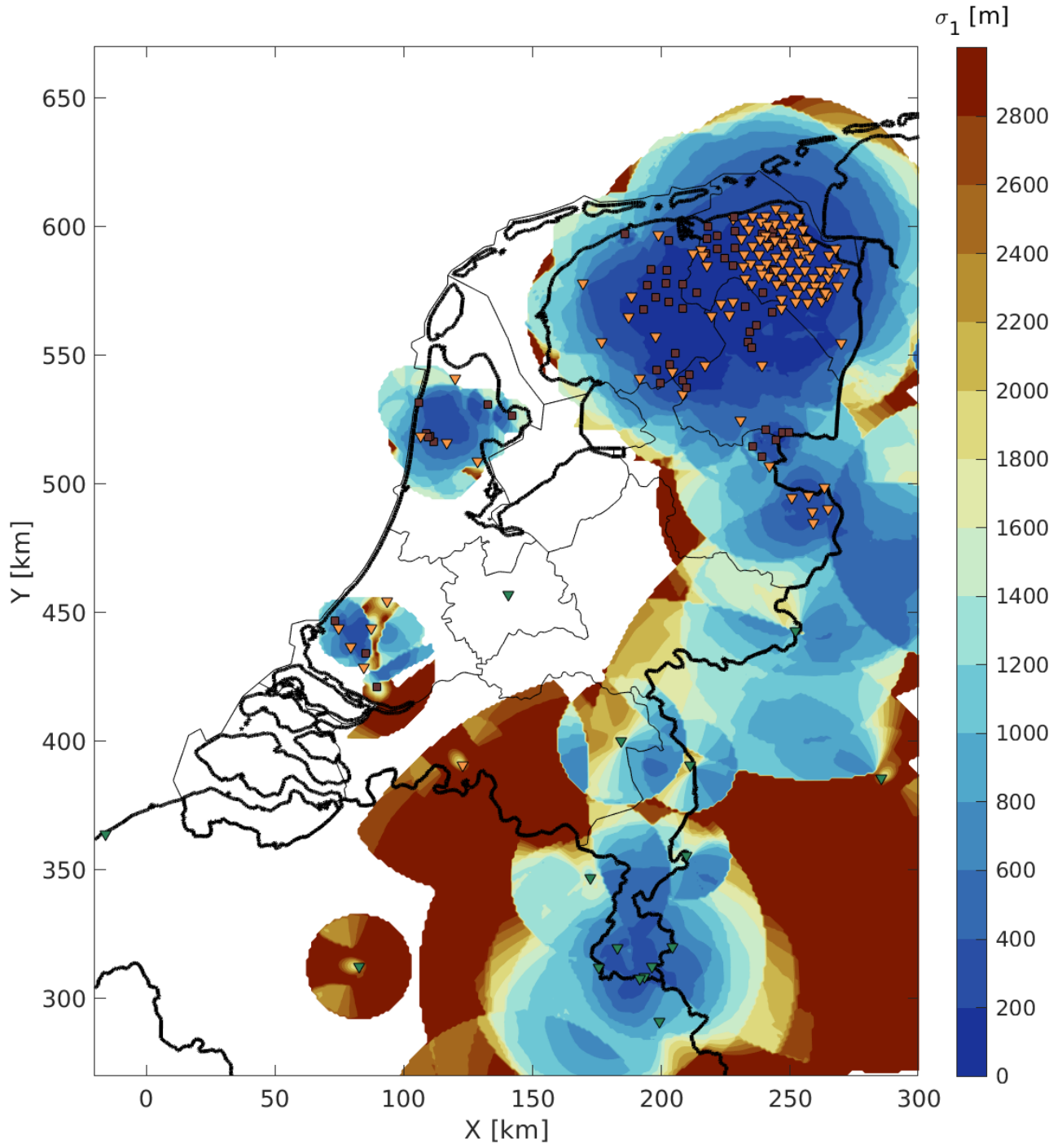


Figure B.3: Spatial distribution of σ_1 in 2021 for $M=1.5$ (color map). White means that events with $M=1.5$ can likely not be detected. Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

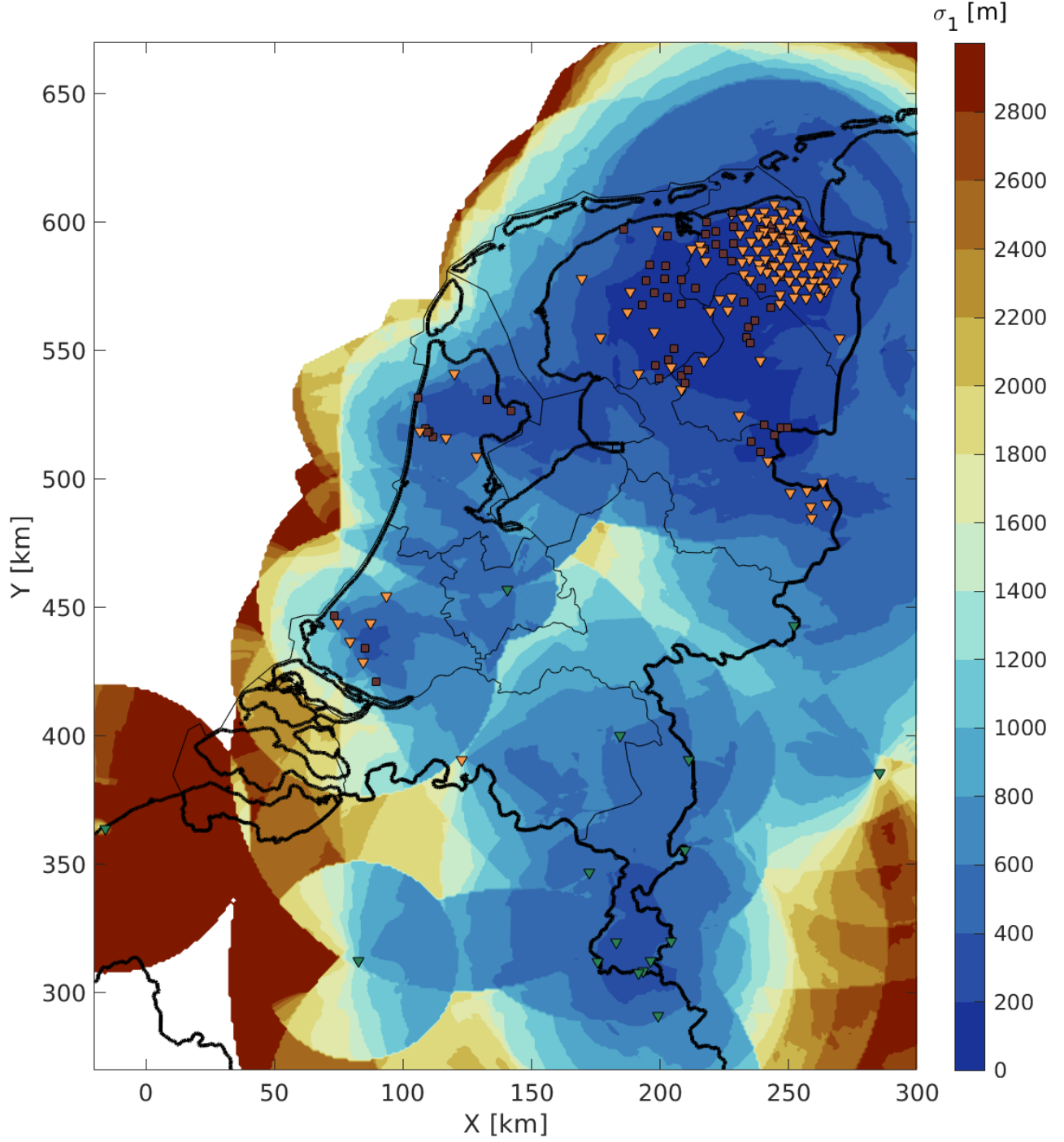


Figure B.4: Spatial distribution of σ_1 in 2021 for $M=2.0$ (color map). White means that events with $M=2.0$ can likely not be detected. Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

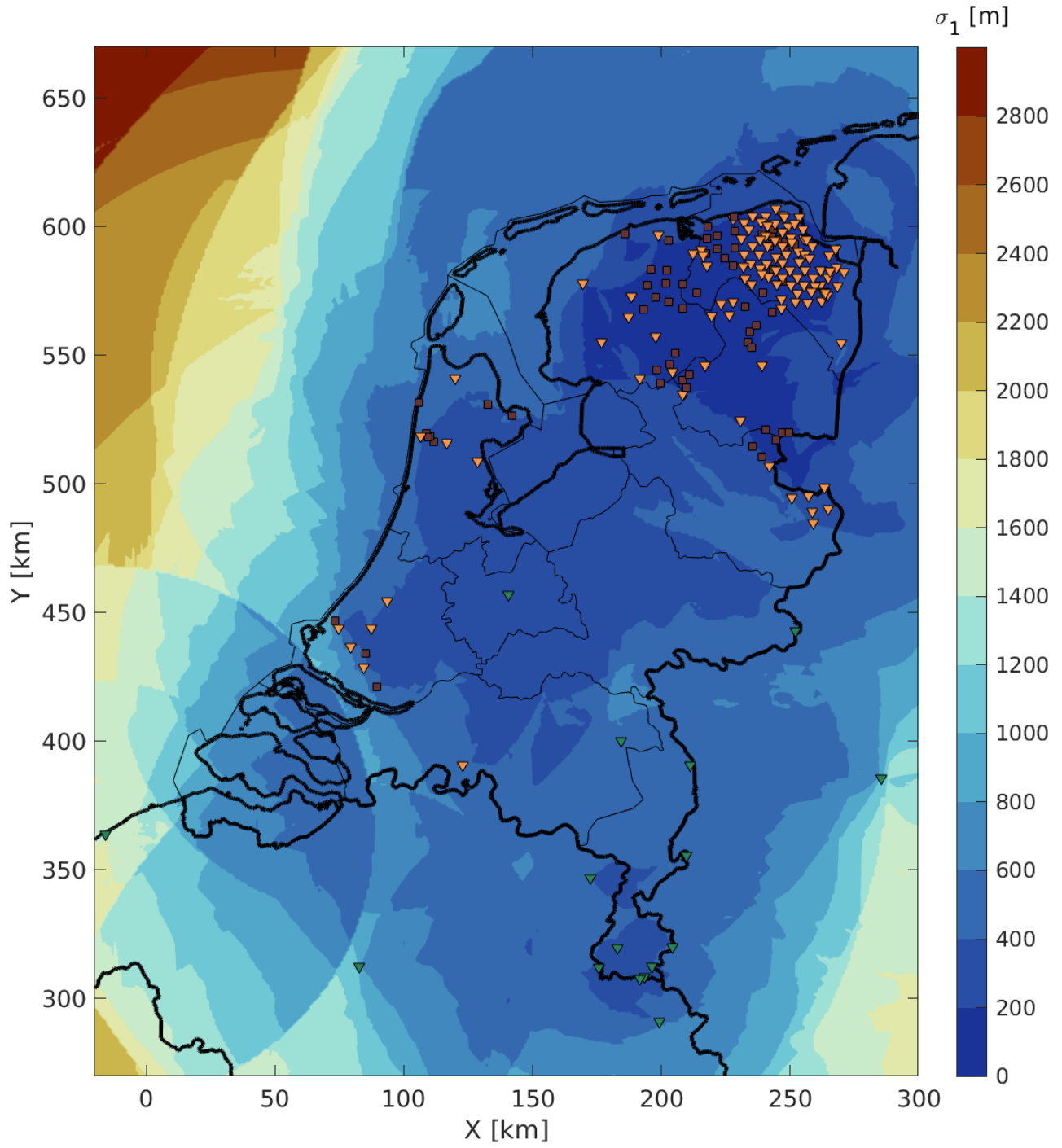


Figure B.5: Spatial distribution of σ_1 in 2021 for $M=2.5$ (color map). White means that events with $M=2.5$ can likely not be detected. Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

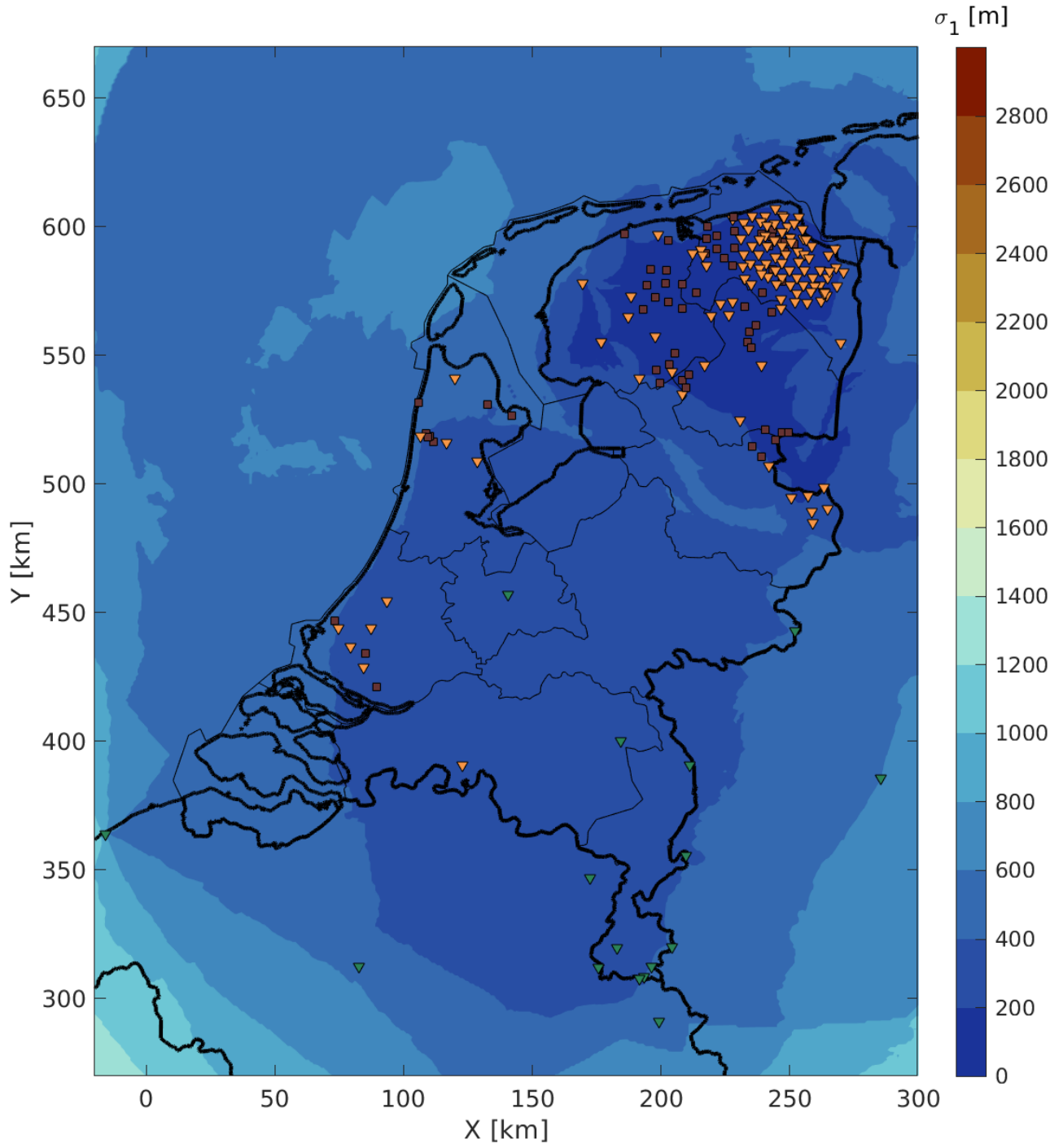


Figure B.6: Spatial distribution of σ_1 in 2021 for $M=3.0$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

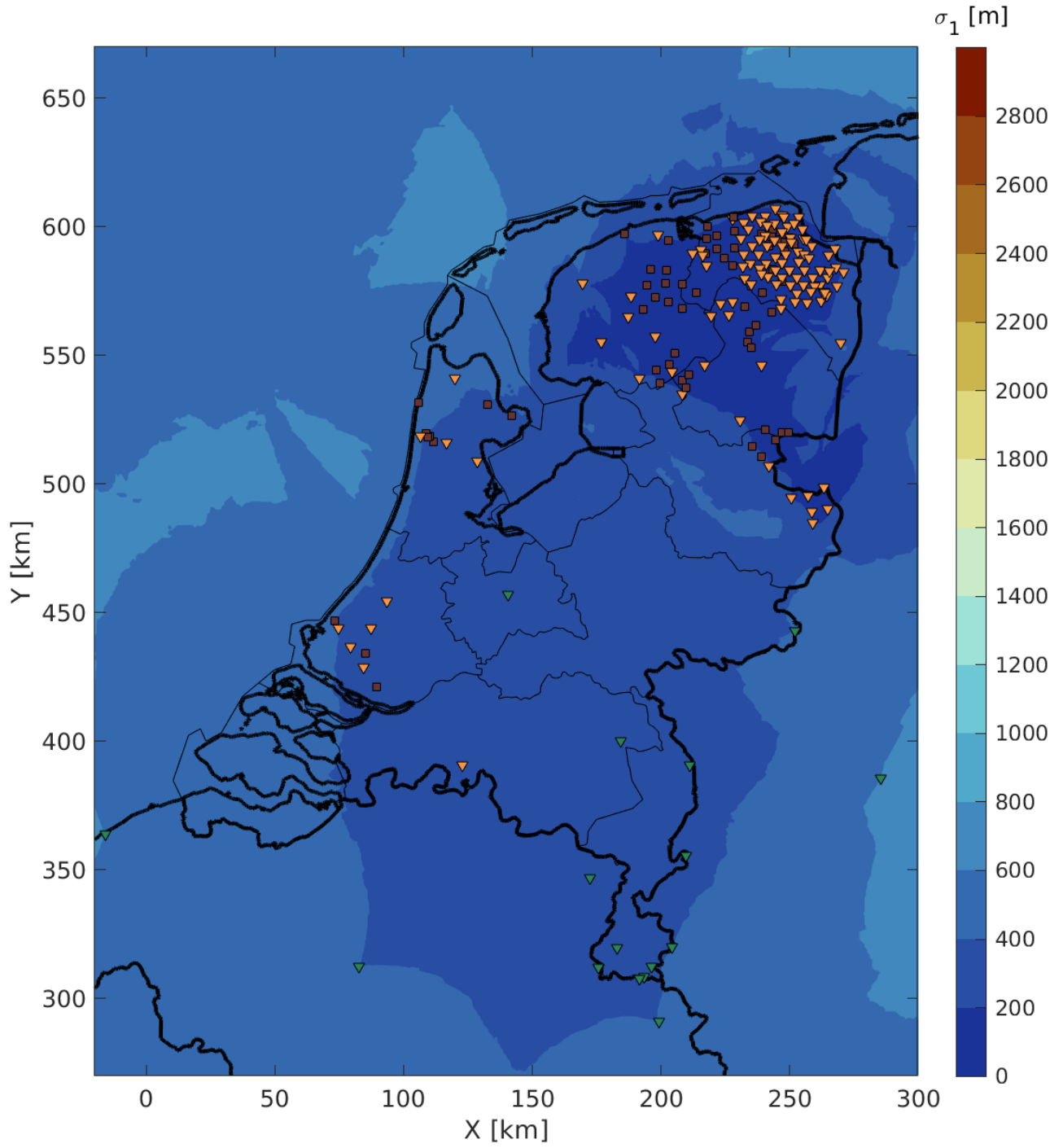


Figure B.7: Spatial distribution of σ_1 in 2021 for $M=3.5$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

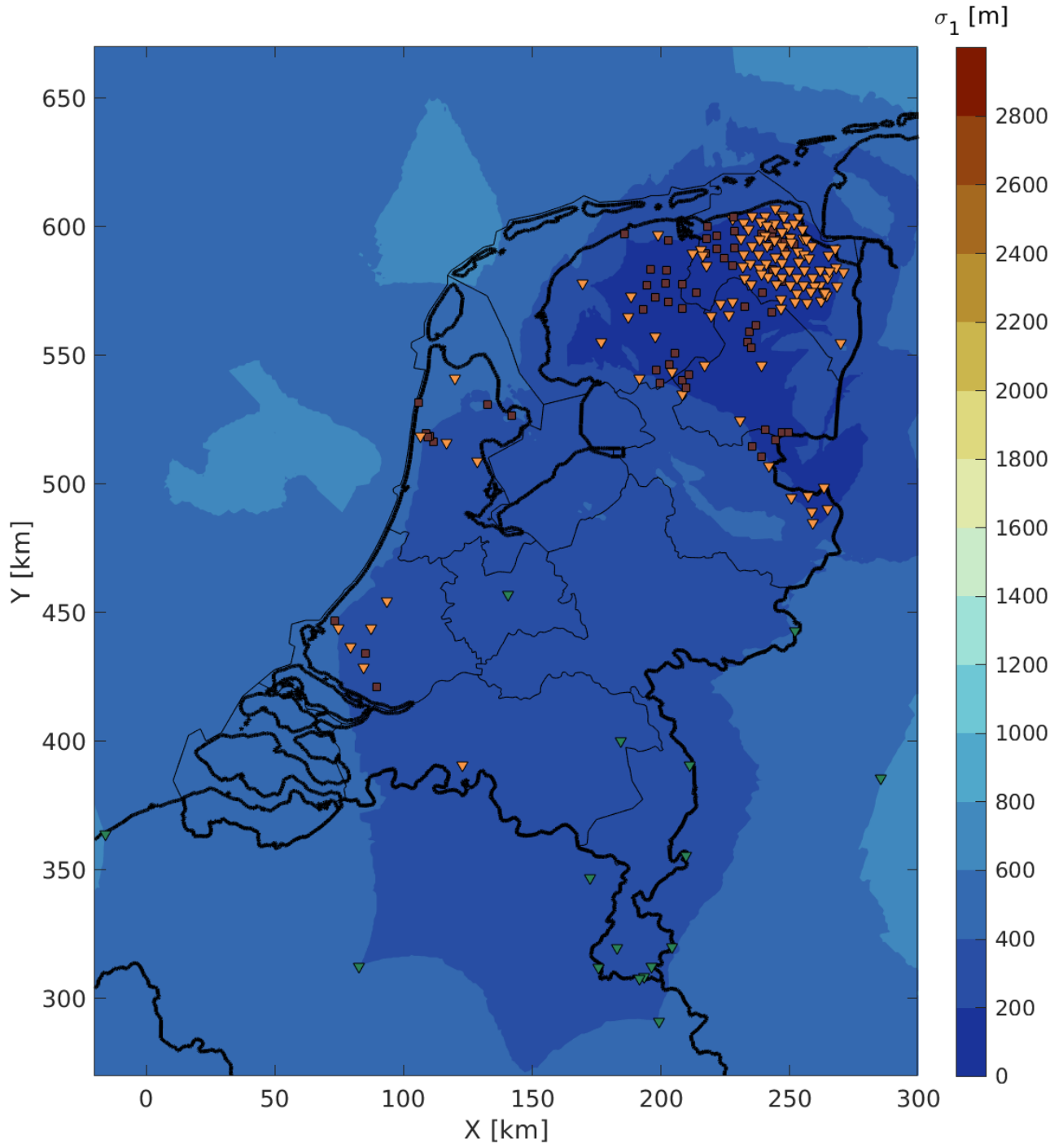


Figure B.8: Spatial distribution of σ_1 in 2021 for $M=4.0$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

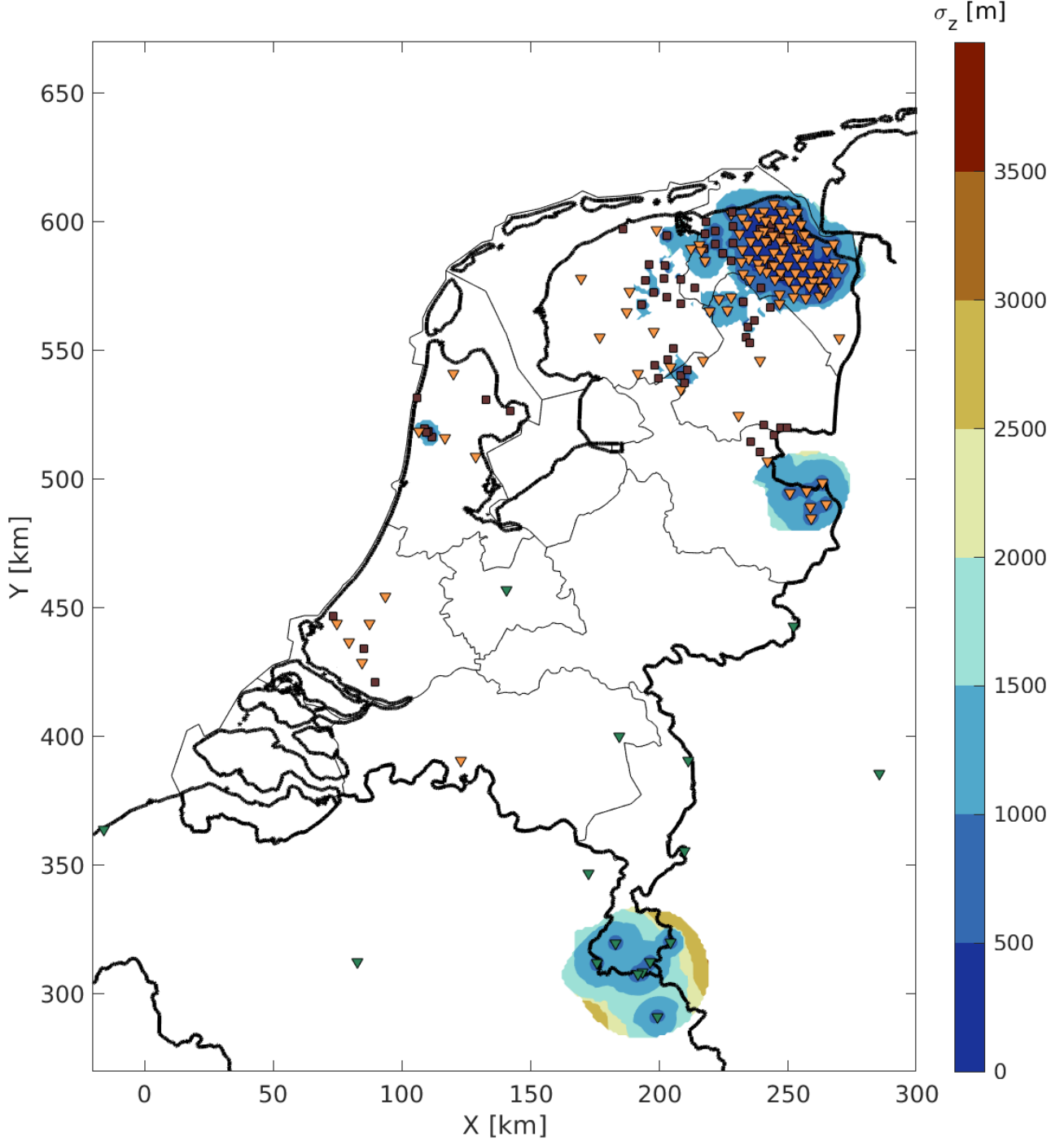
B.2 Vertical-location-uncertainty (σ_Z) maps

Figure B.9: Spatial distribution of σ_Z in 2021 for $M=0.5$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. White means that events with $M=0.5$ can likely not be detected. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers. Dark red depicts uncertainties of 3500 m and higher.

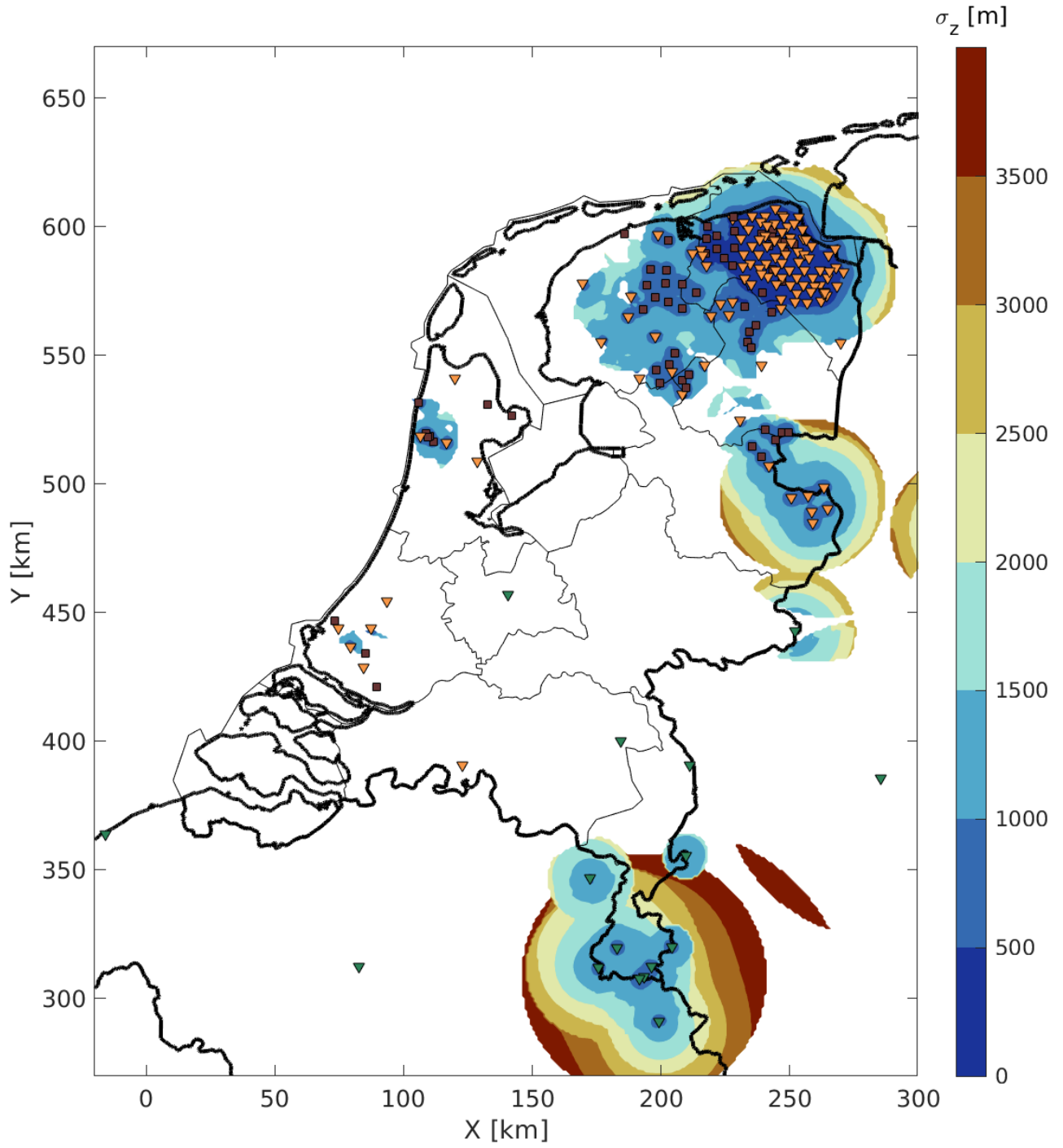


Figure B.10: Spatial distribution of σ_z in 2021 for $M=1.0$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. White means that events with $M=1.0$ can likely not be detected. Uncertainties of 3500 m and higher are shown in red. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

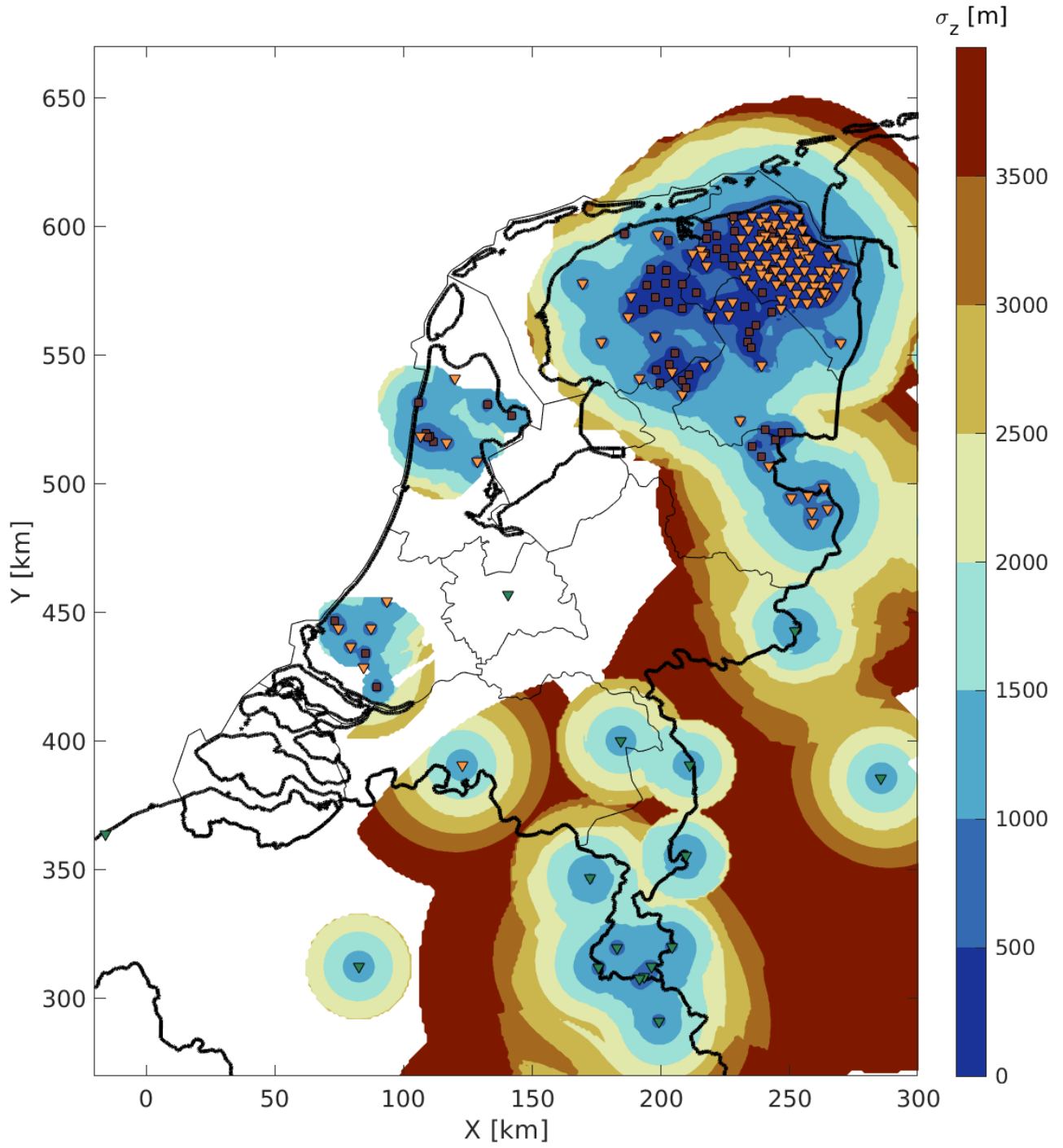


Figure B.11: Spatial distribution of σ_z in 2021 for $M=1.5$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. White means that events with $M=1.5$ can likely not be detected. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

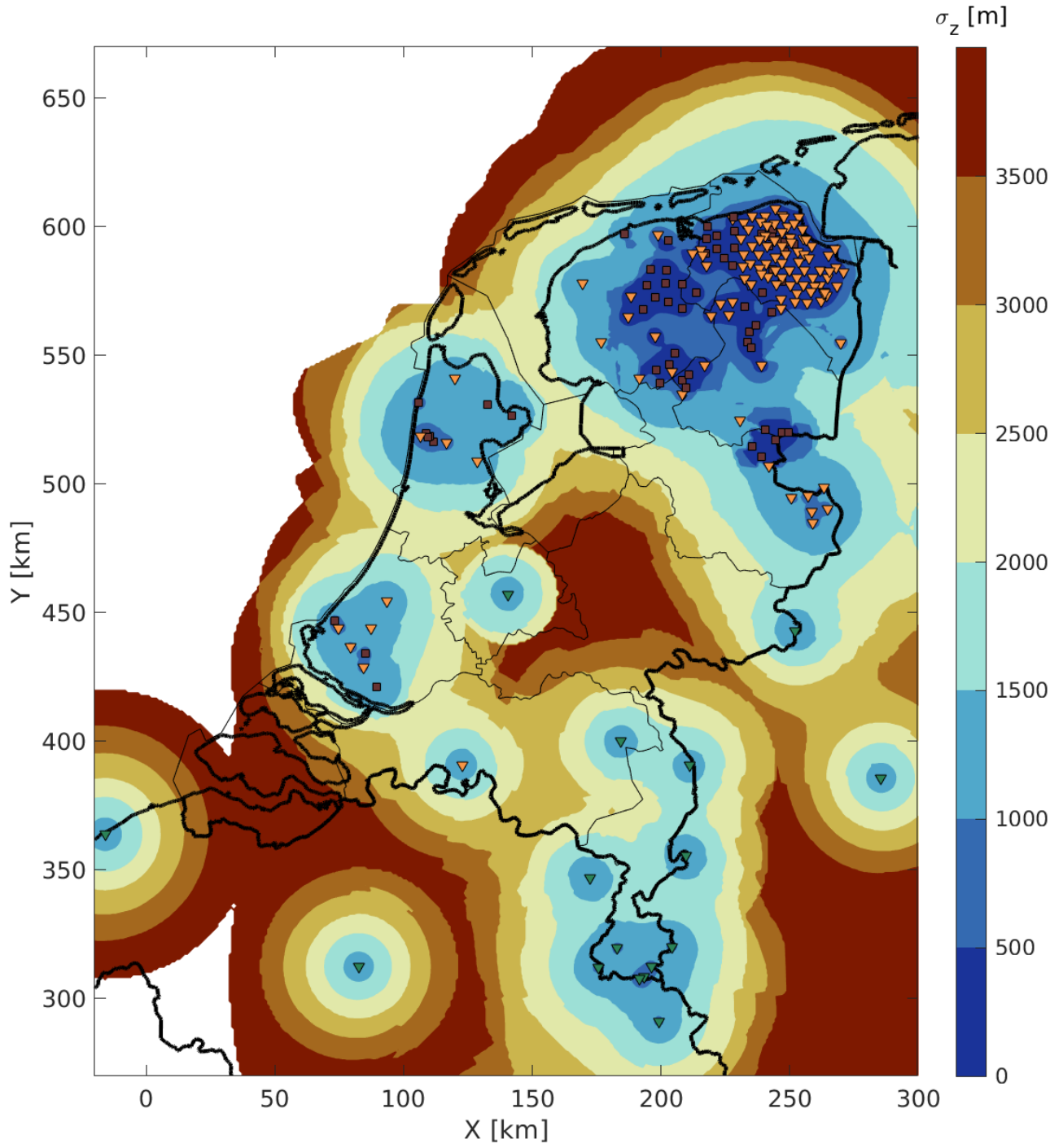


Figure B.12: Spatial distribution of σ_z in 2021 for $M=2.0$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. White means that events with $M=2.0$ can likely not be detected. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

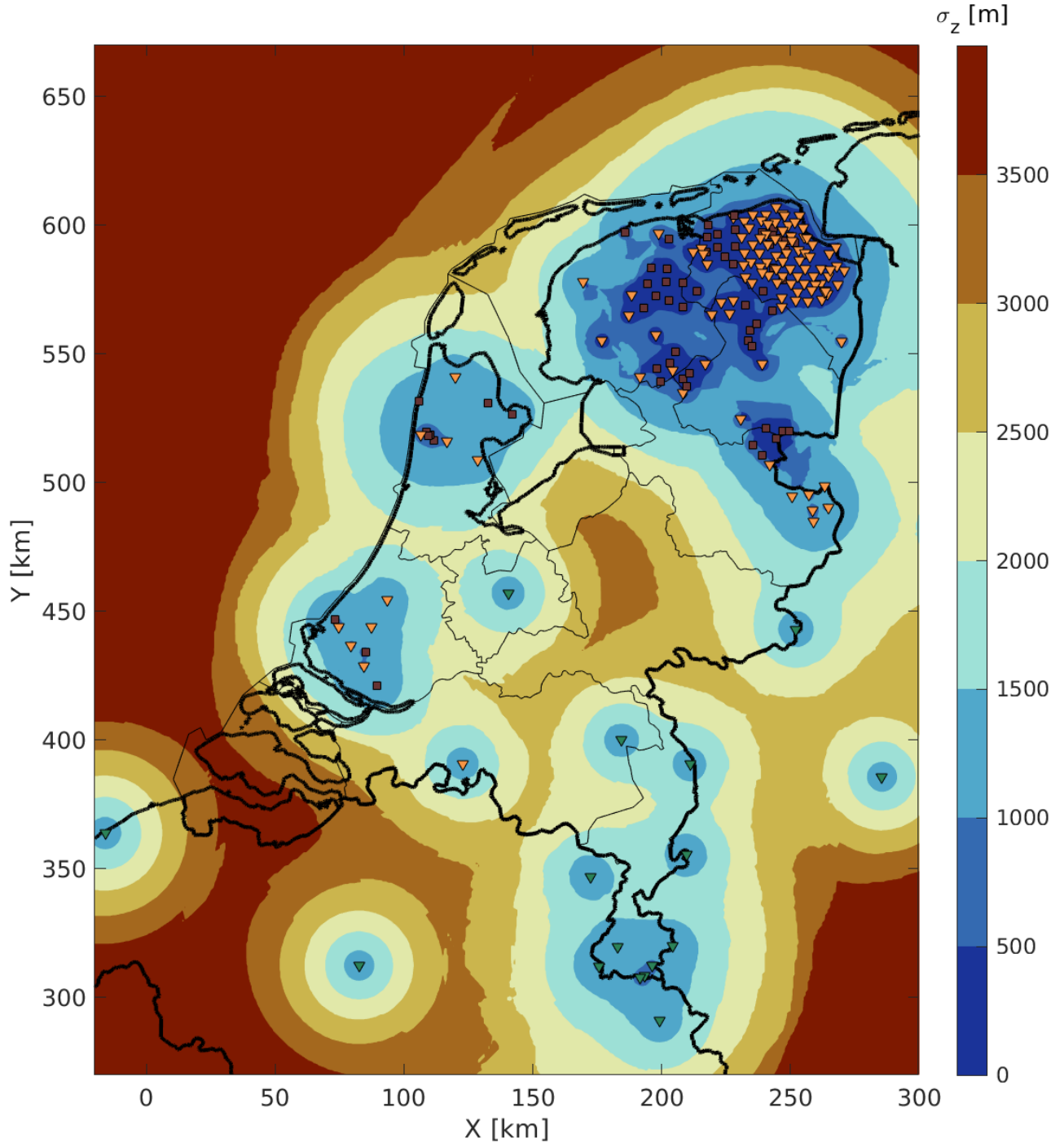


Figure B.13: Spatial distribution of σ_z in 2021 for $M=2.5$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. White means that events with $M=2.5$ can likely not be detected. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

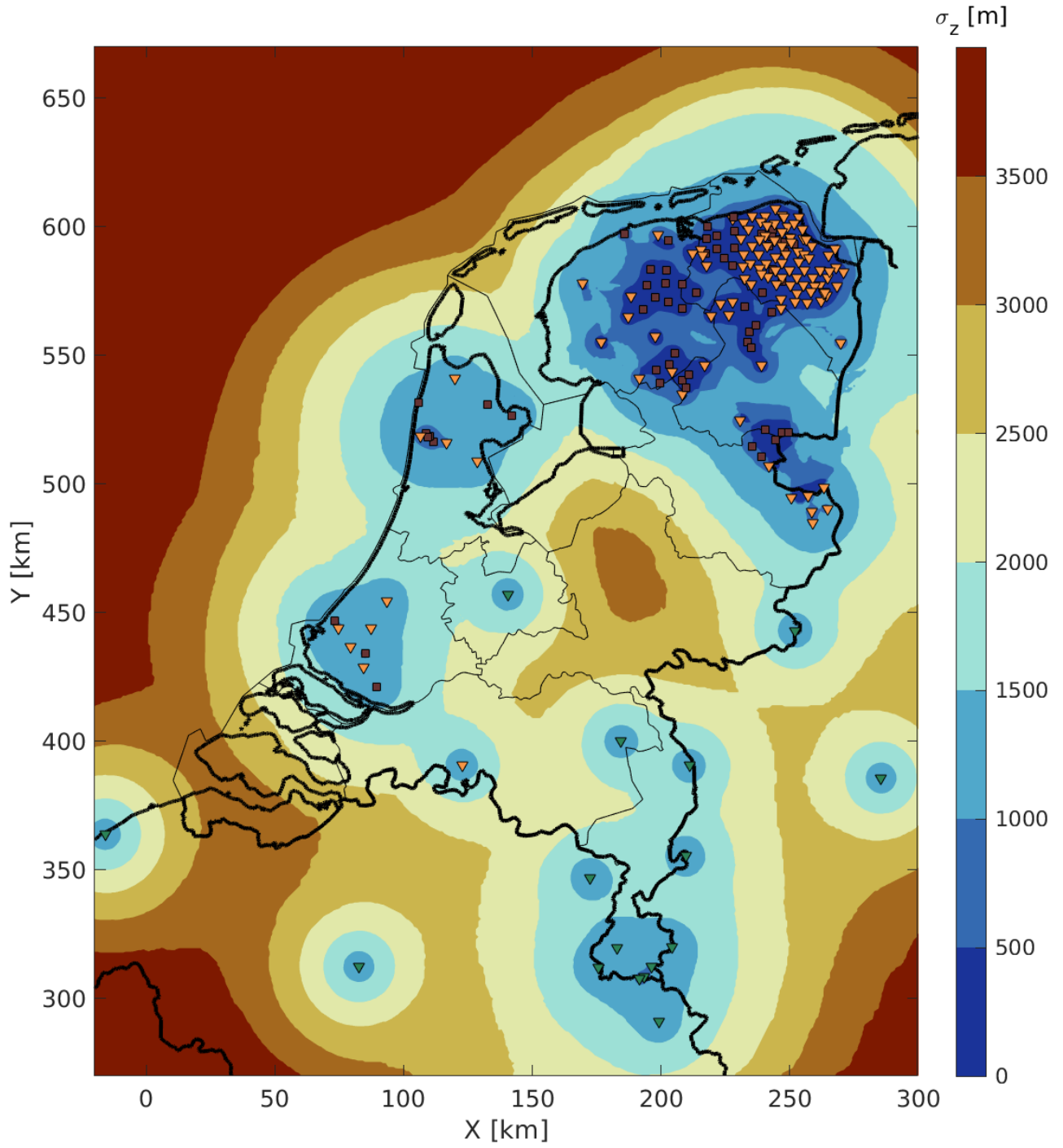


Figure B.14: Spatial distribution of σ_Z in 2021 for $M=3.0$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

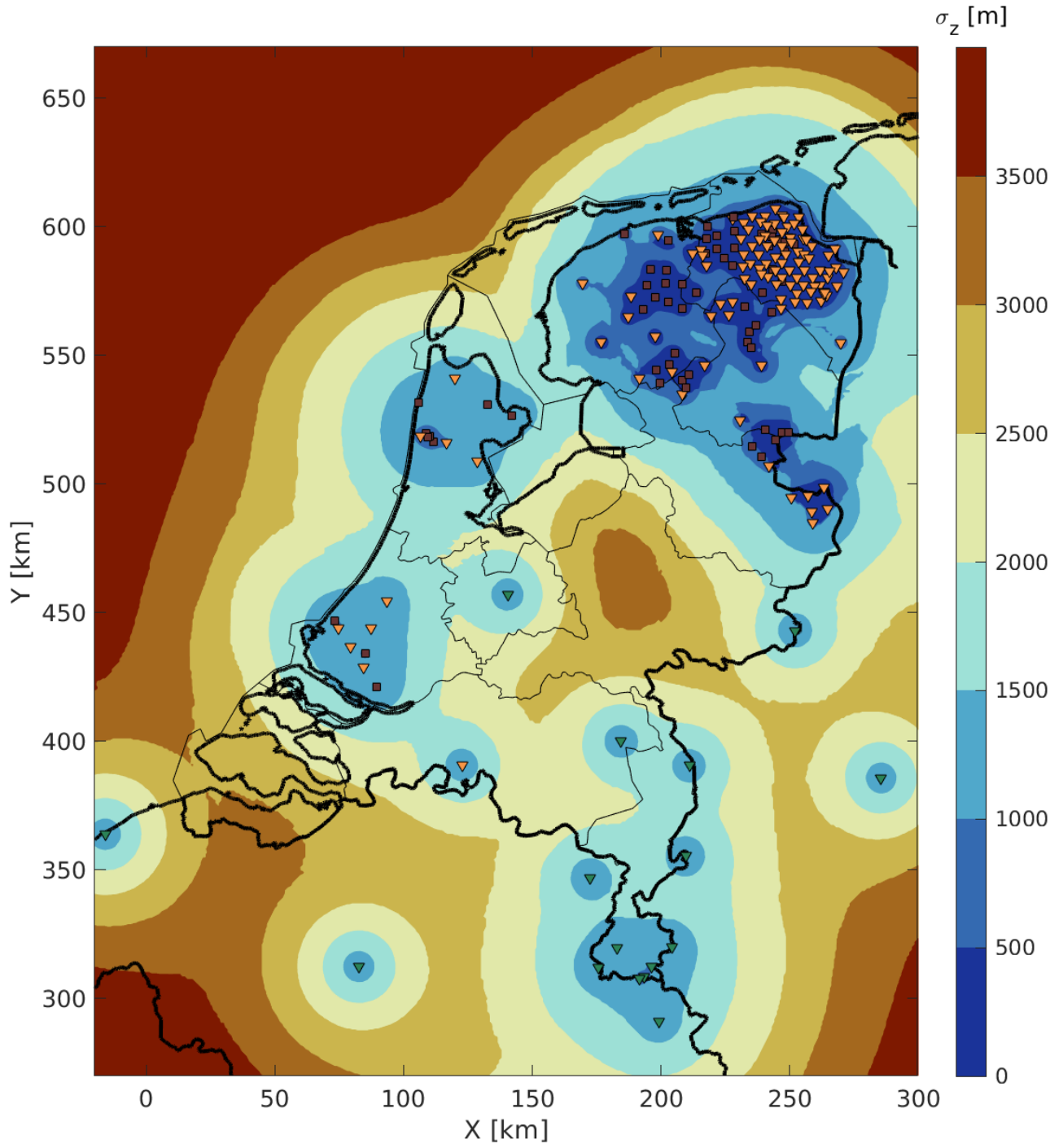


Figure B.15: Spatial distribution of σ_Z in 2021 for $M=3.5$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

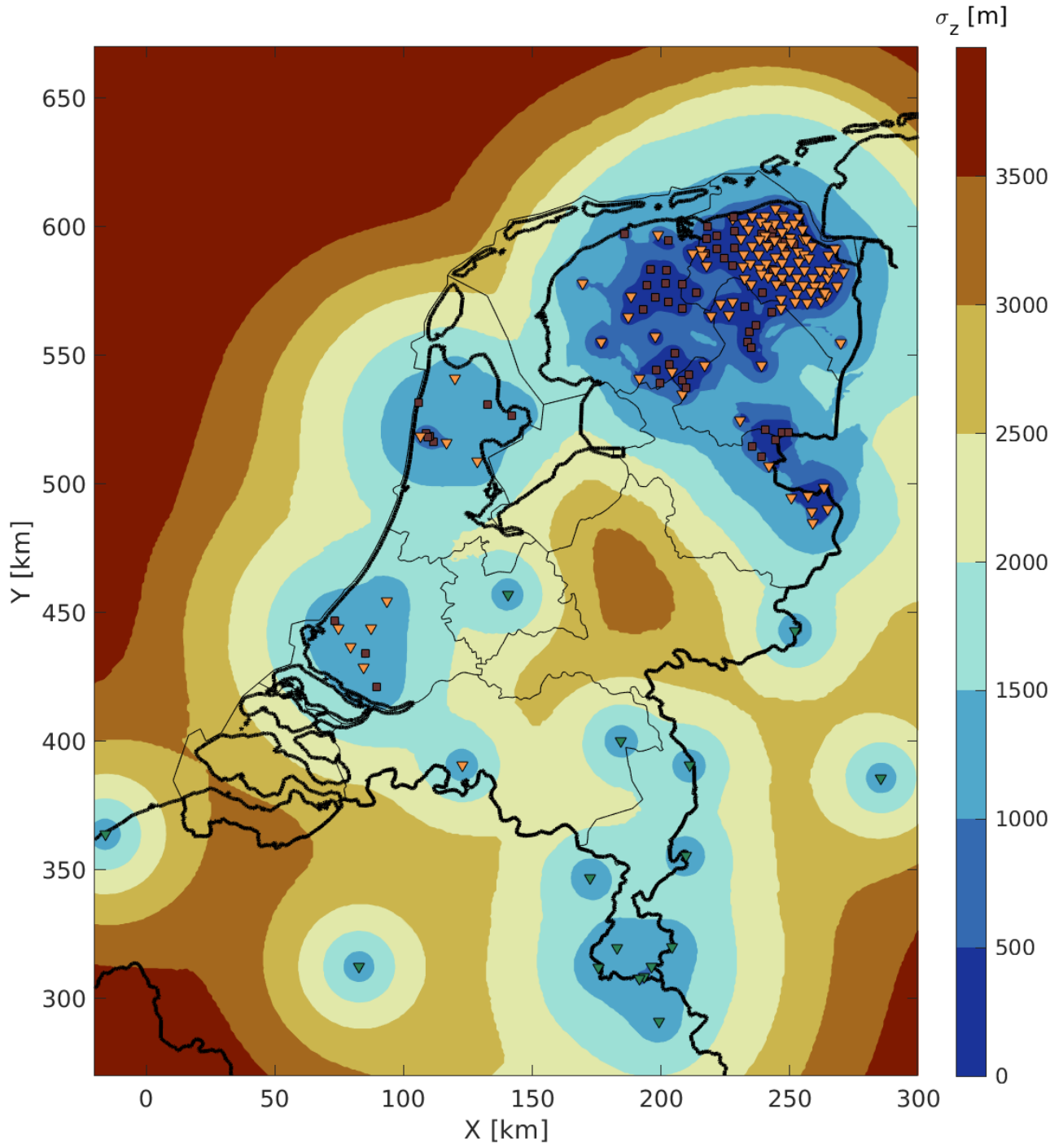


Figure B.16: Spatial distribution of σ_z in 2021 for $M=4.0$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

B.3 Azimuthal-gap (ψ) maps

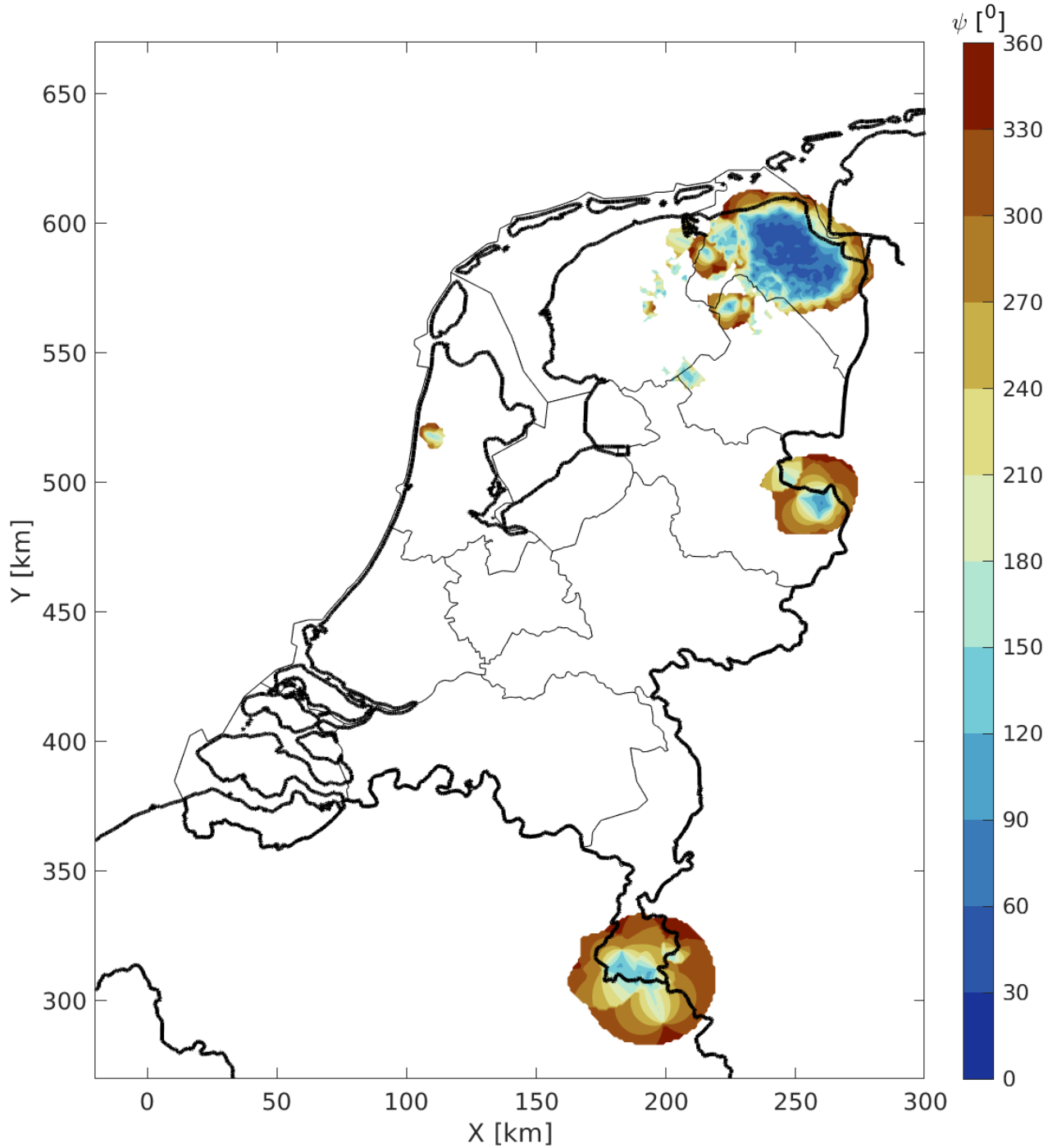


Figure B.17: Spatial distribution of azimuthal gap ψ in 2021 for $M=0.5$ (color map). White means that events with $M=0.5$ can likely not be detected. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

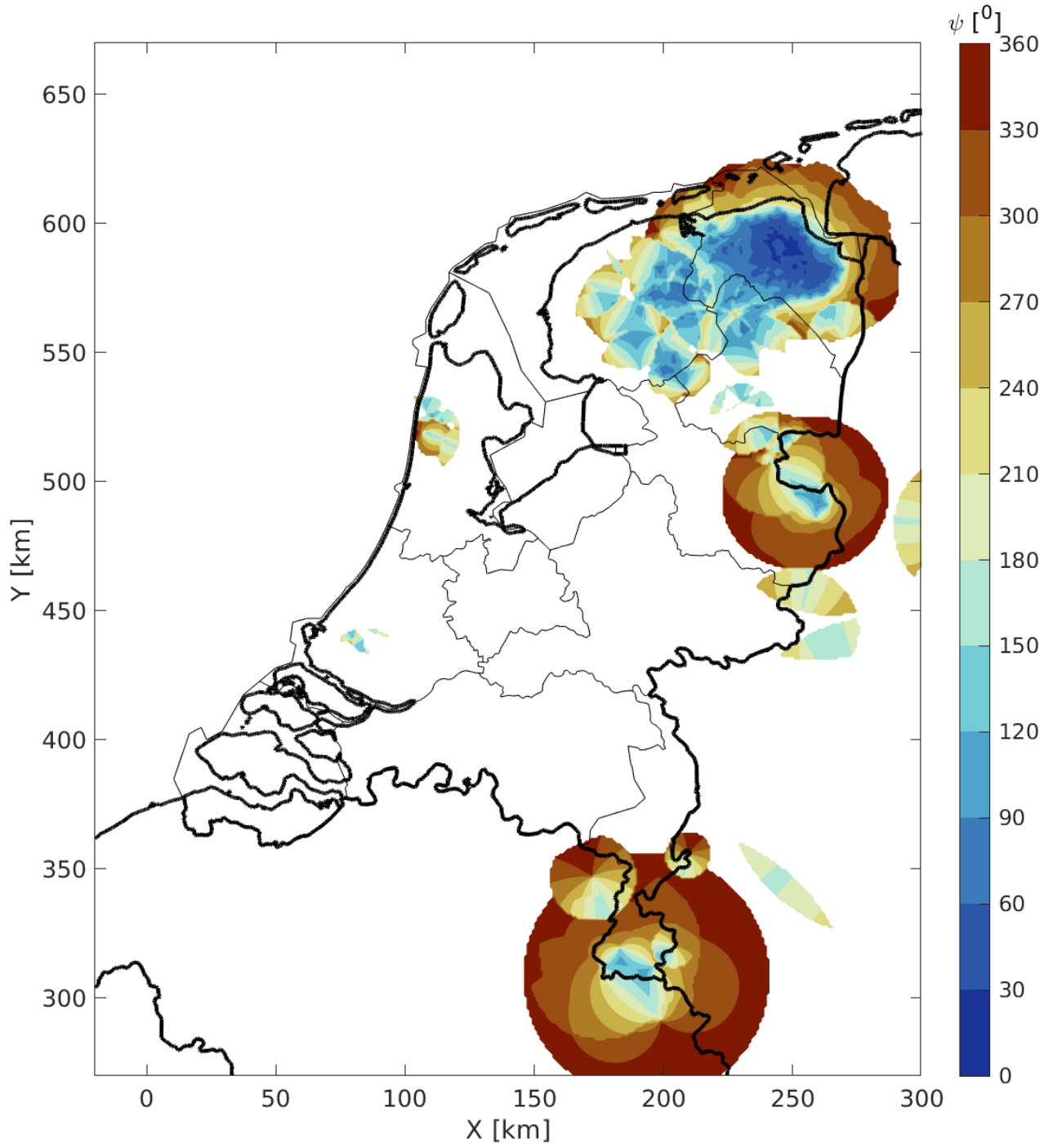


Figure B.18: Spatial distribution of azimuthal gap ψ in 2021 for $M=1.0$ (color map). White means that events with $M=1.0$ can likely not be detected. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

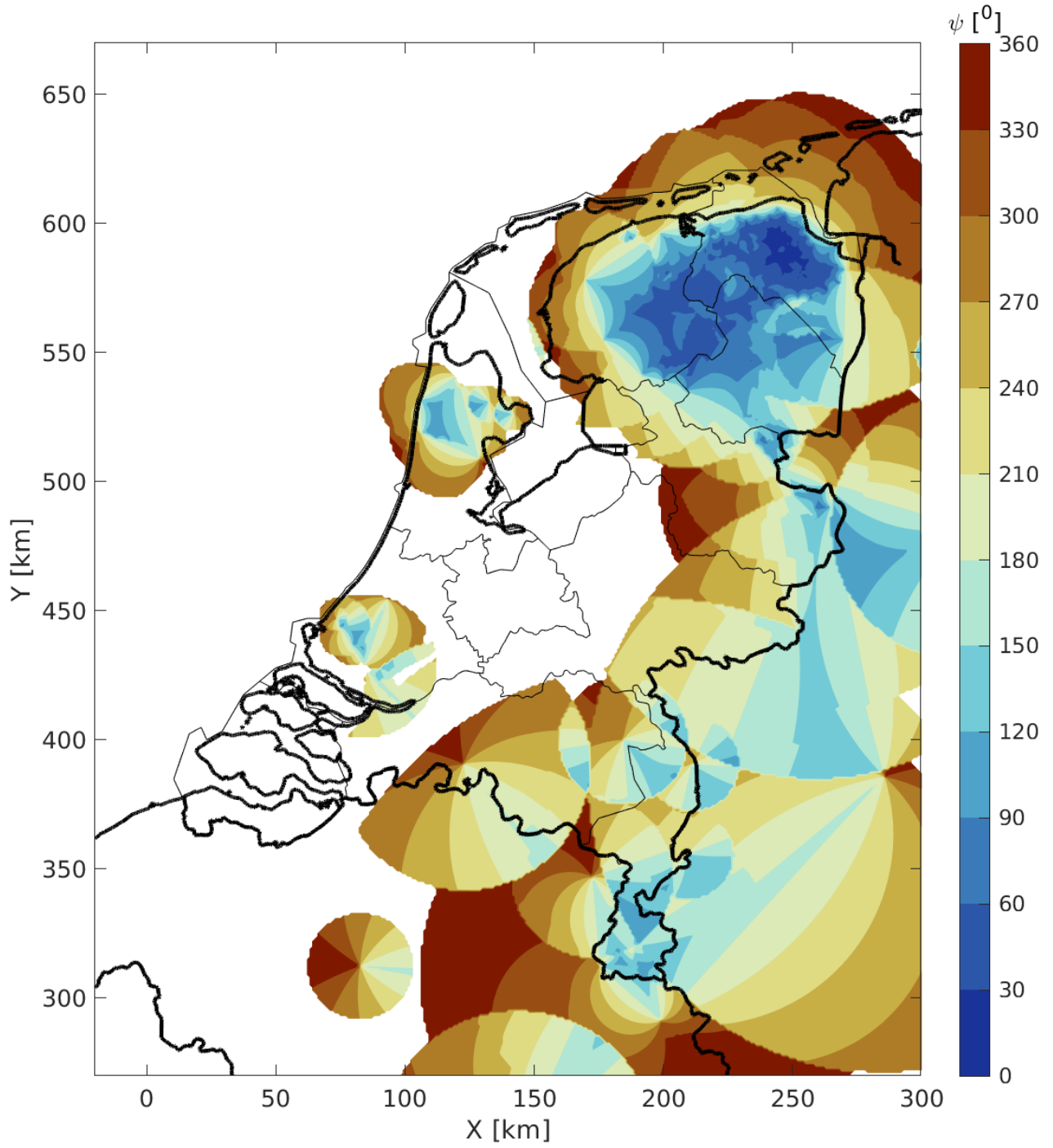


Figure B.19: Spatial distribution of azimuthal gap ψ in 2021 for $M=1.5$ (color map). White means that events with $M=1.5$ can likely not be detected. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

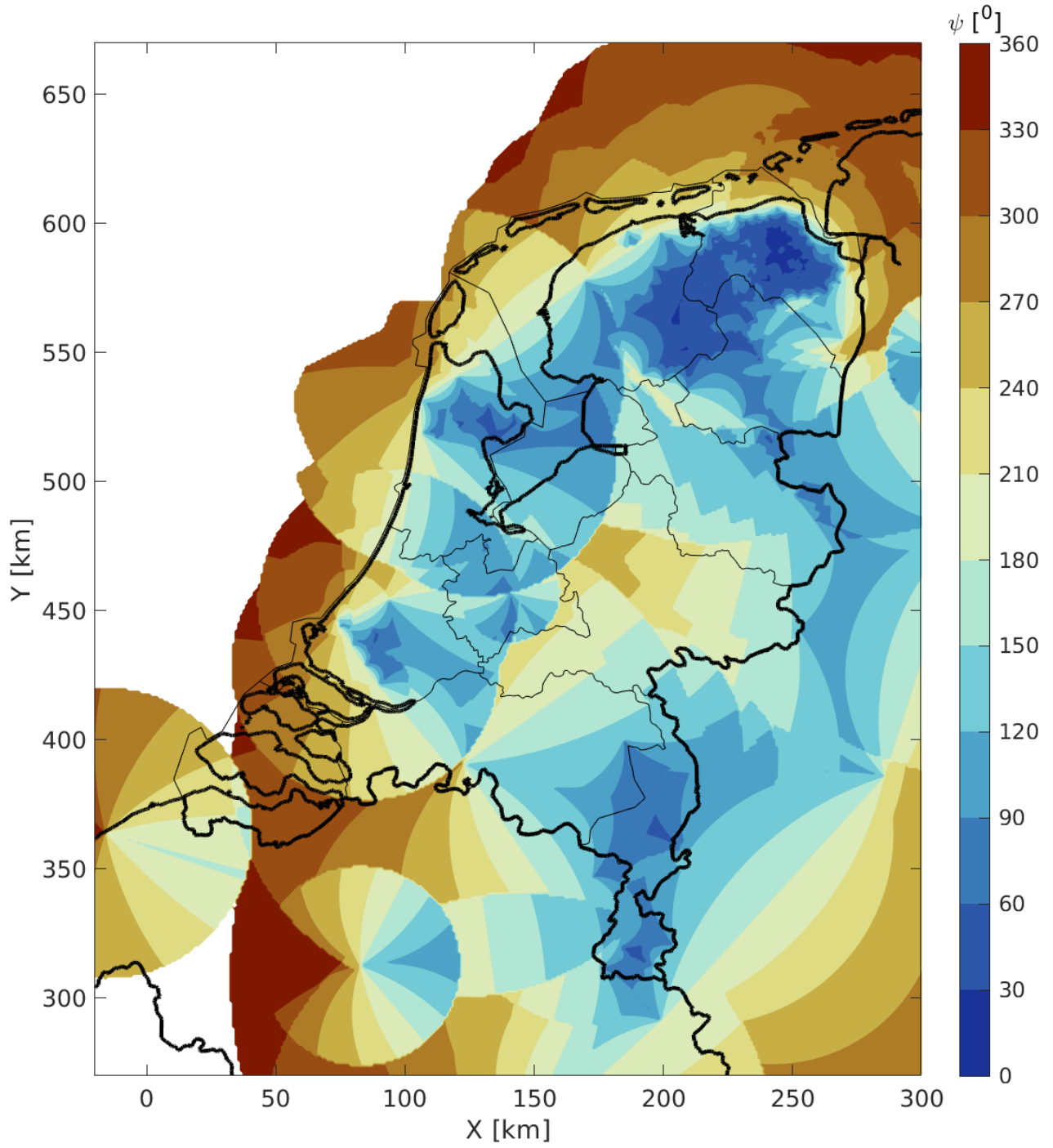


Figure B.20: Spatial distribution of azimuthal gap ψ in 2021 for $M=2.0$ (color map). White means that events with $M=2.0$ can likely not be detected. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

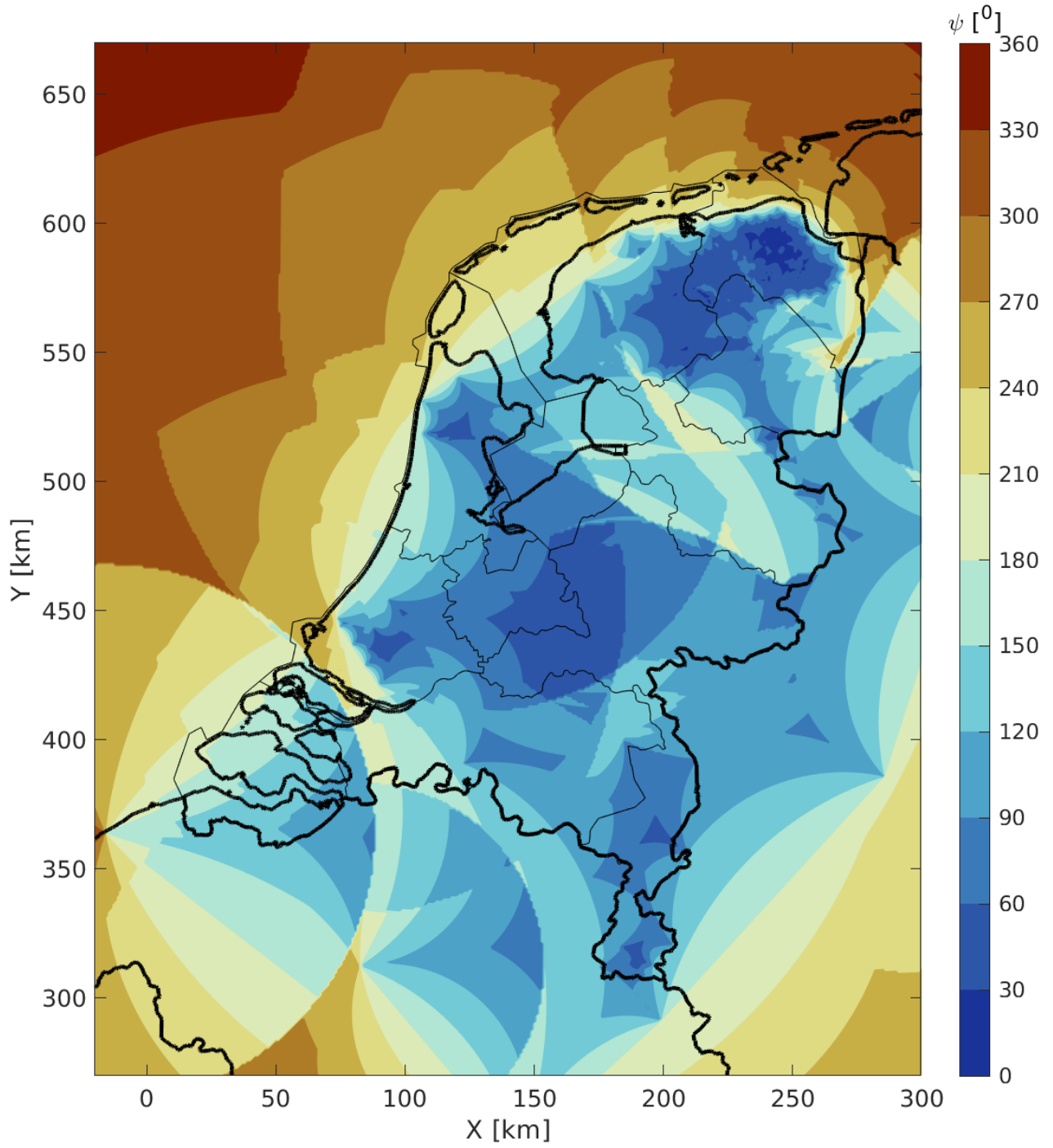


Figure B.21: Spatial distribution of azimuthal gap ψ in 2021 for $M=2.5$ (color map). White means that events with $M=2.5$ can likely not be detected. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

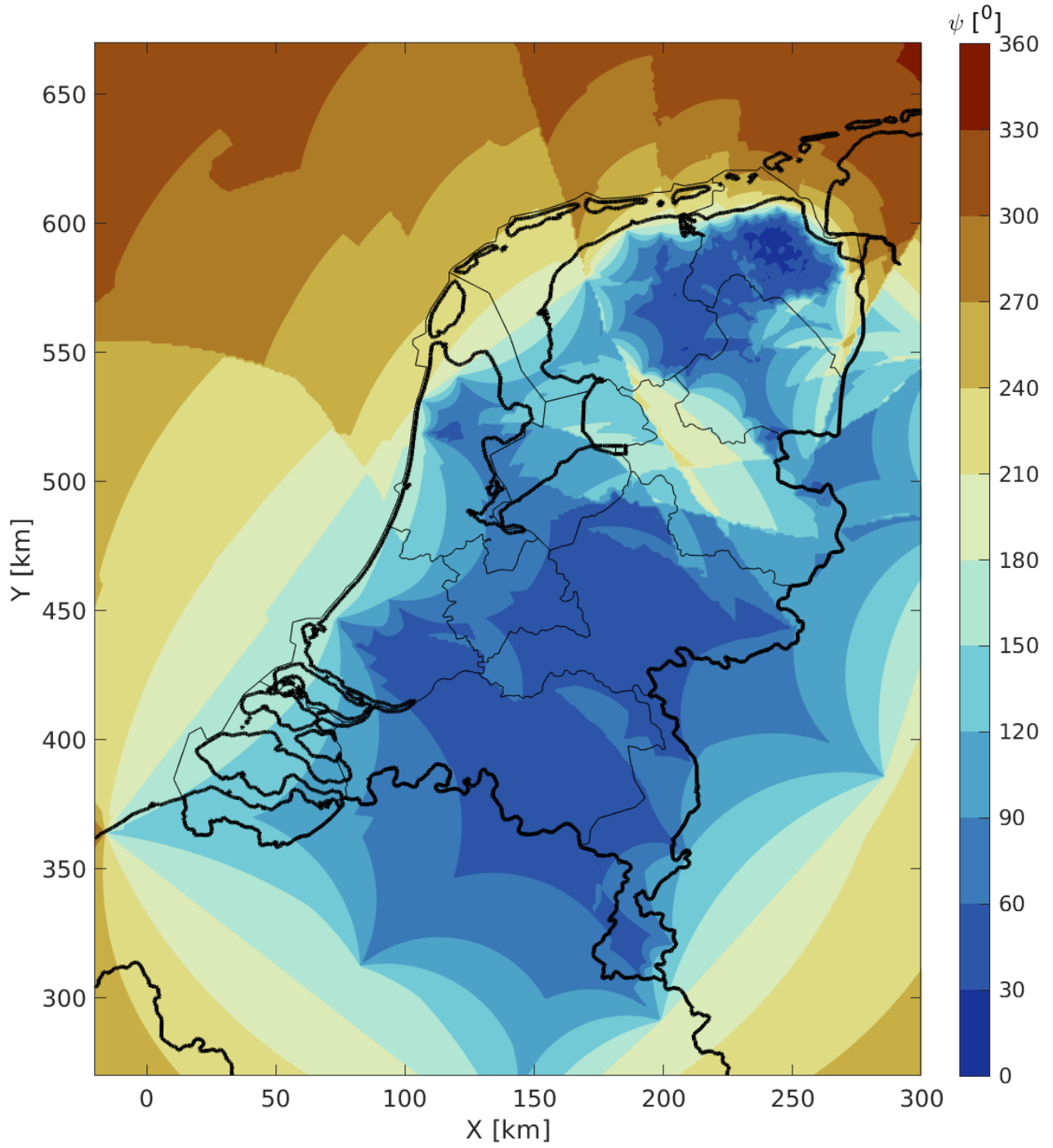


Figure B.22: Spatial distribution of azimuthal gap ψ in 2021 for $M=3.0$ (color map). Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

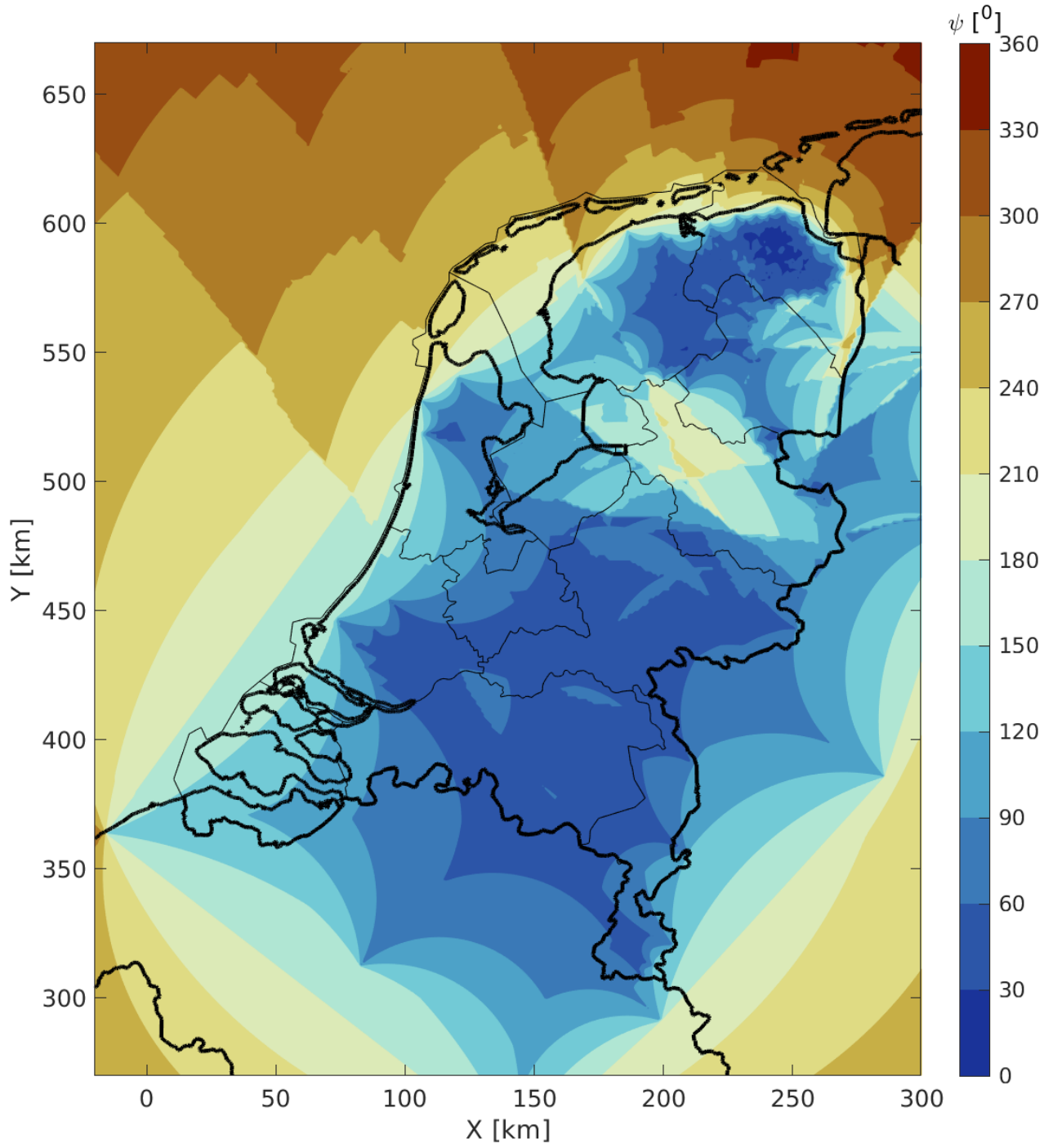


Figure B.23: Spatial distribution of azimuthal gap ψ in 2021 for $M=3.5$ (color map). Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

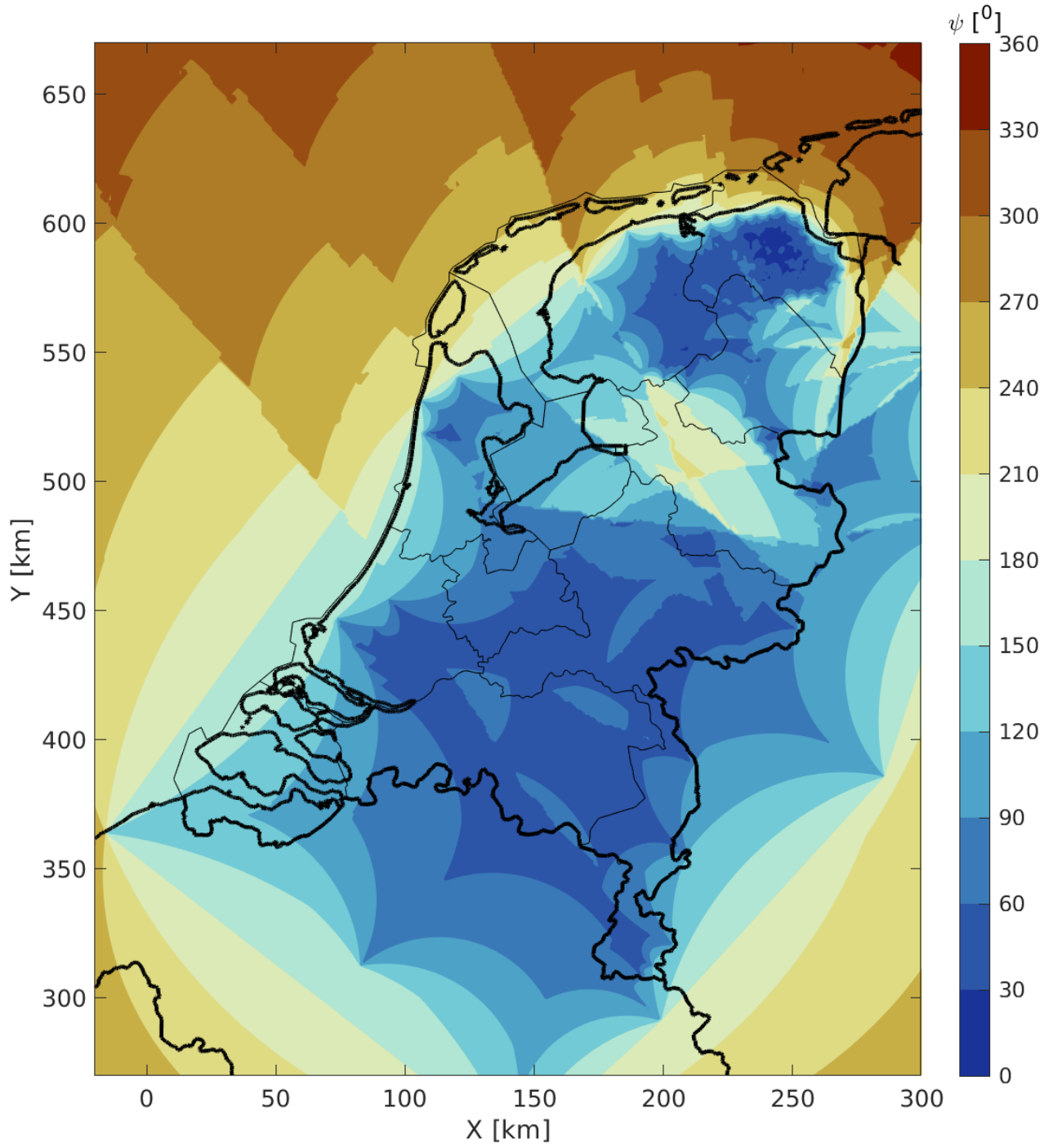


Figure B.24: Spatial distribution of azimuthal gap ψ in 2021 for $M=4.0$ (color map). Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

C 2022 Station list

Table C.1: The year 2022 list of 217 stations. For details on the selection procedure, see Sections 2.1 and 5.4.3. In the last column it is indicated whether the sensor is situated in hardrock conditions.

Station name	Latitude [deg]	Longitude [deg]	Depth [m]	Sensor	Hardrock
NL.ALK1..HGZ	52.6338	4.7444	0	accelerometer	0
NL.ALK2..HGZ	52.6510	4.7234	0	accelerometer	0
NL.ARCN..BHZ	51.5013	6.1942	0	broadband	0
NL.ASS1..HGZ	52.9778	6.5574	0	accelerometer	0
NL.ASS2..HGZ	53.0143	6.5706	0	accelerometer	0
NL.BAPP..HGZ	53.3148	6.8354	0	accelerometer	0
NL.BER1..HGZ	52.6614	4.7037	0	accelerometer	0
NL.BER2..HGZ	52.6478	4.7150	0	accelerometer	0
NL.BFB2..HGZ	53.1875	6.7655	0	accelerometer	0
NL.BGAR..HGZ	53.3679	6.7136	0	accelerometer	0
NL.BHKS..HGZ	53.2919	6.7850	0	accelerometer	0
NL.BKMZ..HGZ	53.2901	6.3189	0	accelerometer	0
NL.BLIJE..HGZ	53.3603	5.8492	0	accelerometer	0
NL.BLOP..HGZ	53.3339	6.7466	0	accelerometer	0
NL.BMD2..HGZ	53.3533	6.6472	0	accelerometer	0
NL.BOWW..HGZ	53.3405	6.8125	0	accelerometer	0
NL.BRTL..HGZ	53.2380	6.3932	0	accelerometer	0
NL.BSTD.00.HGZ	53.3122	6.6921	0	accelerometer	0
NL.BWIR..HGZ	53.3226	6.7861	0	accelerometer	0
NL.BWSE..HGZ	53.3444	6.7099	0	accelerometer	0
NL.COE2..HGZ	52.6328	6.7057	0	accelerometer	0
NL.COE3..HGZ	52.6704	6.6525	0	accelerometer	0
NL.COE4..HGZ	52.6598	6.7451	0	accelerometer	0
NL.COE..HGZ	52.6606	6.7836	0	accelerometer	0
NL.DBLS..HGZ	52.8391	6.0497	0	accelerometer	0
NL.DBN.01.BHZ	52.1017	5.1767	0	broadband	0
NL.DOK..HGZ	53.3273	5.9853	0	accelerometer	0
NL.DON..HGZ	53.1005	6.5429	0	accelerometer	0
NL.DR014..HHZ	52.7987	6.1785	200	geophone	0
NL.DR024..HHZ	52.7034	6.5081	200	geophone	0
NL.DR034..HHZ	52.8976	6.3084	200	geophone	0
NL.DRA..HGZ	53.1202	6.1033	0	accelerometer	0
NL.EETW..HGZ	53.3348	6.1054	0	accelerometer	0
NL.ELE..HGZ	52.9566	6.5776	0	accelerometer	0
NL.ENM4.01.HHZ	53.4064	6.4817	200	geophone	0
NL.ENV2.01.HHZ	52.8944	6.6337	100	geophone	0

NL.EVK..HGZ	53.0797	6.8116	0	accelerometer	0
NL.FDKD..HGZ	52.8483	6.1788	0	accelerometer	0
NL.FR014..HHZ	53.1424	5.8868	200	geophone	0
NL.FR024..HHZ	52.8559	5.9304	200	geophone	0
NL.FR034..HHZ	52.9815	5.7126	200	geophone	0
NL.FR044..HHZ	52.8761	6.1187	200	geophone	0
NL.FR054..HHZ	53.0705	5.8685	200	geophone	0
NL.FR064..HHZ	53.0007	6.0237	200	geophone	0
NL.FSW5.01.HHZ	53.2135	7.1195	300	geophone	0
NL.G014..HHZ	53.4425	6.7312	200	geophone	0
NL.G024..HHZ	53.4490	6.8058	200	geophone	0
NL.G034..HHZ	53.4159	6.5976	200	geophone	0
NL.G044..HHZ	53.4147	6.6747	200	geophone	0
NL.G054..HHZ	53.4184	6.7713	200	geophone	0
NL.G064..HHZ	53.4111	6.8709	200	geophone	0
NL.G074..HHZ	53.3946	6.5458	200	geophone	0
NL.G084..HHZ	53.3954	6.6439	200	geophone	0
NL.G094..HHZ	53.3878	6.7245	200	geophone	0
NL.G104..HHZ	53.3785	6.8044	200	geophone	0
NL.G114..HHZ	53.3669	6.8891	200	geophone	0
NL.G124..HHZ	53.3699	6.5714	200	geophone	0
NL.G134..HHZ	53.3489	6.6686	200	geophone	0
NL.G144..HHZ	53.3586	6.7708	200	geophone	0
NL.G164..HHZ	53.3385	6.5314	200	geophone	0
NL.G174..HHZ	53.3362	6.6363	200	geophone	0
NL.G184..HHZ	53.3306	6.7217	200	geophone	0
NL.G194..HHZ	53.3378	6.8170	200	geophone	0
NL.G204..HHZ	53.3353	6.8952	200	geophone	0
NL.G214..HHZ	53.3134	6.5937	200	geophone	0
NL.G224..HHZ	53.3103	6.6776	200	geophone	0
NL.G234..HHZ	53.3095	6.7675	200	geophone	0
NL.G244..HHZ	53.2900	6.8562	200	geophone	0
NL.G254..HHZ	53.3081	6.9433	200	geophone	0
NL.G264..HHZ	53.2821	6.5458	200	geophone	0
NL.G274..HHZ	53.2832	6.6281	200	geophone	0
NL.G284..HHZ	53.2722	6.7292	200	geophone	0
NL.G294..HHZ	53.2814	6.7870	200	geophone	0
NL.G304..HHZ	53.2792	6.8951	200	geophone	0
NL.G314..HHZ	53.2962	7.0218	200	geophone	0
NL.G324..HHZ	53.2498	6.5792	200	geophone	0
NL.G334..HHZ	53.2499	6.6708	200	geophone	0
NL.G344..HHZ	53.2525	6.7651	200	geophone	0
NL.G354..HHZ	53.2549	6.8621	200	geophone	0
NL.G364..HHZ	53.2651	6.9245	200	geophone	0
NL.G374..HHZ	53.2740	7.0350	200	geophone	0
NL.G384..HHZ	53.2005	6.5476	200	geophone	0
NL.G394..HHZ	53.2253	6.7241	200	geophone	0
NL.G404..HHZ	53.2250	6.8099	200	geophone	0
NL.G414..HHZ	53.2234	6.8919	200	geophone	0
NL.G424..HHZ	53.2217	6.9886	200	geophone	0
NL.G434..HHZ	53.2319	7.0791	200	geophone	0
NL.G444..HHZ	53.2042	6.6840	200	geophone	0
NL.G454..HHZ	53.1995	6.7645	200	geophone	0

NL.G464..HHZ	53.1978	6.8498	200	geophone	0
NL.G474..HHZ	53.1964	6.9413	200	geophone	0
NL.G484..HHZ	53.1973	7.0290	200	geophone	0
NL.G494..HHZ	53.1765	6.7306	200	geophone	0
NL.G504..HHZ	53.1745	6.8093	200	geophone	0
NL.G514..HHZ	53.1732	6.8848	200	geophone	0
NL.G524..HHZ	53.1699	6.9837	200	geophone	0
NL.G534..HHZ	53.1674	7.0801	200	geophone	0
NL.G544..HHZ	53.1243	6.7580	200	geophone	0
NL.G554..HHZ	53.1433	6.8461	200	geophone	0
NL.G564..HHZ	53.1484	6.9309	200	geophone	0
NL.G574..HHZ	53.1351	7.0222	200	geophone	0
NL.G584..HHZ	53.1103	6.9071	200	geophone	0
NL.G594..HHZ	53.1160	6.9877	200	geophone	0
NL.G604..HHZ	53.2979	7.0752	200	geophone	0
NL.G614..HHZ	53.3819	6.6813	200	geophone	0
NL.G624..HHZ	53.3870	6.8419	200	geophone	0
NL.G634..HHZ	53.3314	6.9109	200	geophone	0
NL.G644..HHZ	53.2207	7.0380	200	geophone	0
NL.G654..HHZ	53.1692	6.9543	200	geophone	0
NL.G664..HHZ	53.1794	6.5852	200	geophone	0
NL.G674..HHZ	53.3213	6.8228	200	geophone	0
NL.G684..HHZ	53.2351	6.6347	200	geophone	0
NL.G694..HHZ	53.2128	6.6438	200	geophone	0
NL.G704..HHZ	53.2407	6.5379	200	geophone	0
NL.G710..HGZ	53.1483	6.6507	0	accelerometer	0
NL.G720..HGZ	53.2462	6.4792	0	accelerometer	0
NL.G730..HGZ	53.2719	6.4301	0	accelerometer	0
NL.G740..HGZ	53.3043	6.3894	0	accelerometer	0
NL.G750..HGZ	53.3078	6.4875	0	accelerometer	0
NL.G760..HGZ	53.3507	6.3901	0	accelerometer	0
NL.G770..HGZ	53.3403	6.3304	0	accelerometer	0
NL.G780..HGZ	53.3648	6.4894	0	accelerometer	0
NL.G790..HGZ	53.3836	6.3319	0	accelerometer	0
NL.G800..HGZ	53.4154	6.4876	0	accelerometer	0
NL.G81B..HHZ	53.3103	6.6776	100	geophone	0
NL.G82B..HHZ	53.2253	6.7241	100	geophone	0
NL.G83B..HHZ	53.1484	6.9309	100	geophone	0
NL.G84B..HHZ	53.2207	7.0380	100	geophone	0
NL.GK014..HHZ	53.2906	6.2471	200	geophone	0
NL.GK024..HHZ	53.3010	6.2927	200	geophone	0
NL.GK034..HHZ	53.2827	6.3056	200	geophone	0
NL.GK044..HHZ	53.2462	6.3261	200	geophone	0
NL.GRP..HGZ	53.1788	5.9779	0	accelerometer	0
NL.HGN.02.BHZ	50.7640	5.9317	4	broadband	1
NL.HLB.11.EPZ	53.1465	7.0036	60	geophone	0
NL.HLG.11.EPZ	53.1491	7.0071	58	geophone	0
NL.HLH.11.EPZ	53.1404	7.0065	60	geophone	0
NL.HLK.11.EPZ	53.1450	7.0096	60	geophone	0
NL.HOEV..HGZ	52.9025	6.1055	0	accelerometer	0
NL.HRDB..HGZ	52.5770	6.6259	0	accelerometer	0
NL.HRKB..BHZ	51.1879	6.1678	0	broadband	0
NL.HWF4.01.HHZ	53.0710	6.3512	200	geophone	0

NL.J01..HHZ	52.7650	5.0545	0	geophone	0
NL.J02.01.HHZ	52.7255	5.1948	0	geophone	0
NL.KLWL..HGZ	51.7755	4.4406	0	accelerometer	0
NL.KMP..HGZ	53.3206	6.1920	0	accelerometer	0
NL.LUTT..HGZ	52.6133	6.5715	0	accelerometer	0
NL.MAME..BHZ	50.8000	5.9727	0	broadband	1
NL.MAR..HGZ	53.1540	6.2631	0	accelerometer	0
NL.N014..HHZ	53.1185	6.4733	200	geophone	0
NL.N024..HHZ	53.0720	6.4530	200	geophone	0
NL.N034..HHZ	53.1110	6.4045	200	geophone	0
NL.NGA..HGZ	53.1364	6.0257	0	accelerometer	0
NL.NH014..HHZ	52.5657	4.9965	200	geophone	0
NL.NIW4.01.HHZ	53.3548	6.0430	120	geophone	0
NL.NLDW..HGZ	52.0050	4.1957	0	accelerometer	0
NL.ODBK..HGZ	52.9405	6.1339	0	accelerometer	0
NL.ODHP..HGZ	52.8852	6.0249	0	accelerometer	0
NL.OLK..HGZ	52.6293	4.8495	0	accelerometer	0
NL.OOTH..HGZ	52.7295	6.7534	0	accelerometer	0
NL.OPLO.01.BHZ	51.5888	5.8121	0	broadband	0
NL.OSTM..HGZ	53.1875	6.0857	0	accelerometer	0
NL.OTL2.01.HHZ	52.6289	4.8227	100	geophone	0
NL.PET..HGZ	52.7706	4.6595	0	accelerometer	0
NL.PMR..HGZ	52.5116	5.0106	0	accelerometer	0
NL.PPB3.01.HHZ	52.6526	4.6700	165	geophone	0
NL.ROLD..BHZ	50.8694	6.0847	0	broadband	0
NL.SL014..HHZ	52.8551	4.8683	200	geophone	0
NL.SNB..HGZ	52.6842	6.8719	0	accelerometer	0
NL.SPY4.01.HHZ	53.4098	6.7838	120	geophone	0
NL.SRHV..HGZ	53.1833	6.1849	0	accelerometer	0
NL.STR..HGZ	51.7437	4.5573	0	accelerometer	0
NL.SUH4.01.HHZ	53.2113	6.2110	120	geophone	0
NL.T014..HHZ	52.4643	6.7401	200	geophone	0
NL.T024..HHZ	52.4634	6.9787	200	geophone	0
NL.T034..HHZ	52.4297	6.7918	200	geophone	0
NL.T044..HHZ	52.4375	6.8905	200	geophone	0
NL.T054..HHZ	52.3806	6.9057	200	geophone	0
NL.T064..HHZ	52.3432	6.9097	200	geophone	0
NL.T084..HHZ	52.3893	6.9965	200	geophone	0
NL.TERZ.01.BHZ	50.7568	5.9061	250	broadband	1
NL.TWZL..HGZ	53.2325	6.0927	0	accelerometer	0
NL.UTRP..HGZ	53.0969	6.1842	0	accelerometer	0
NL.VBG4.01.HHZ	52.5440	6.6693	192	geophone	0
NL.VDM24..HHZ	53.0844	6.8336	200	geophone	0
NL.VHP..HGZ	53.0957	5.9553	0	accelerometer	0
NL.VKB..BHZ	50.8669	5.7847	0	broadband	1
NL.VLD..HGZ	52.8668	6.2157	0	accelerometer	0
NL.VLW3.01.HHZ	52.9682	7.0972	150	geophone	0
NL.VNDM4..HHZ	53.1151	6.8390	200	geophone	0
NL.VNWD..HGZ	53.2368	5.9984	0	accelerometer	0
NL.VPR..HGZ	51.8915	4.3715	0	accelerometer	0
NL.VRS..HGZ	53.0360	6.6046	0	accelerometer	0
NL.WDB4.01.HHZ	53.2083	6.7355	197	geophone	0
NL.WMH4.01.HHZ	52.7096	4.7498	200	geophone	0

NL.WSVN..HGZ	52.8204	6.1983	0	accelerometer	0
NL.WTSB.01.BHZ	51.9663	6.7989	0	broadband	1
NL.WYN4.00.HHZ	53.1995	5.4578	120	geophone	0
NL.ZDL..HGZ	53.0791	6.7026	0	accelerometer	0
NL.ZH014..HHZ	51.9137	4.2881	200	geophone	0
NL.ZH024..HHZ	52.0744	4.4884	200	geophone	0
NL.ZH034..HHZ	51.9783	4.4001	200	geophone	0
NL.ZH044..HHZ	51.8417	4.3626	200	geophone	0
NL.ZH064..HHZ	51.9781	4.2204	200	geophone	0
NL.ZH074..HHZ	52.1037	4.2946	200	geophone	0
NL.ZLV4.01.HHZ	53.0921	6.7533	200	geophone	0
GE.IBBN..HHZ	52.3063	7.7592	1	broadband	1
GR.AHRW..HHZ	50.5410	7.0760	0	broadband	1
GR.BUG..HHZ	51.4455	7.2643	0	broadband	1
GR.RAST..EHZ	52.9096	7.7164	0	geophone	0
NR.NE116..BHZ	51.5042	4.9209	0	broadband	0
BE.BOST..HHZ	51.2382	2.9387	304	broadband	1
BE.MEM..HHZ	50.6087	6.0096	0	broadband	1
BE.BEBN..HHZ	50.7970	5.6778	0	broadband	1
BE.OPTB..HHZ	51.1115	5.6360	379	broadband	1
BE.RCHB..HHZ	50.1552	5.2268	43	broadband	1
BE.UCC..HHZ	50.7972	4.3604	141	broadband	1

D 2022 NL-wide maps with location-uncertainty parameters

D.1 Largest-horizontal-location-uncertainty (σ_1) maps

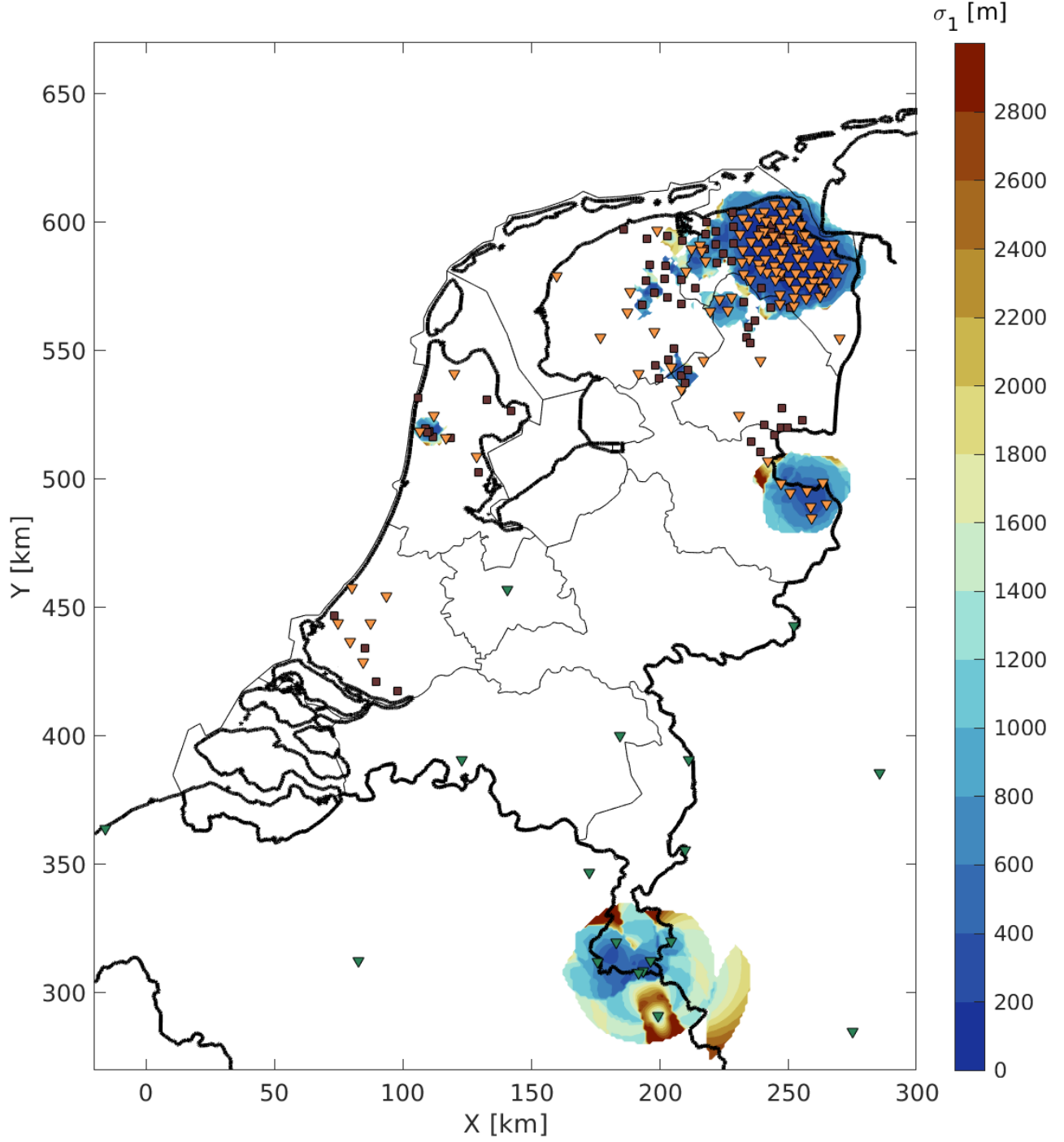


Figure D.1: Spatial distribution of σ_1 in 2022 for $M=0.5$ (color map). White means that events with $M=0.5$ can likely not be detected. Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers. Dark red depicts uncertainties of 2800 m and higher.

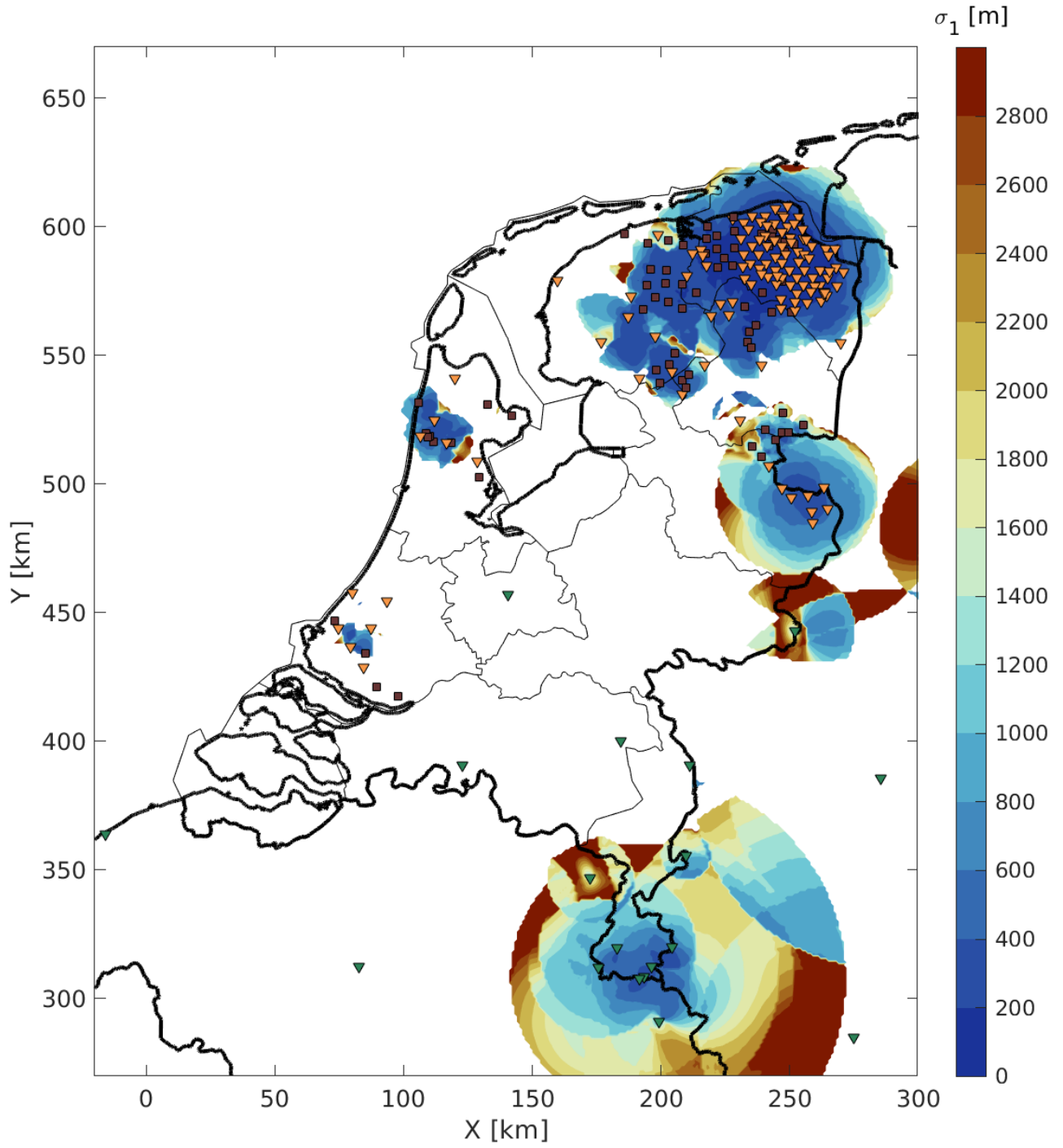


Figure D.2: Spatial distribution of σ_1 in 2022 for $M=1.0$ (color map). White means that events with $M=1.0$ can likely not be detected. Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

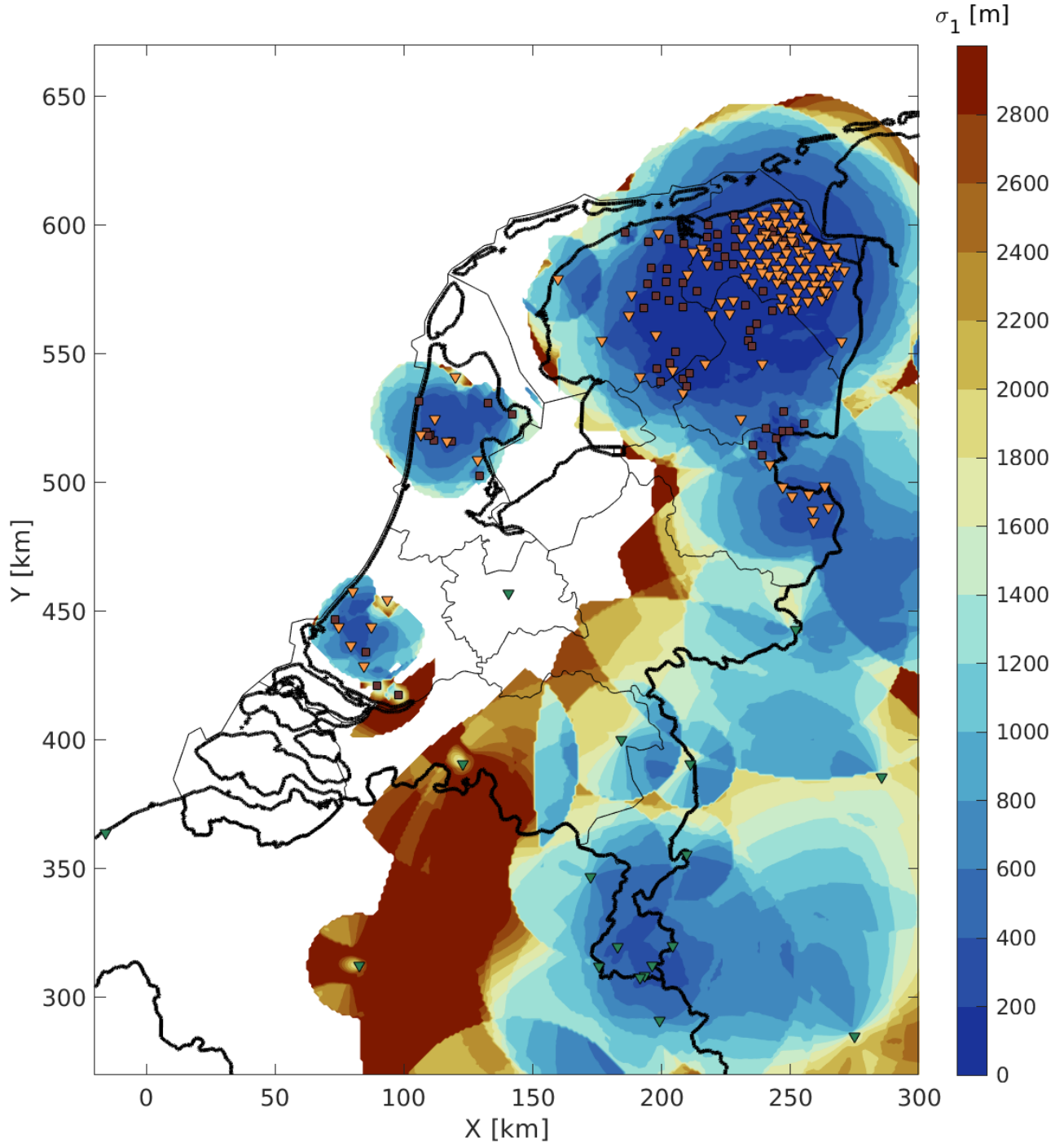


Figure D.3: Spatial distribution of σ_1 in 2022 for $M=1.5$ (color map). White means that events with $M=1.5$ can likely not be detected. Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

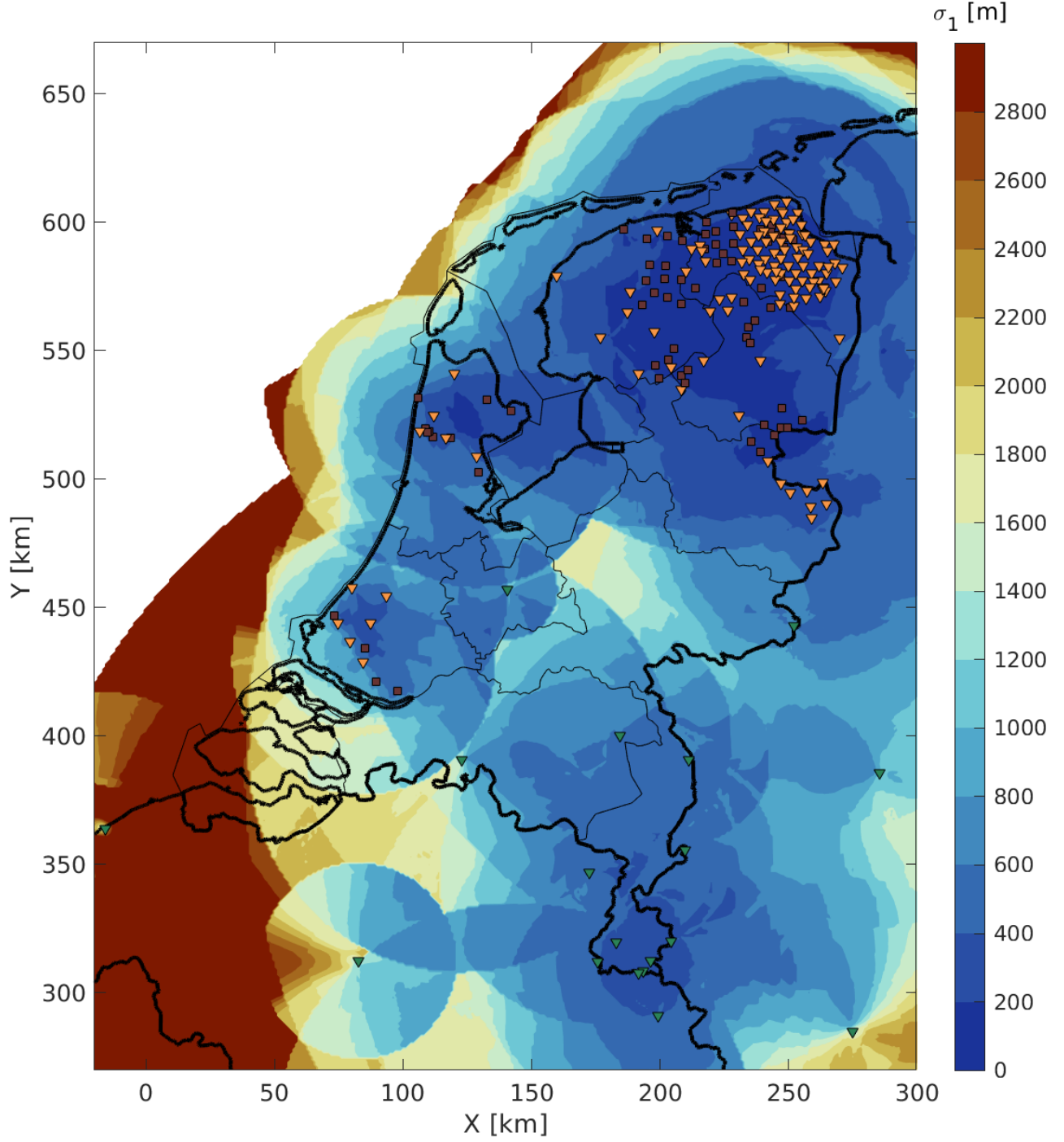


Figure D.4: Spatial distribution of σ_1 in 2022 for $M=2.0$ (color map). White means that events with $M=2.0$ can likely not be detected. Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

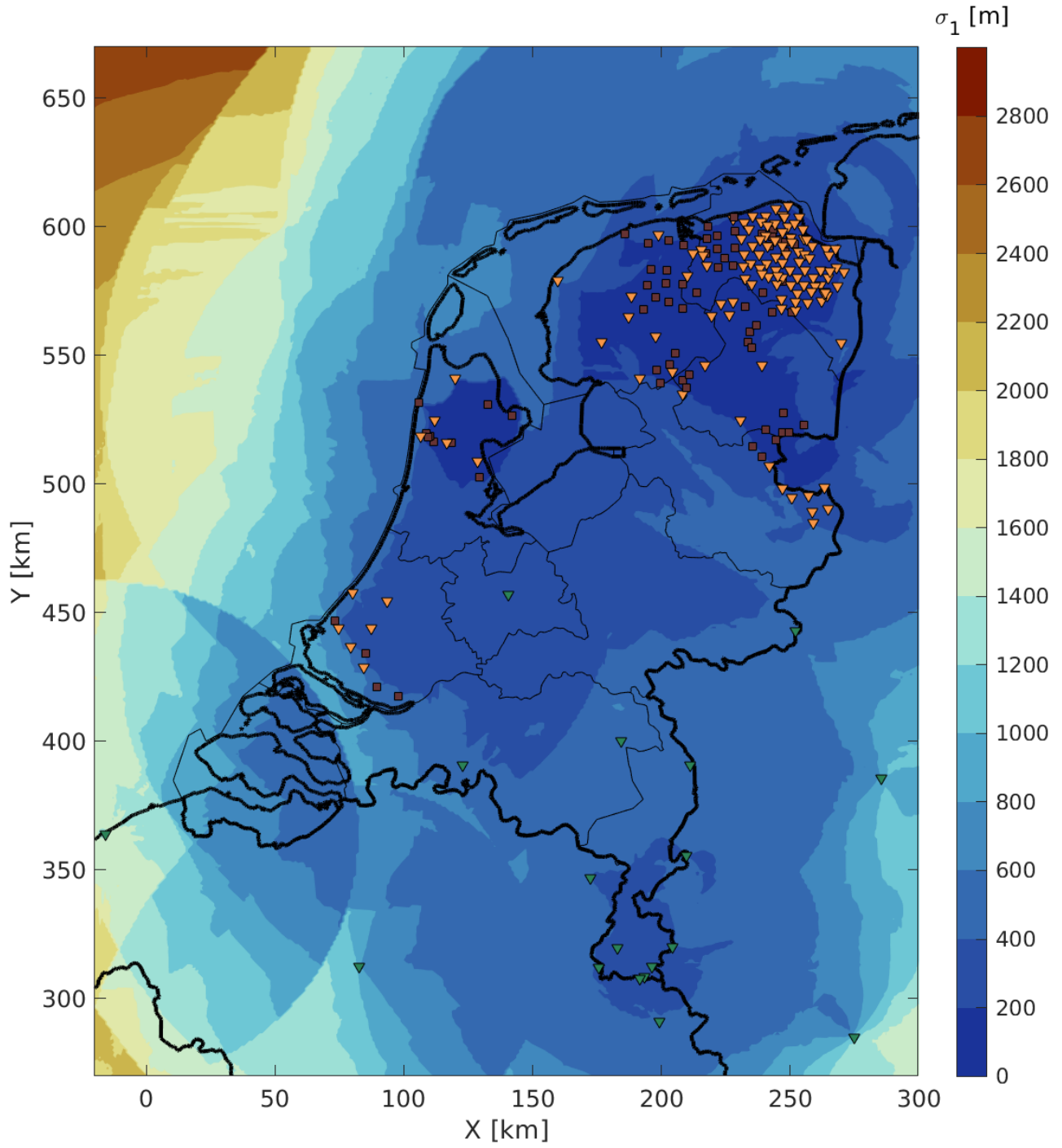


Figure D.5: Spatial distribution of σ_1 in 2022 for $M=2.5$ (color map). White means that events with $M=2.5$ can likely not be detected. Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

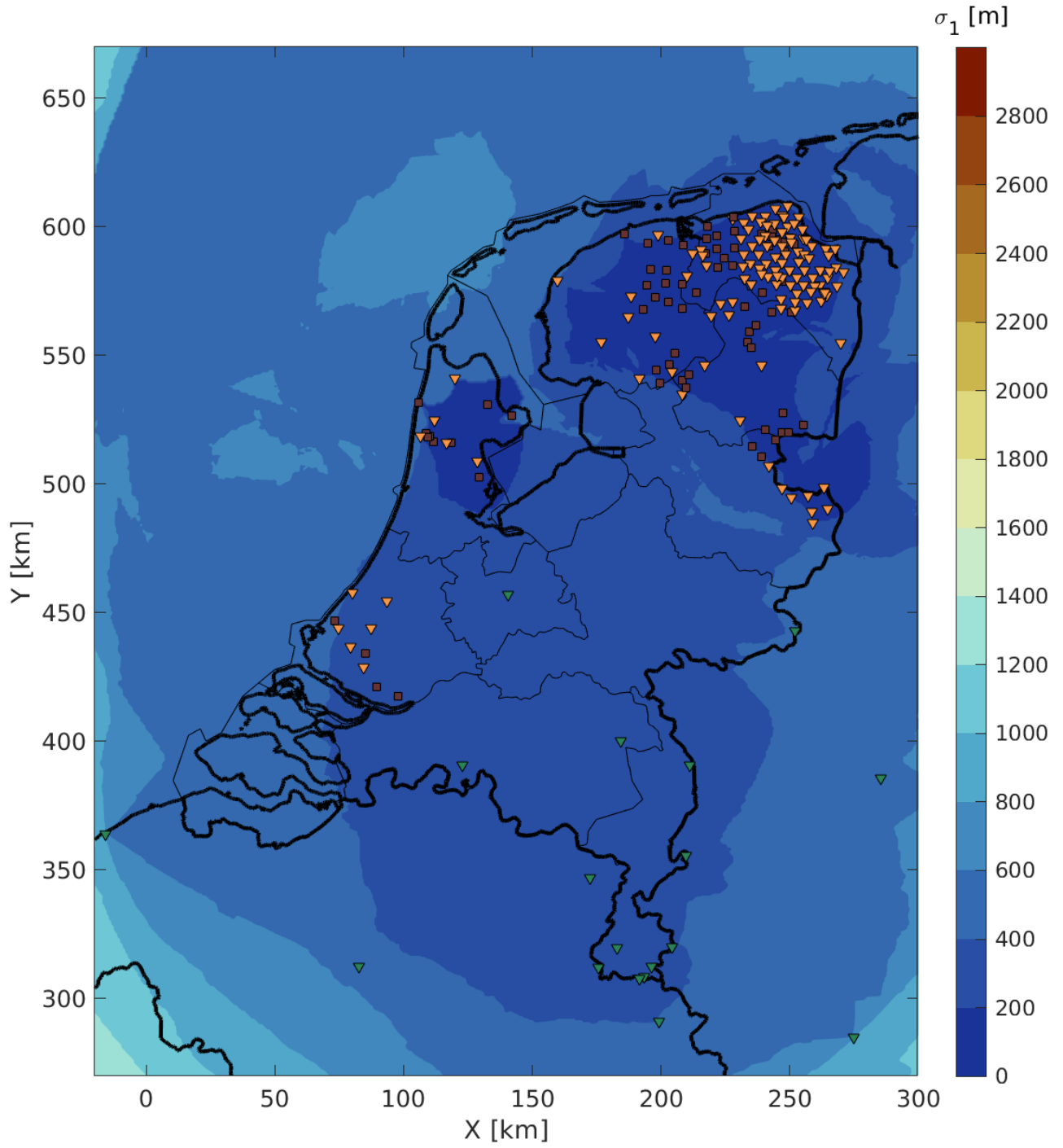


Figure D.6: Spatial distribution of σ_1 in 2022 for $M=3.0$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

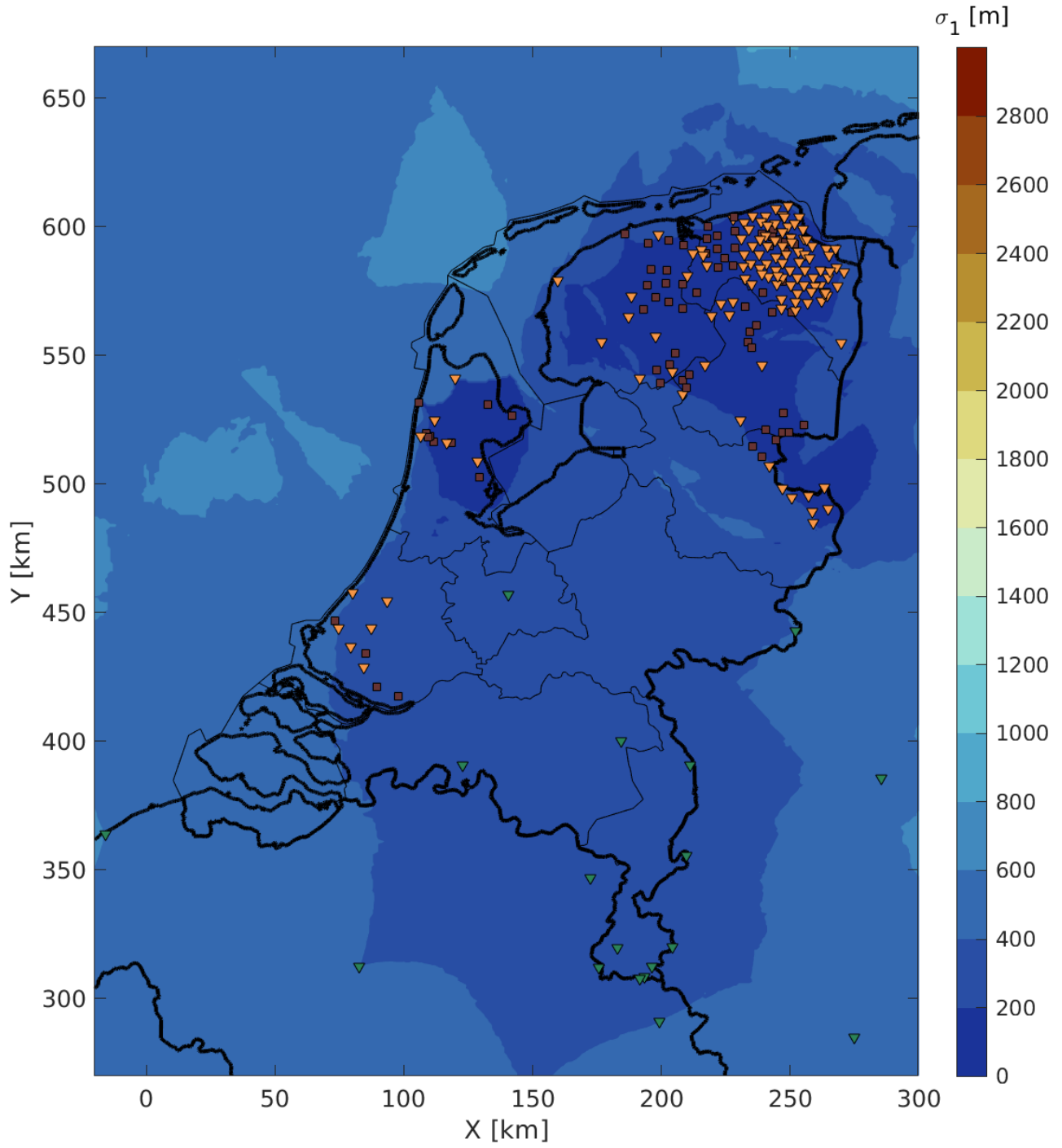


Figure D.7: Spatial distribution of σ_1 in 2022 for $M=3.5$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

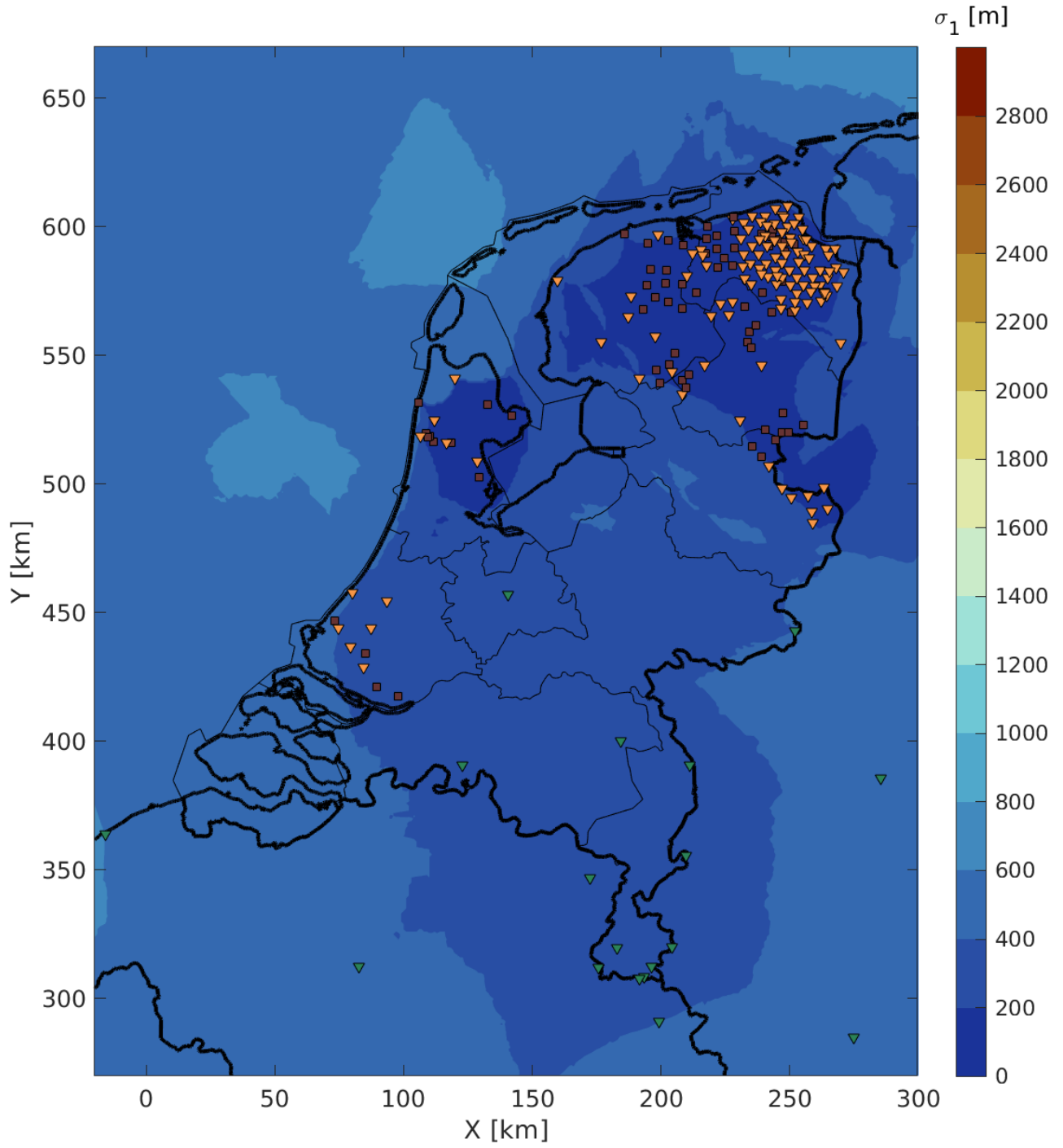


Figure D.8: Spatial distribution of σ_1 in 2022 for $M=4.0$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

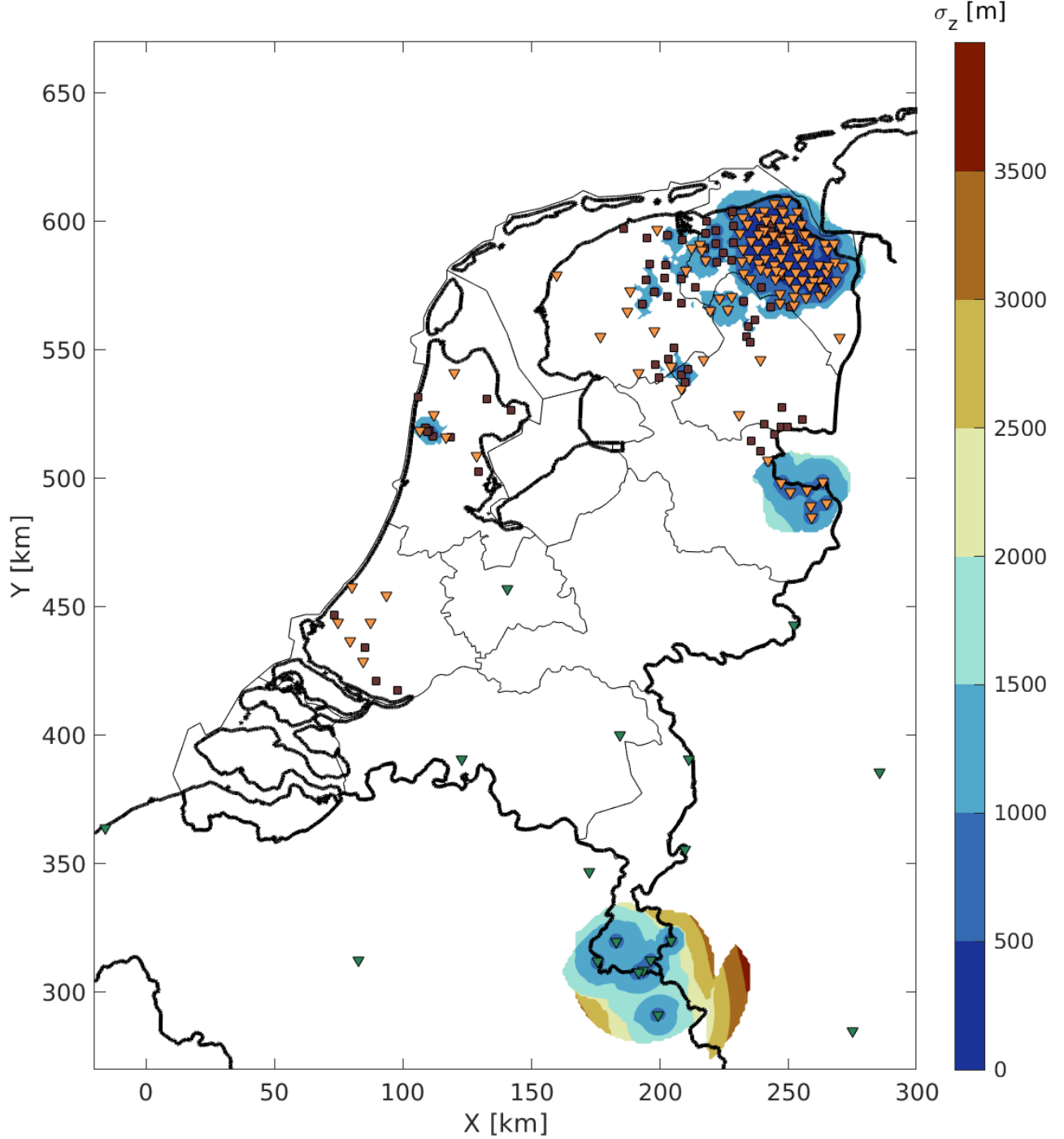
D.2 Vertical-location-uncertainty (σ_Z) maps

Figure D.9: Spatial distribution of σ_Z in 2022 for $M=0.5$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. White means that events with $M=0.5$ can likely not be detected. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers. Dark red depicts uncertainties of 3500 m and higher.

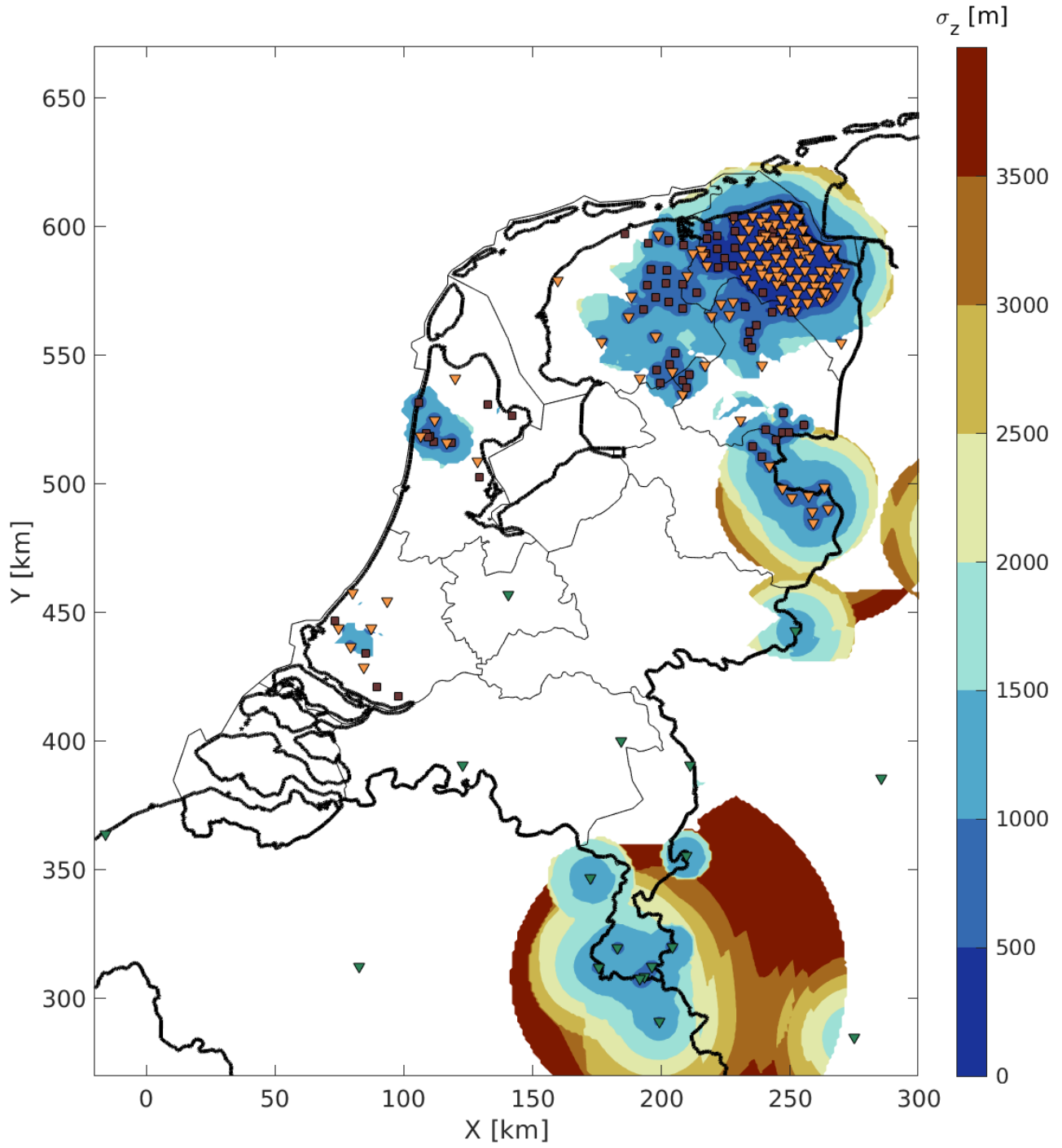


Figure D.10: Spatial distribution of σ_z in 2022 for $M=1.0$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. White means that events with $M=1.0$ can likely not be detected. Uncertainties of 3500 m and higher are shown in red. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

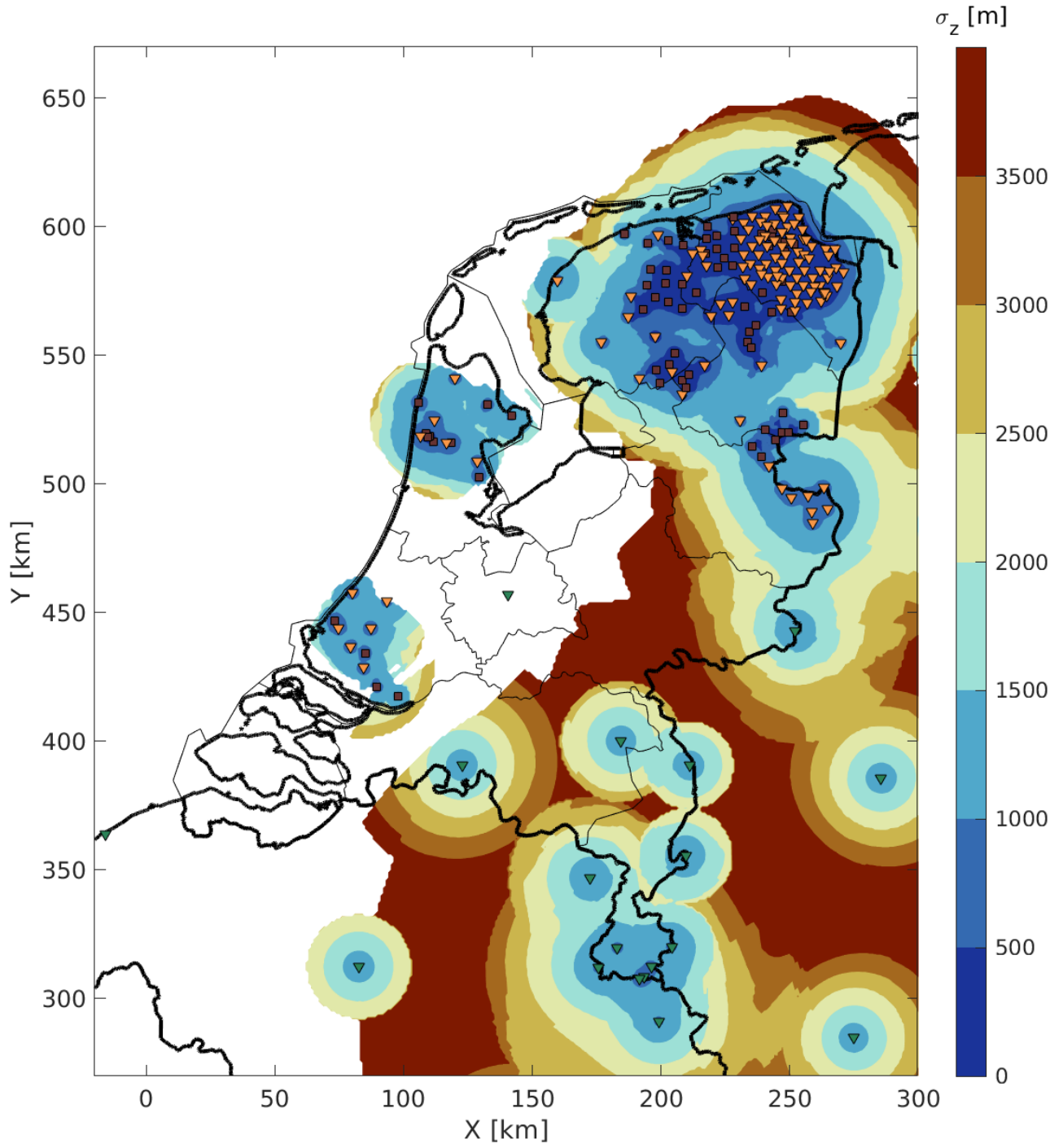


Figure D.11: Spatial distribution of σ_z in 2022 for $M=1.5$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. White means that events with $M=1.5$ can likely not be detected. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

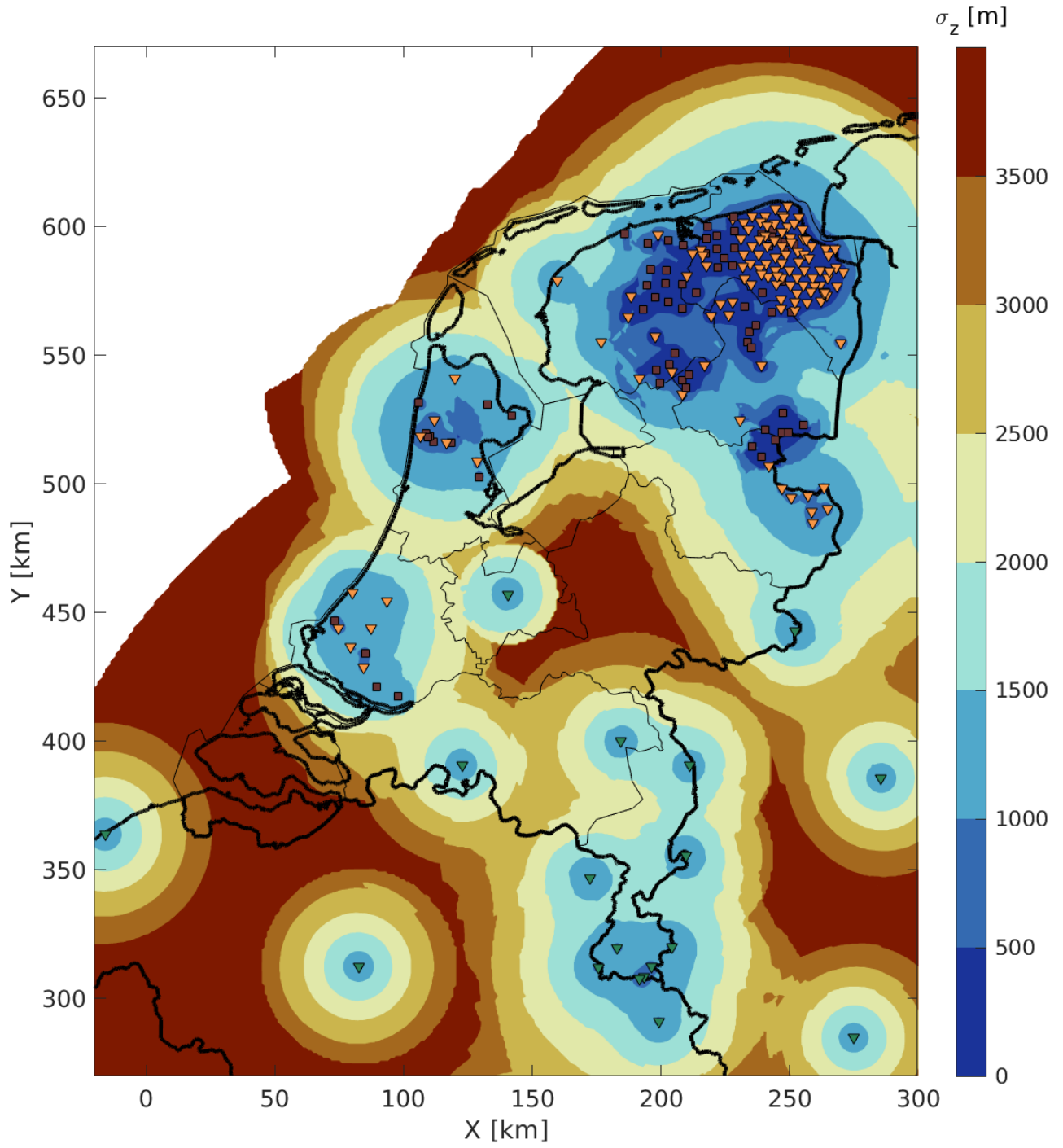


Figure D.12: Spatial distribution of σ_z in 2022 for $M=2.0$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. White means that events with $M=2.0$ can likely not be detected. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

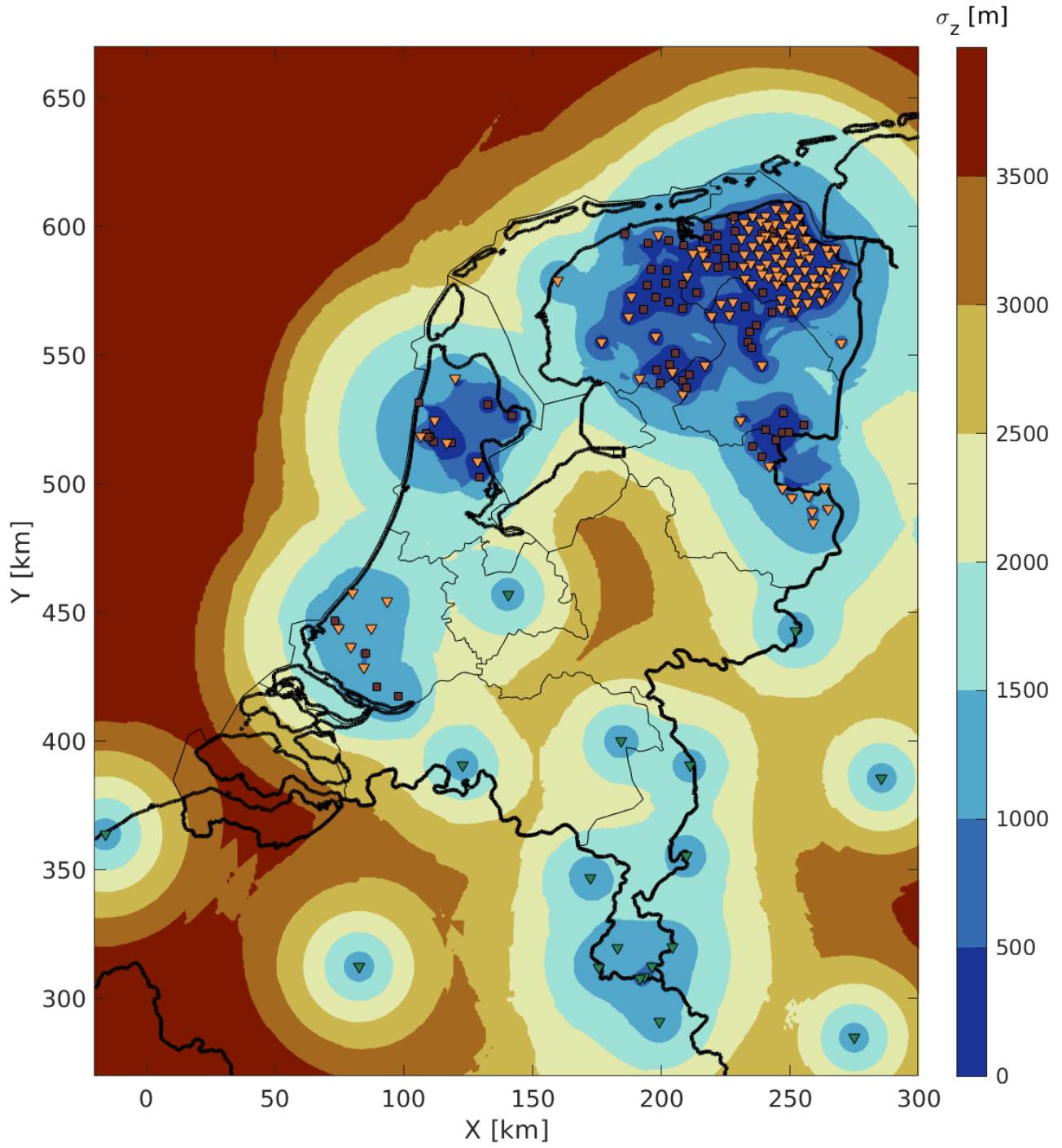


Figure D.13: Spatial distribution of σ_z in 2022 for $M=2.5$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. White means that events with $M=2.5$ can likely not be detected. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

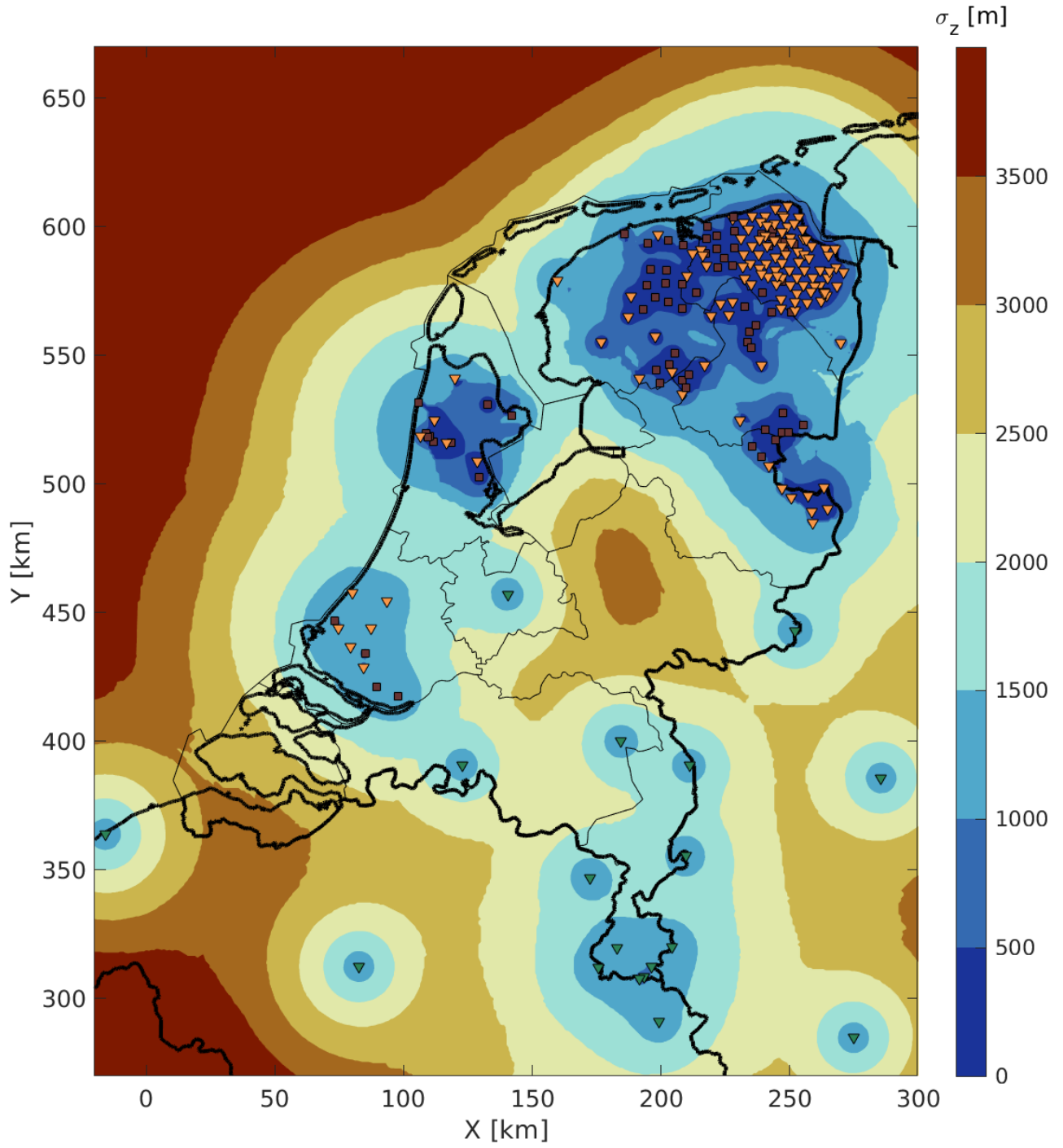


Figure D.14: Spatial distribution of σ_Z in 2022 for $M=3.0$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

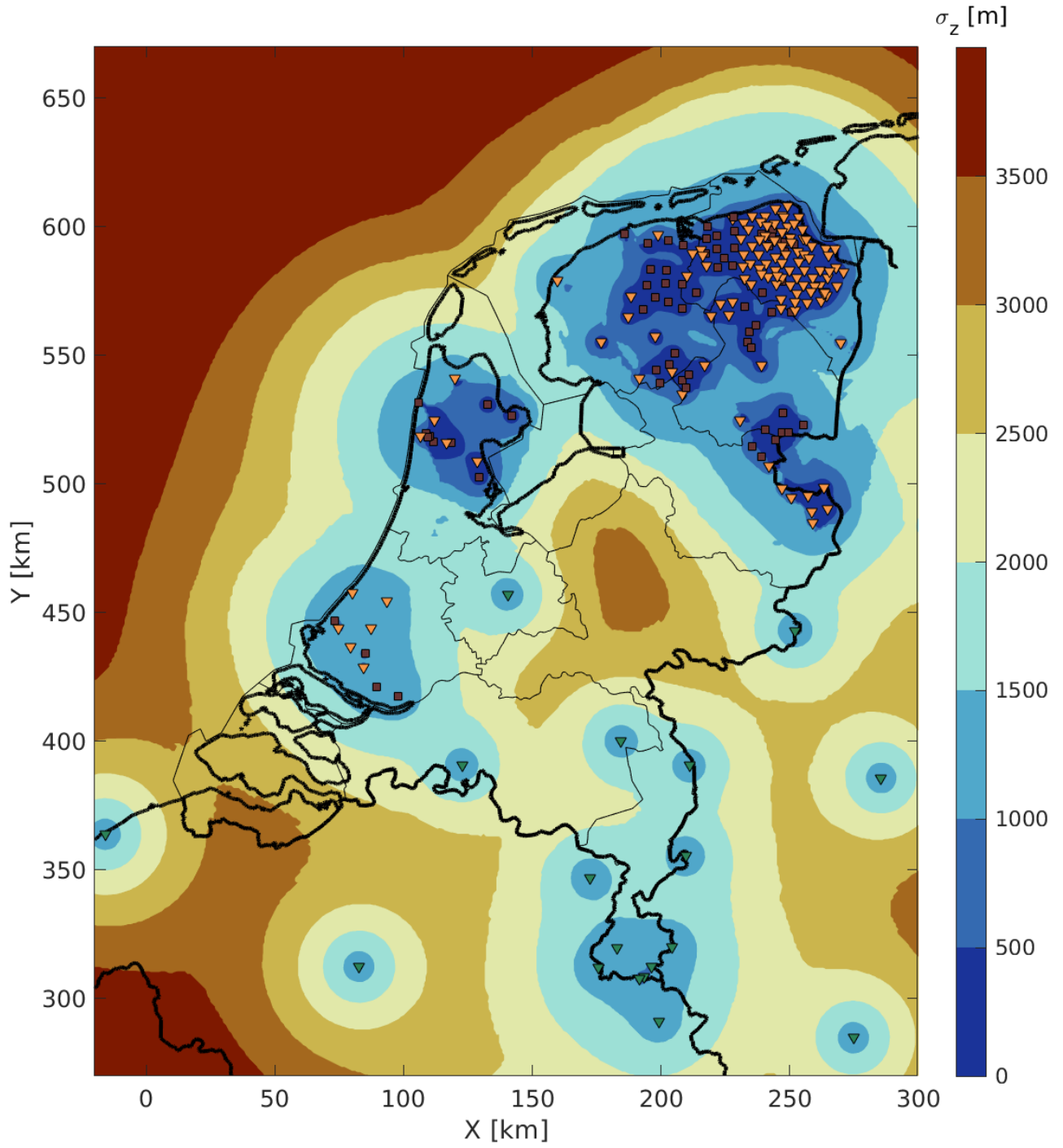


Figure D.15: Spatial distribution of σ_z in 2022 for $M=3.5$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

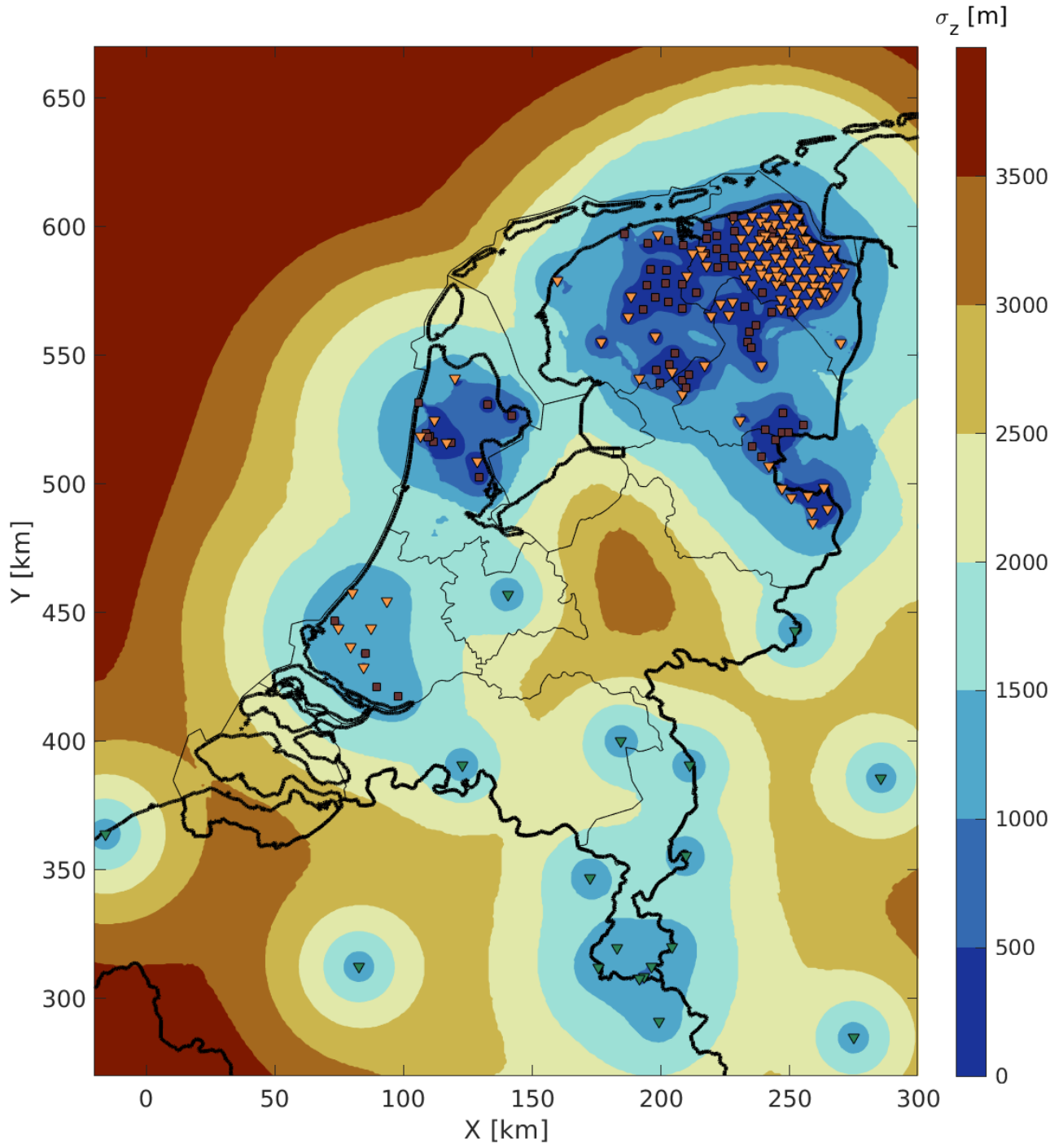


Figure D.16: Spatial distribution of σ_z in 2022 for $M=4.0$ (color map). Surface sensors are shown with squares, deep geophones with orange triangles and broadbands with green triangles. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

D.3 Azimuthal-gap (ψ) maps

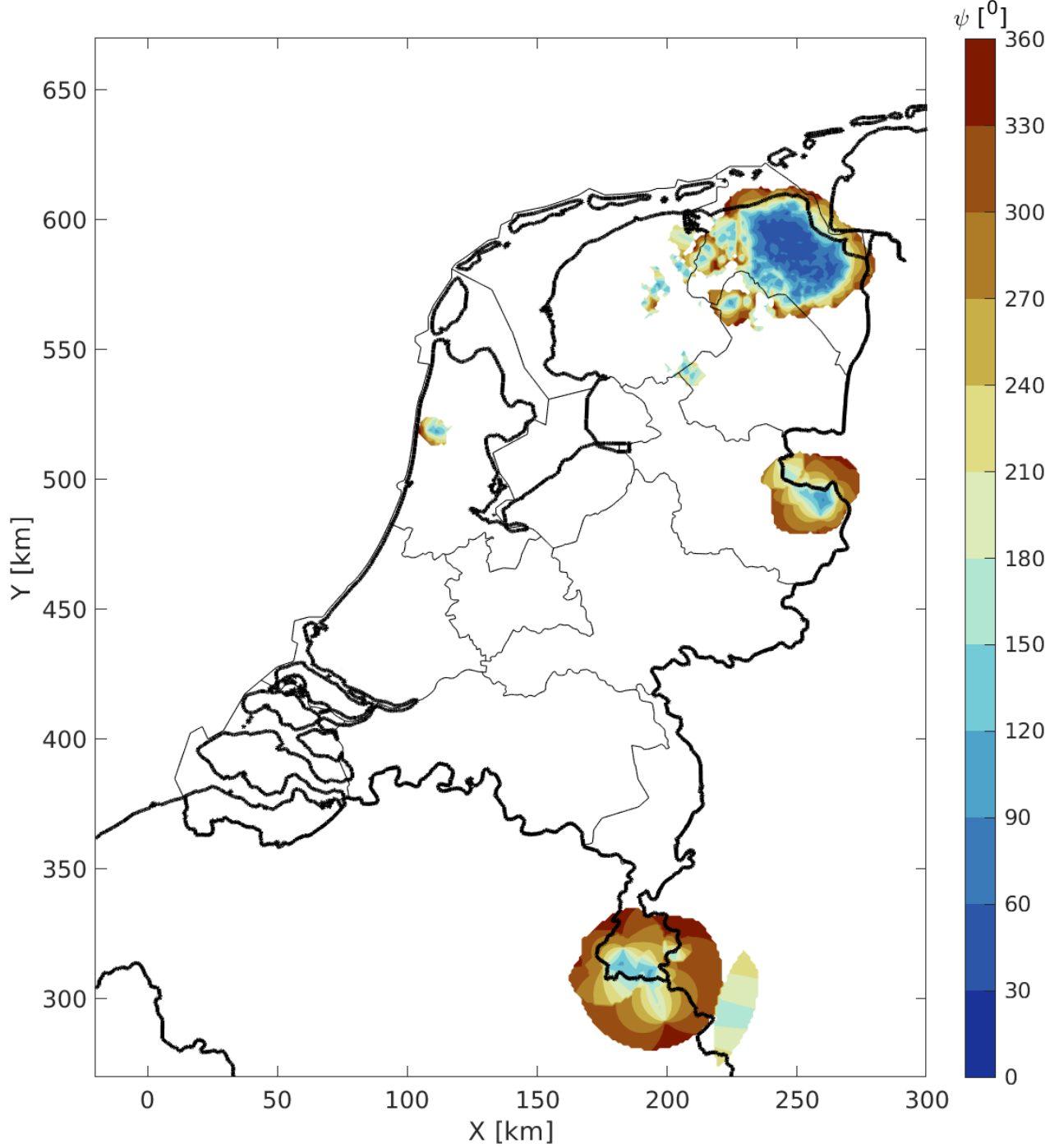


Figure D.17: Spatial distribution of azimuthal gap ψ in 2022 for $M=0.5$ (color map). White means that events with $M=0.5$ can likely not be detected. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

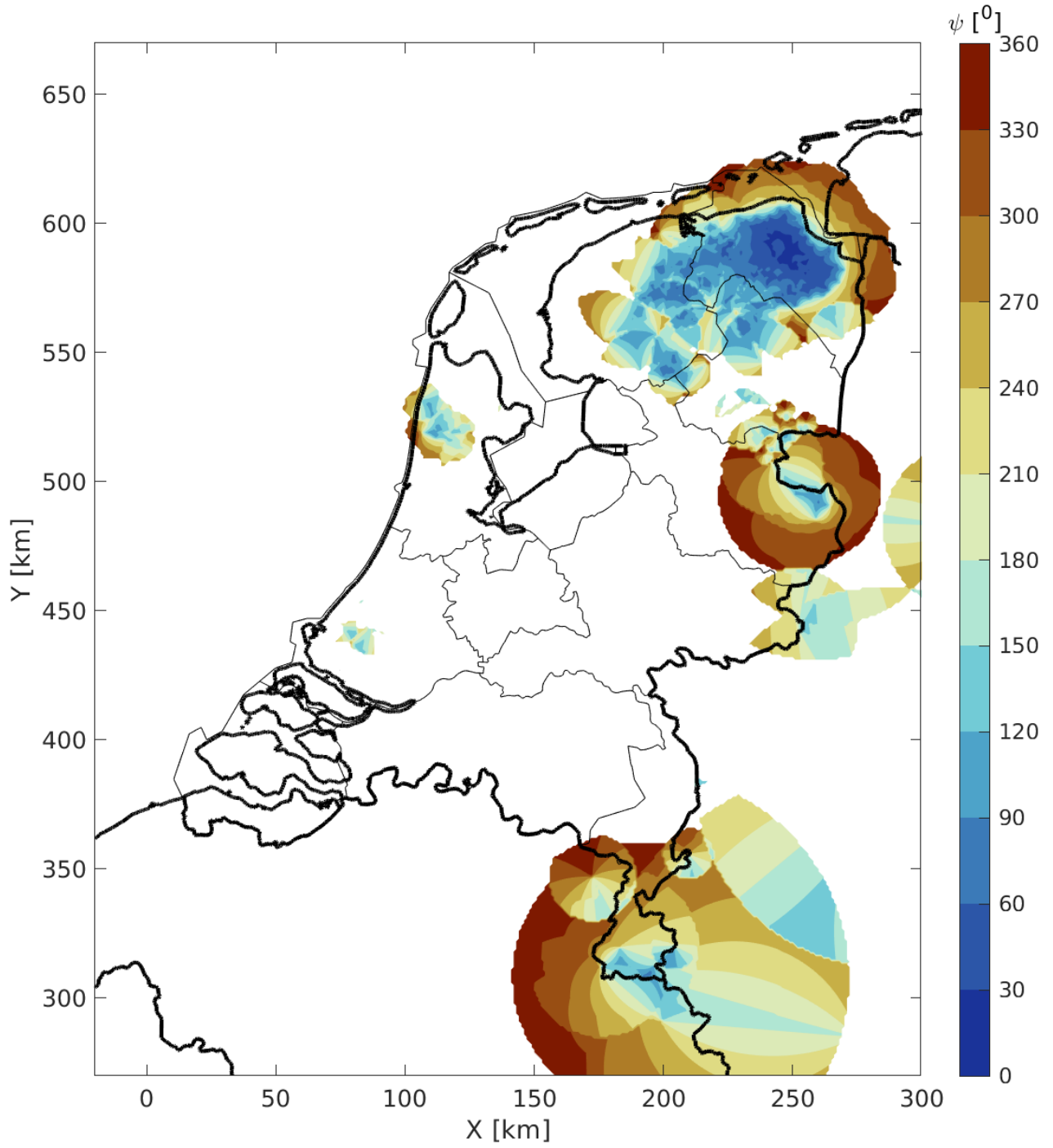


Figure D.18: Spatial distribution of azimuthal gap ψ in 2022 for $M=1.0$ (color map). White means that events with $M=1.0$ can likely not be detected. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

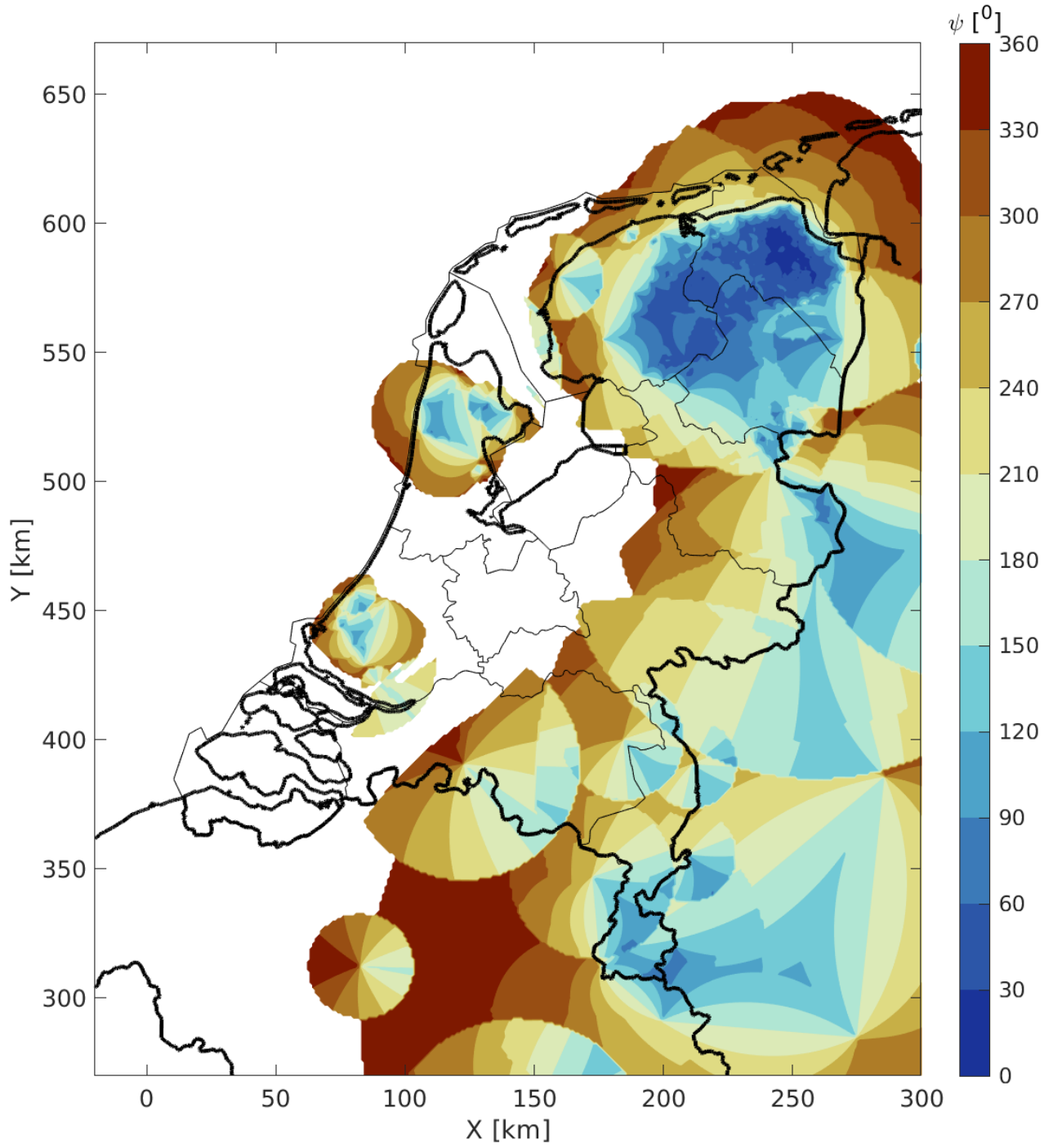


Figure D.19: Spatial distribution of azimuthal gap ψ in 2022 for $M=1.5$ (color map). White means that events with $M=1.5$ can likely not be detected. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

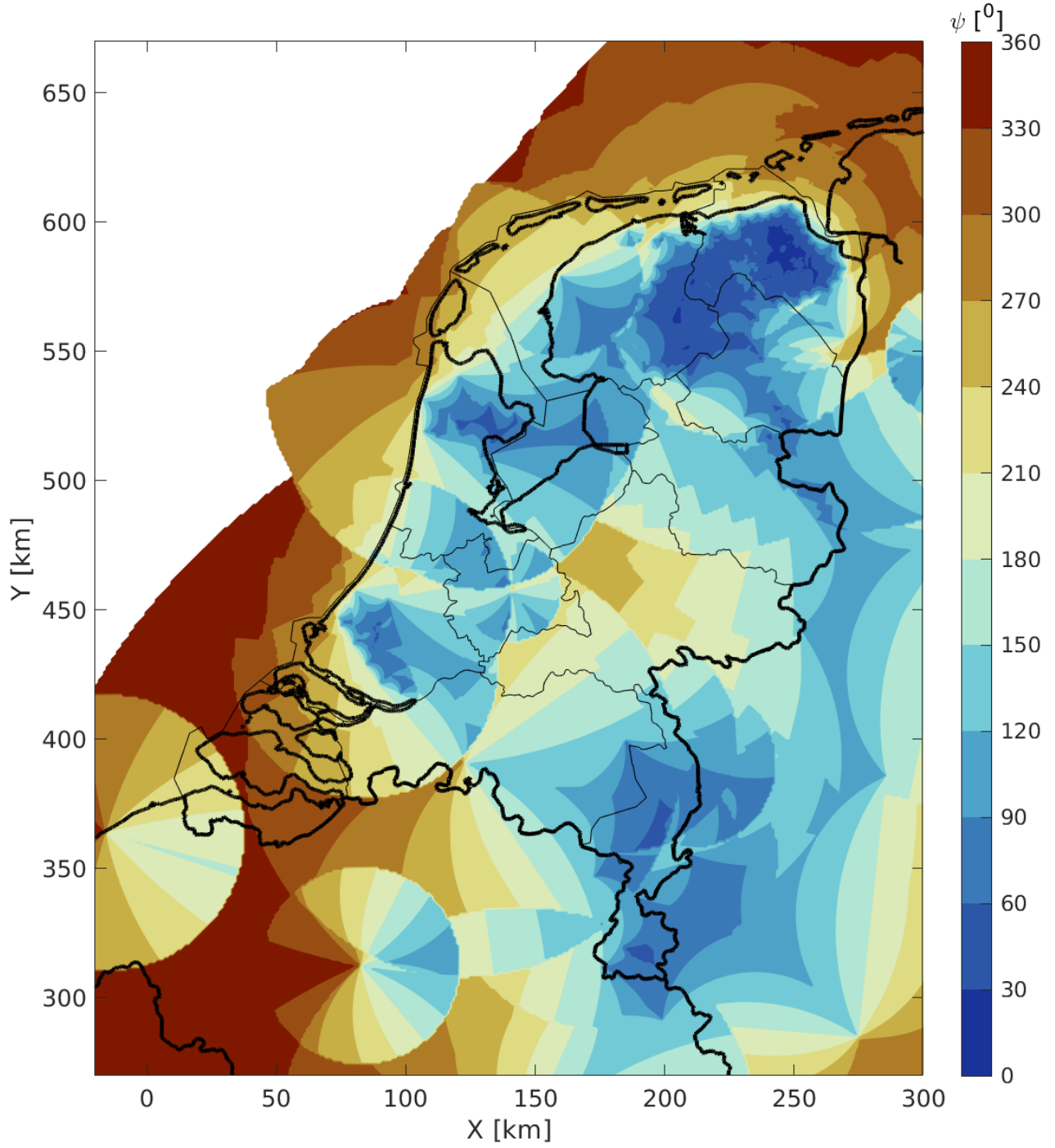


Figure D.20: Spatial distribution of azimuthal gap ψ in 2022 for $M=2.0$ (color map). White means that events with $M=2.0$ can likely not be detected. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

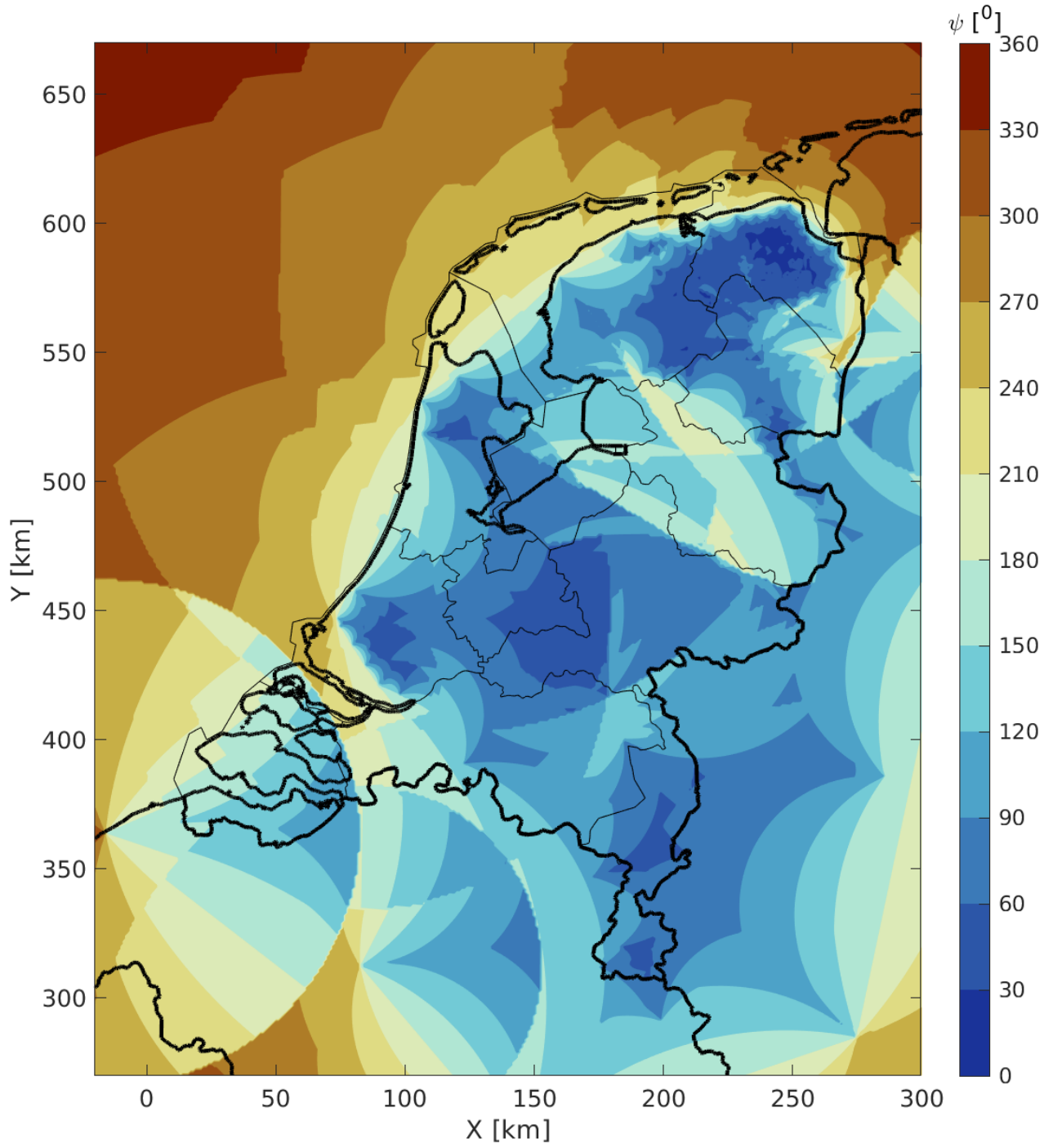


Figure D.21: Spatial distribution of azimuthal gap ψ in 2022 for $M=2.5$ (color map). White means that events with $M=2.5$ can likely not be detected. Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

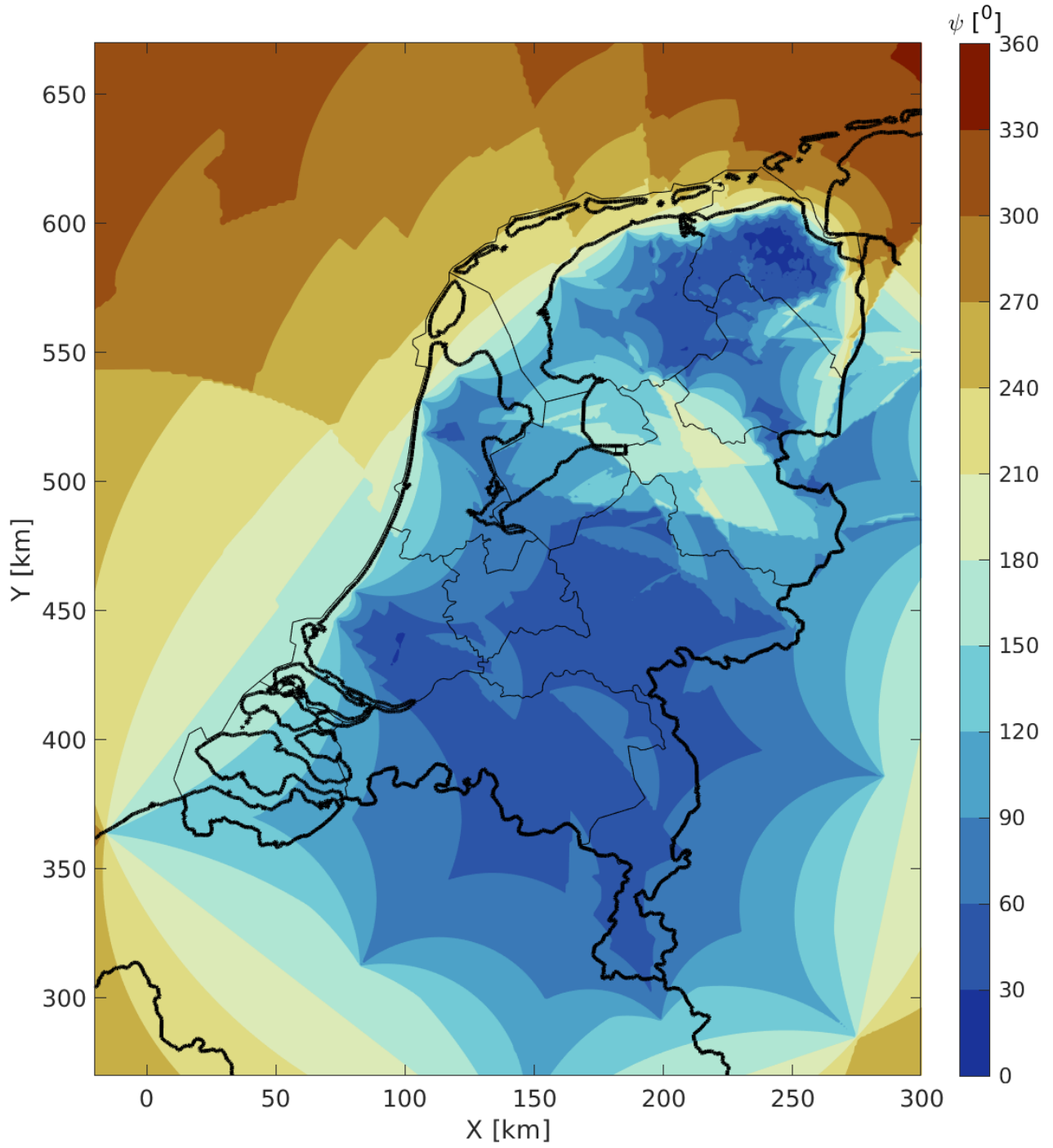


Figure D.22: Spatial distribution of azimuthal gap ψ in 2022 for $M=3.0$ (color map). Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

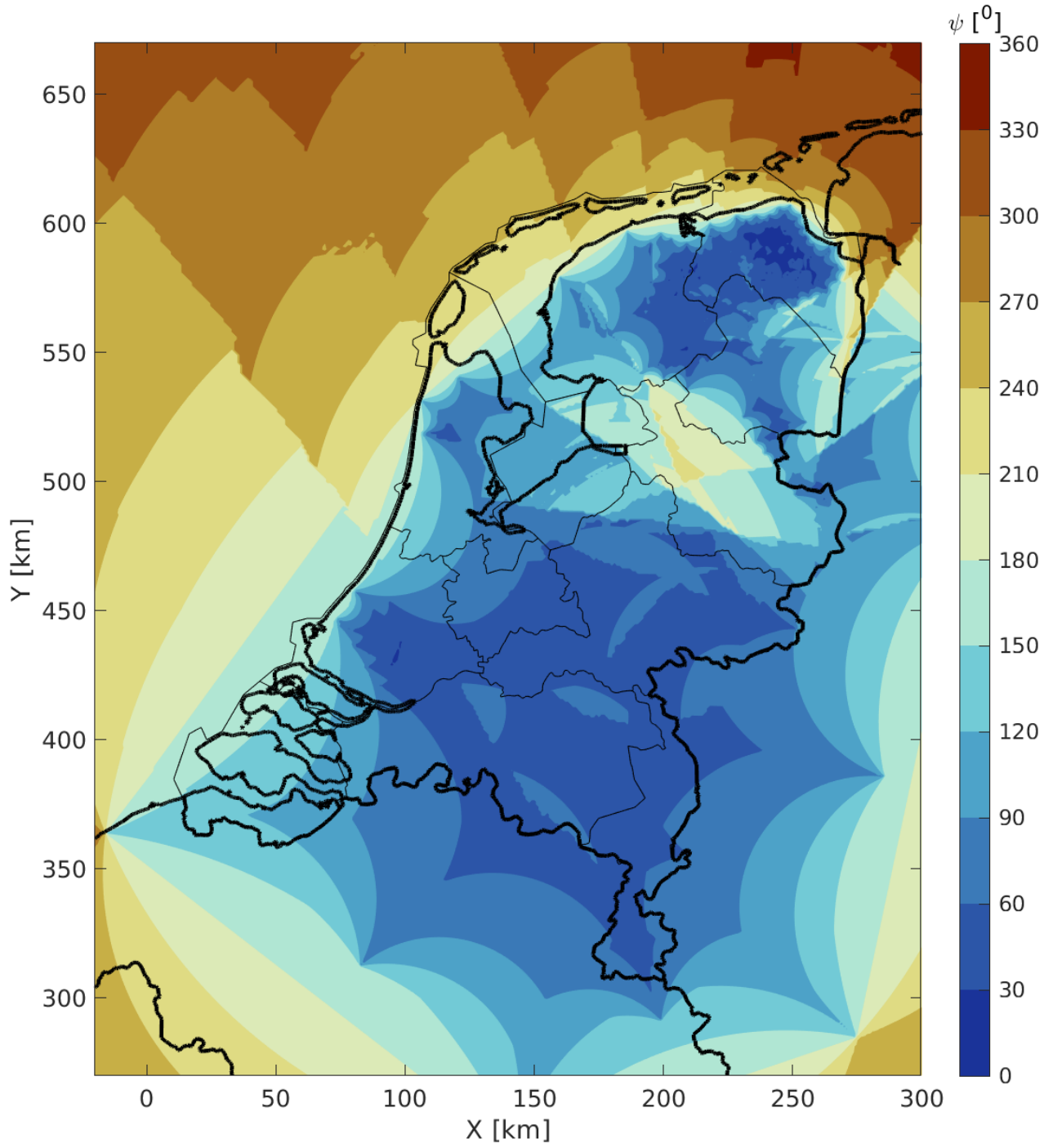


Figure D.23: Spatial distribution of azimuthal gap ψ in 2022 for $M=3.5$ (color map). Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.

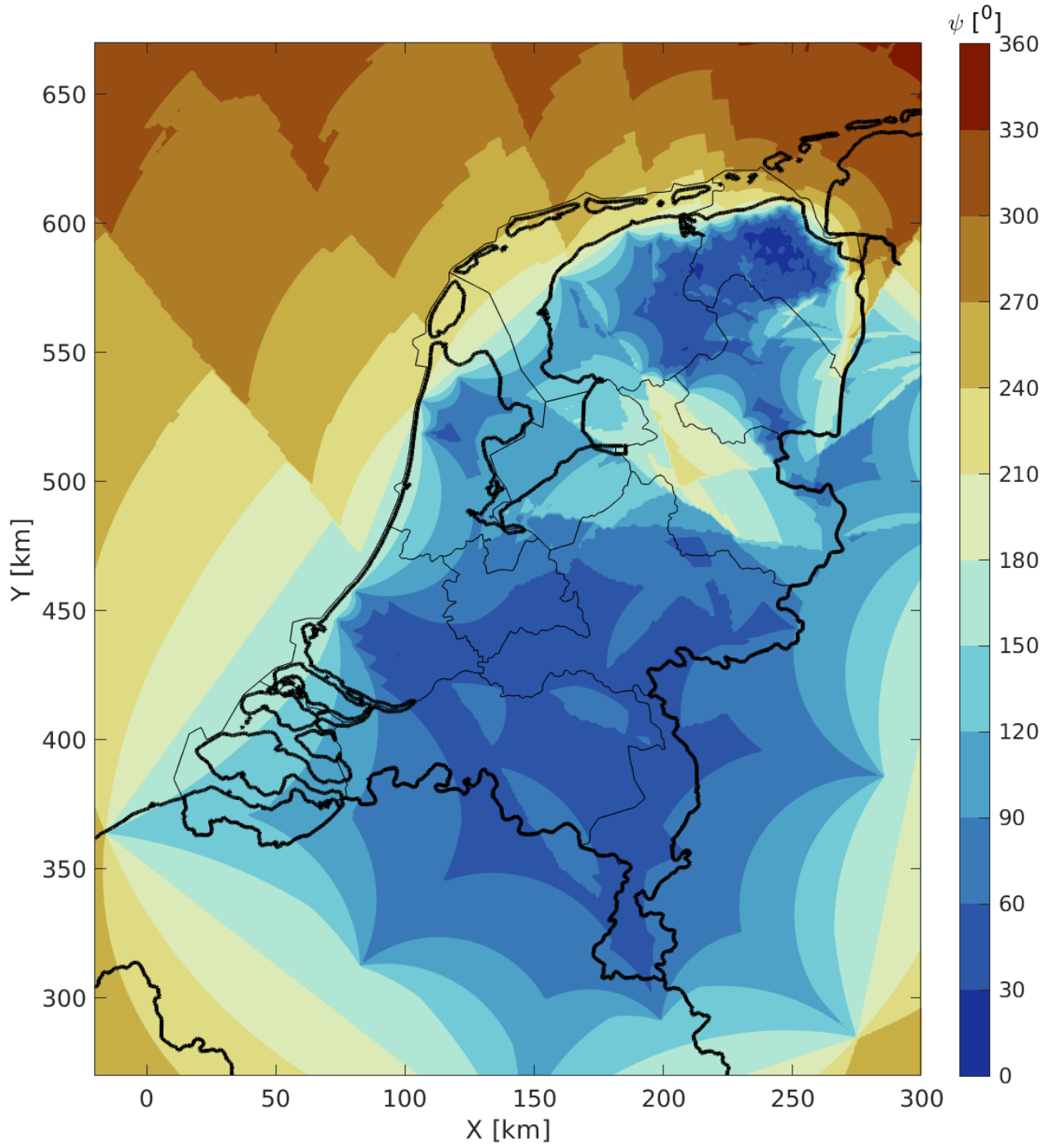


Figure D.24: Spatial distribution of azimuthal gap ψ in 2022 for $M=4.0$ (color map). Thin black lines denote province borders within the Netherlands. Thick black lines are country borders. The coordinates are in the local rectangular grid (Rijksdriehoekstelsel) in kilometers.